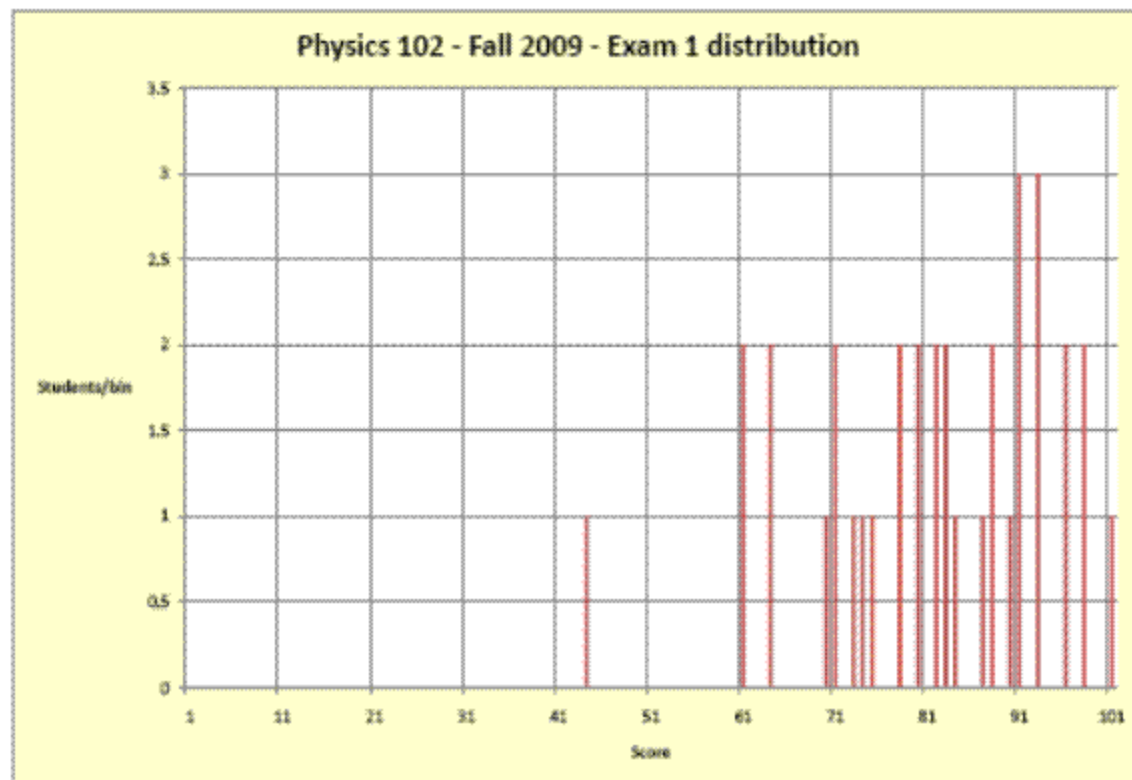


# Physics 102 - October 21, 2009

- Exam 1 graded - Available at end of class today... here
  - Will be in box outside my office door
  - After that

mean Score  $\sim 80$   
Median Score  $\sim 81$

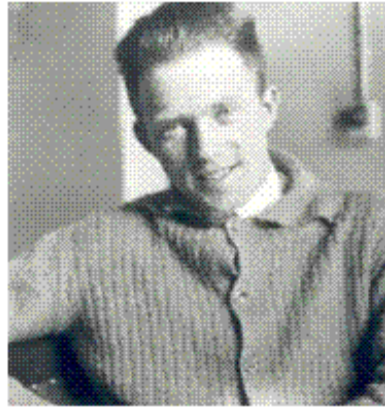


- Solns and distribution posted on web

Last Time

$$\Delta x \Delta p \geq \frac{h}{2\pi}$$

$$\Delta E \Delta t \geq \frac{h}{2\pi}$$



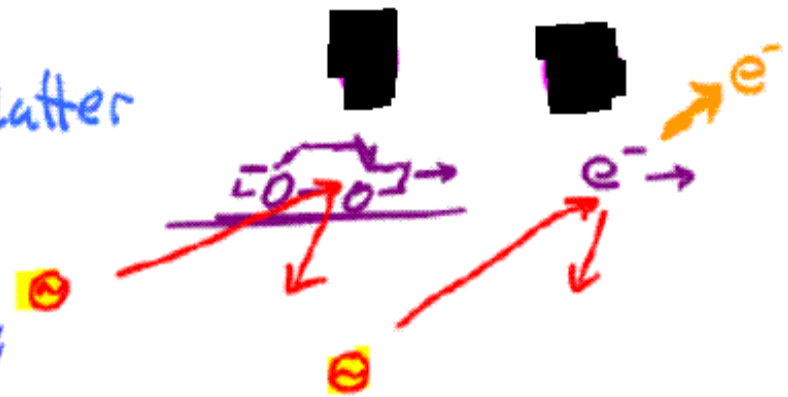
Heisenberg's  
Uncertainty  
Principle  
~1927



Say goodbye to the deterministic Universe

Size Really **DOES** matter

Say goodbye to the  
deterministic  
Universe!



"apparent" loophole in energy conservation

holds the key to understand forces and  
much of cosmology

Say hello to the Harry Potter Universe

QUANTUM Field Theory  $\rightarrow$  Exchange force



We'll come back  
to this ...

# The Journey into Inner Space

Nuclear physics



By the way, the word  
'nuclear' is pronounced

**nūk-lē-ər**

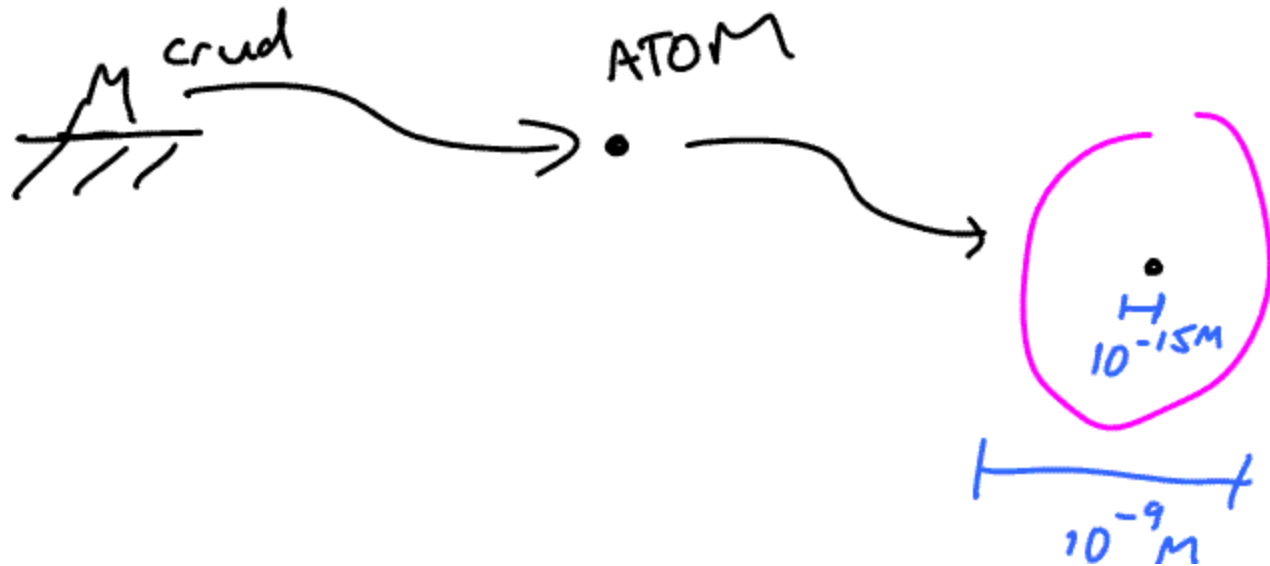
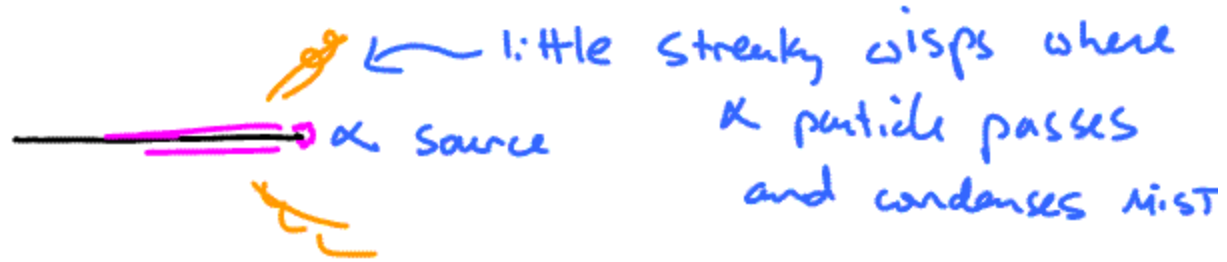
and not

**nūk-ū-lər.**

Got that Mr.  
President!?



# Cloud chamber



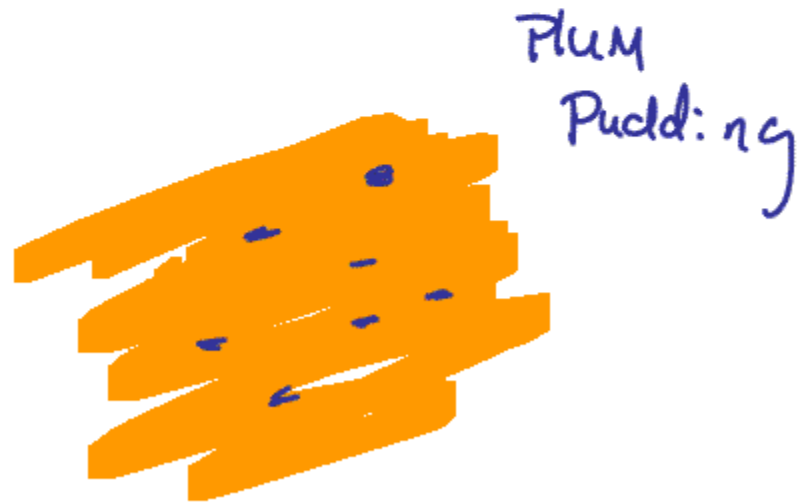
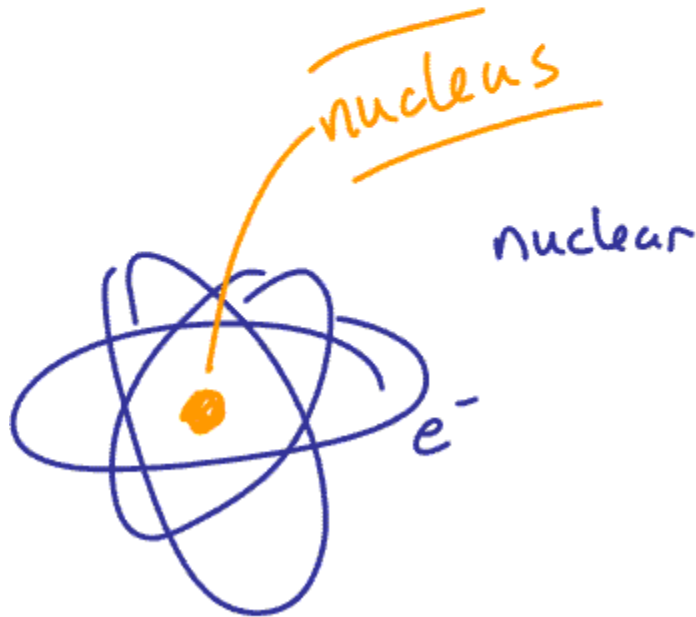
# The nuclear Model of the atom

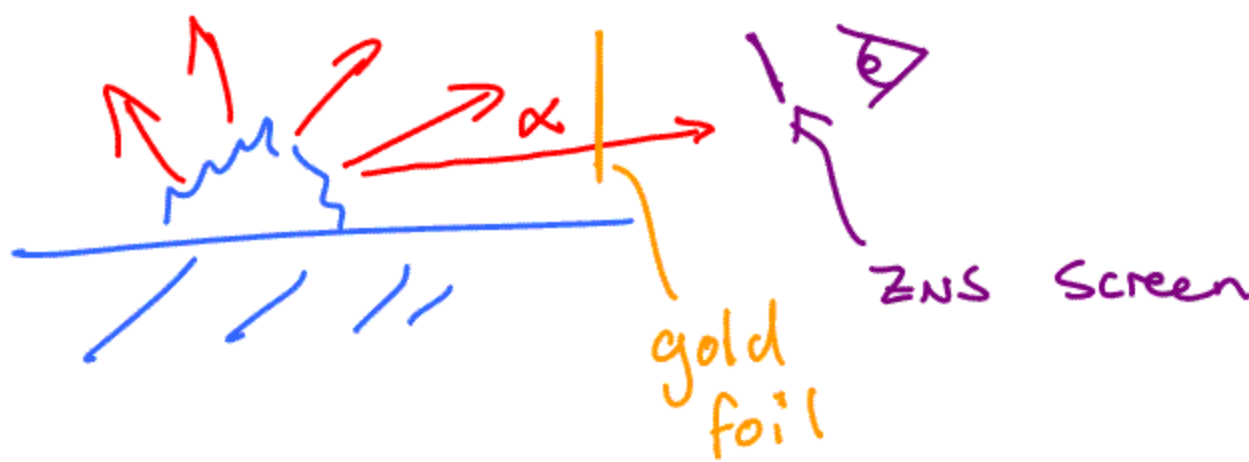


Ernest Rutherford  
(1871 - 1937)

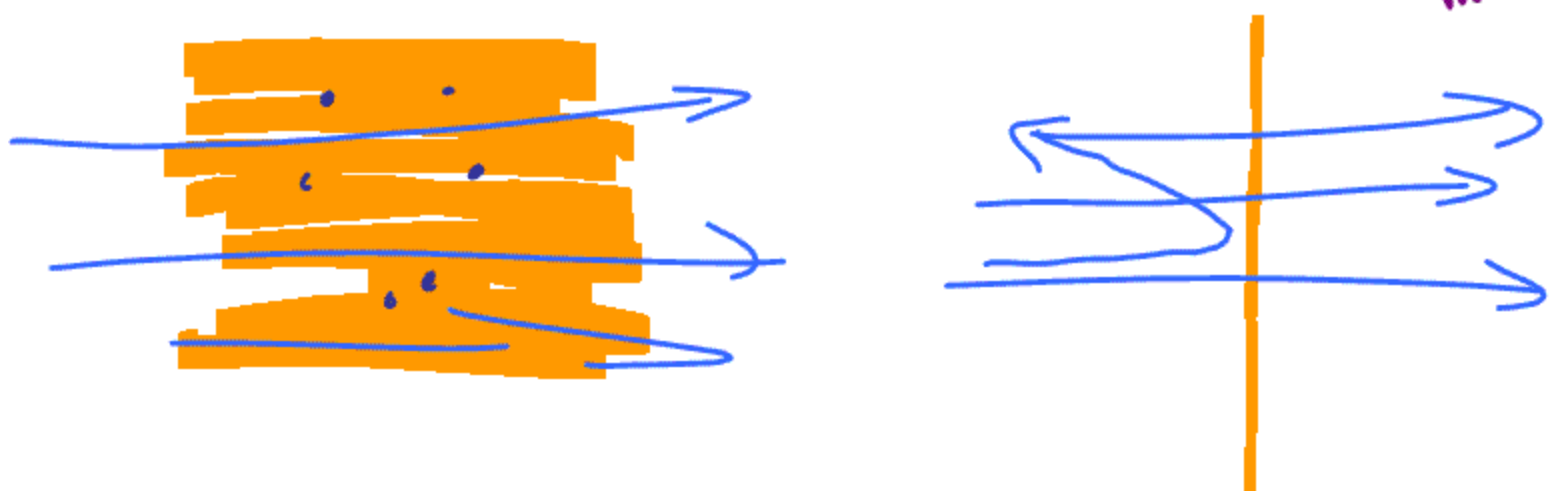
New Zealand Farm boy

↳ Manchester, England

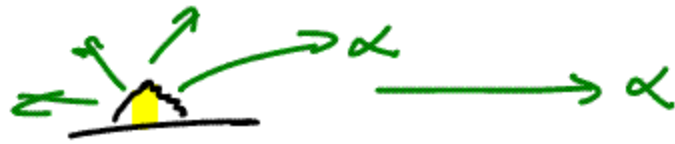




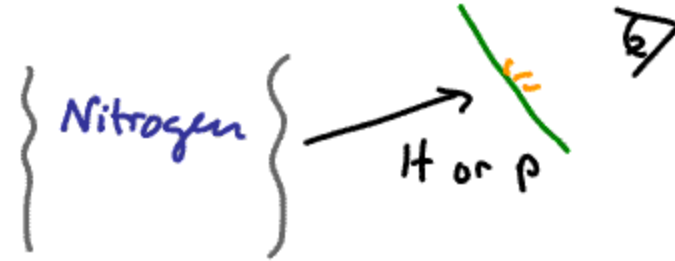
The dramatic "hard"  
scatters prove that  
positive charge compressed  
in small  
nucleus



# The Proton



Rutherford 1918



Mass  $\sim 938 \text{ MeV}/c^2$

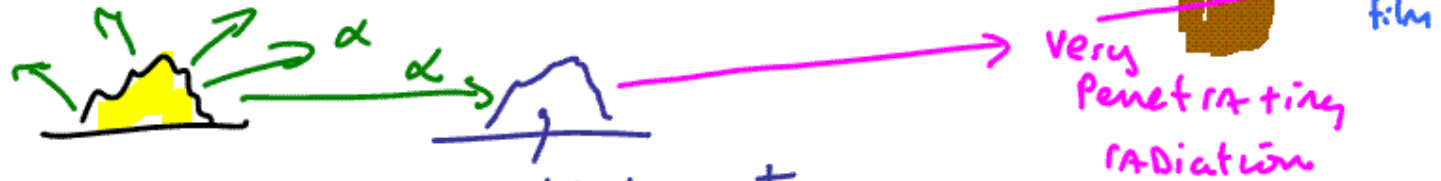
Protons  $q = +1 \text{ |e|}$

Fermion Spin  $\frac{1}{2}$

for comparison  
 $M_{e^-} \sim 0.511 \text{ MeV}/c^2$



# The neutron



1930 Walther Bothe

↙ H. Becker  
(Germany)



1954 Nobel Prize in Physics

"For the coincidence method and his discoveries made therewith"

1932 James Chadwick  
(England)



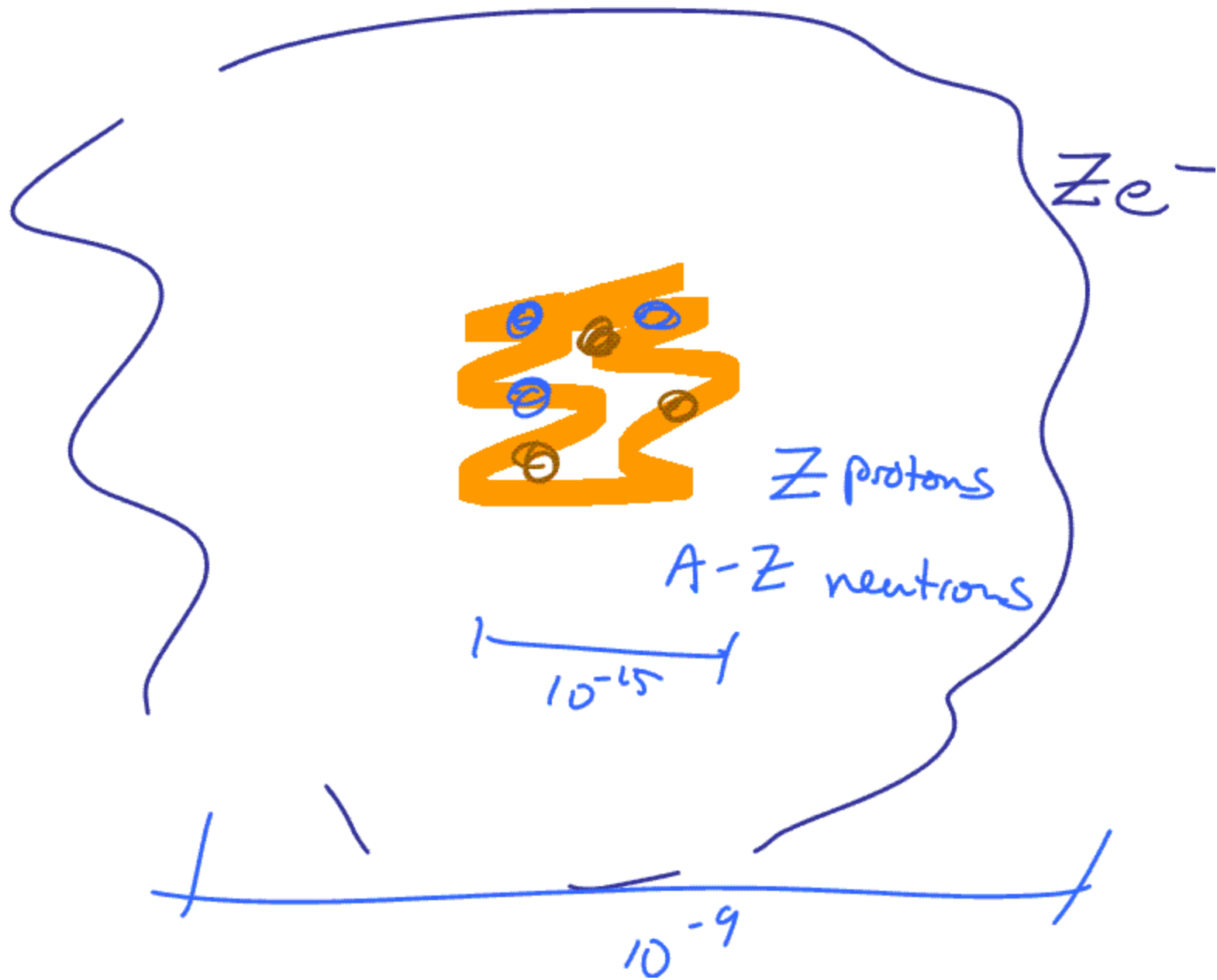
Showed this penetrating radiation to NOT be  $\gamma$ -rays

Showed had mass similar to proton and was uncharged

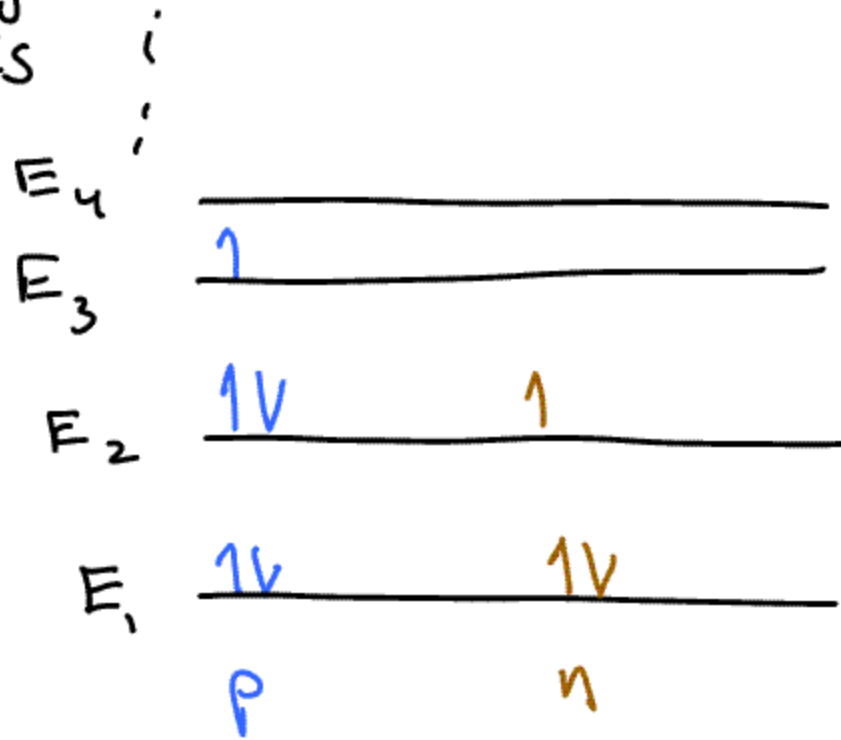
→ 1935 Nobel Prize "for discovery of the neutron"

Mass of neutron  $\sim 940 \text{ MeV}/c^2$

electric chg = 0  
Spin =  $1/2$  fermion



Discrete  
energy  
levels

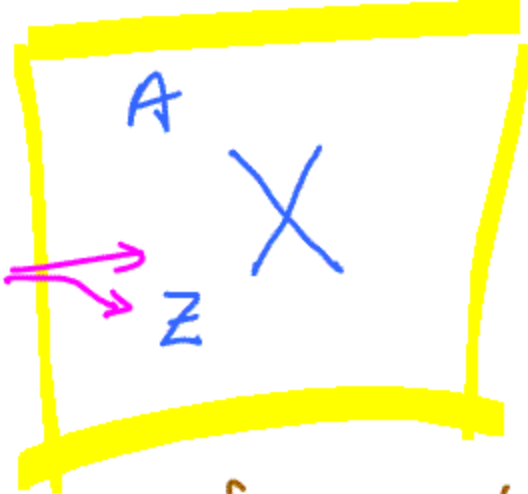


ATOMS  
10's of eV

Millions  
of  
eV

In a fashion similar to the electron energy levels in an atom — neutrons and protons occupy discrete energy levels in a nucleus. Typical energy differences between levels is Millions of electron-volts (as opposed to tens of electron-volts for atomic levels).

X, Z  
SAME  
INFO



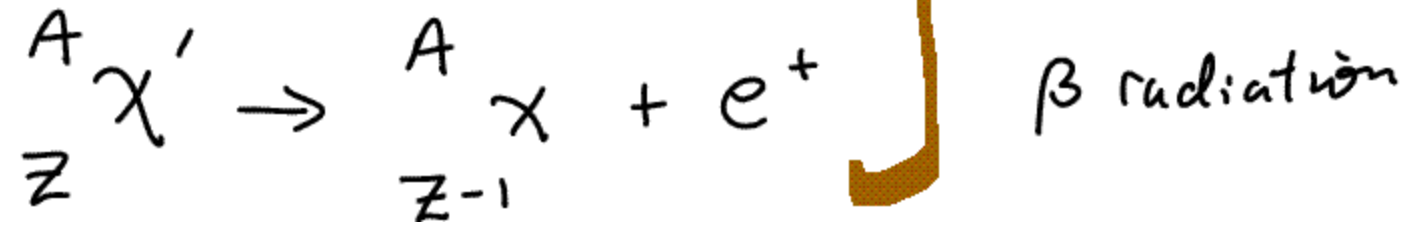
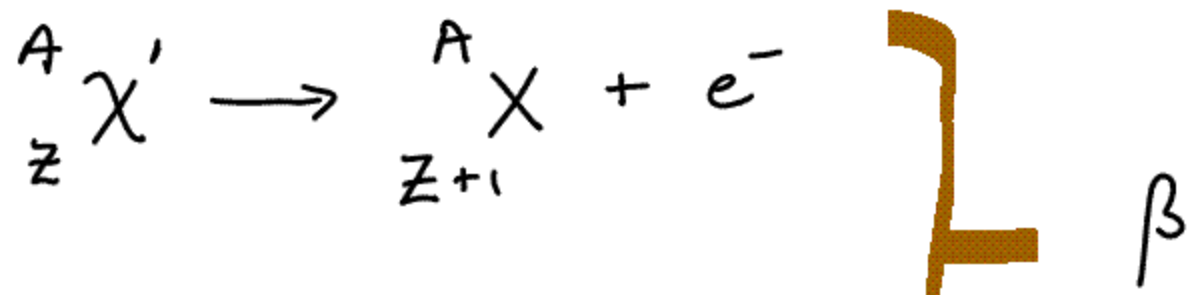
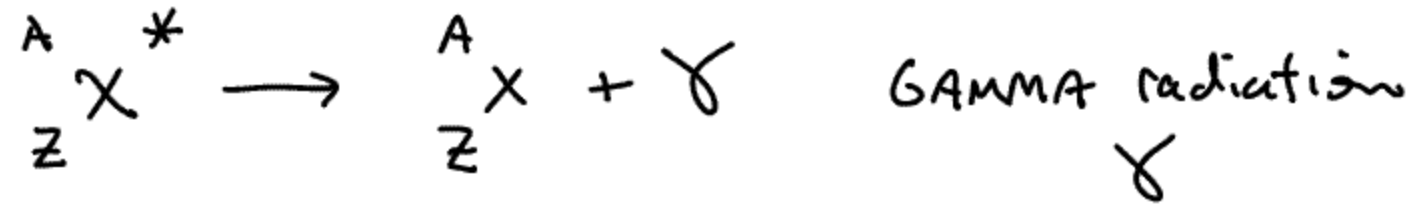
X  $\equiv$  Atomic Symbol

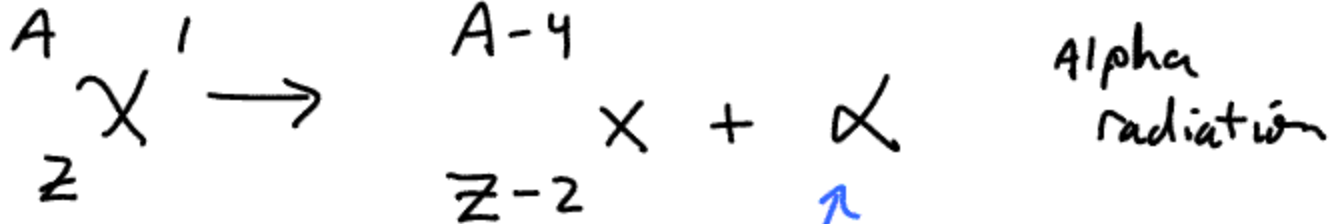
Z  $\equiv$  # protons

Atomic #

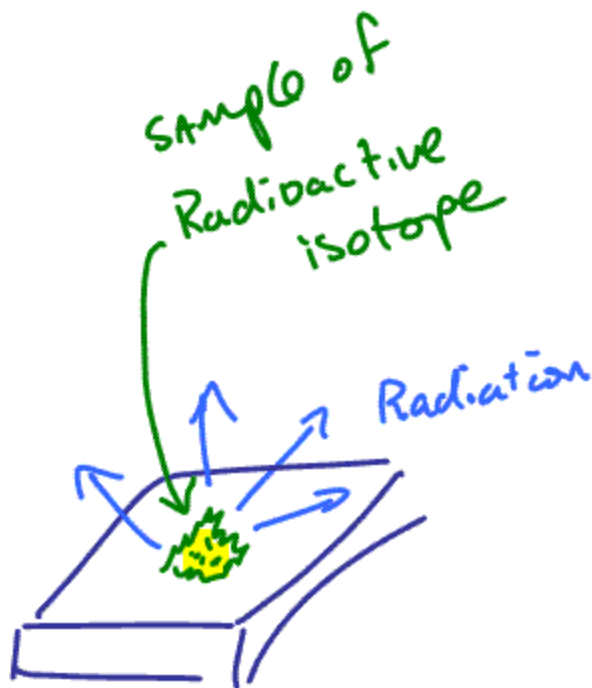
A  $\equiv$  Atomic Mass

Symbol for a nucleus N  $\equiv$  A - Z  $\equiv$  # neutrons





decay constant



$$\frac{\Delta N}{\Delta t} = \lambda N$$

$N \equiv$  # radioactive nuclei in sample

Activity  $\equiv$  # nuclei decaying per second

$$t_{1/2} = \frac{0.693}{\lambda} \equiv \text{half-life} \equiv \text{time for } \frac{1}{2} \text{ the nuclei to decay}$$

$t_{1/2}$   
 $\lambda$  }  $\equiv$  both characterize how prone a specific type of nucleus is to break down (decay)

large  $\lambda$ , small  $t_{1/2}$  means nuclei prone to decay quickly

Small  $\lambda$ , large  $t_{1/2}$  nuclei tend to stick around longer

$\lambda$  and  $t_{1/2}$  ... measures of the same characteristic

Sort of like a fingerprint for specific type of radioactive nucleus



$N$  atoms in sample  
at time = 0

$$\text{Activity} \equiv \frac{\# \text{decays}}{\text{second}} = \frac{\Delta N}{\Delta t} = \lambda N$$

↑  
decay CONSTANT

half life  $\equiv t_{1/2}$  = time for  $1/2$  sample to decay

$$t_{1/2} = \frac{0.693}{\lambda}$$

## Radioactive Dating

Normal  $^{12}_6\text{C}$  carbon



$^{14}_6\text{C}$  Produced by cosmic rays hitting atmosphere

$^{14}_6\text{C}$  is naturally radioactive  
 $\beta$ -emitter

$t_{1/2} = 5730$  years

- $^{14}_6\text{C}$  incorporated into living tissue
- Stops at death
- $^{14}_6\text{C}/^{12}_6\text{C}$  ratio gives estimate of time since death
- $^{14}_6\text{C}$  concentration in atmosphere varies, calibrate w/ tree rings