

# Physics 100 - April 13, 2009

Group	topic	15-Apr-09	20-Apr-09	27-Apr-09	29-Apr-09
A	Nuclear physics and society	3	1	✓ 2	4
B	String theory	3	2	✓ 1	4
C	Global Positioning System				✓ 1
D	Search for extra-terrestrial life	✓ 1	2	3	4
E	Music	3	4	1	✓ 2
F	Life and times of a great physicist	3	✓ 1	2	4
G	Radiation - dangers and uses	4	✓ 1	2	3

- Presentations begin this Wednesday
- EXAM 2 one week from Wednesday  
plan to schedule a Q + A session

# This is your life ... Starting at the very beginning

In the beginning there was ...

Space-time foam

NOT empty space

quantum fluctuations

Very small  $\sim 10^{-33}$  cm possibly

what was there and  
what exactly  
caused  
inflation  
not established  
yet.  
models  
exist

Fluctuation occurs w/ property that include  
tremendous repulsive pressure



AT  $\sim 10^{-43}$  s  $\rightarrow 10^{-35}$  s exponential expansion

■ How big did inflation make universe by  $10^{-35}$  s ?

Much bigger than observable universe

at the time

$$(10^{-35} \text{ s})(3 \times 10^8 \text{ m/s})$$

or

Bigger than now

Maybe factor of  $10^{10^{12}}$

■ What was curved  $\rightarrow$  now flat

geometry of Universe flat

■ What was small  $\rightarrow$  now big  
quantum fluctuations become large  
density/energy fluctuations

ACTS as nucleus for large-scale structure formation

■ As "inflaton field" properties change during inflation, inflation ends

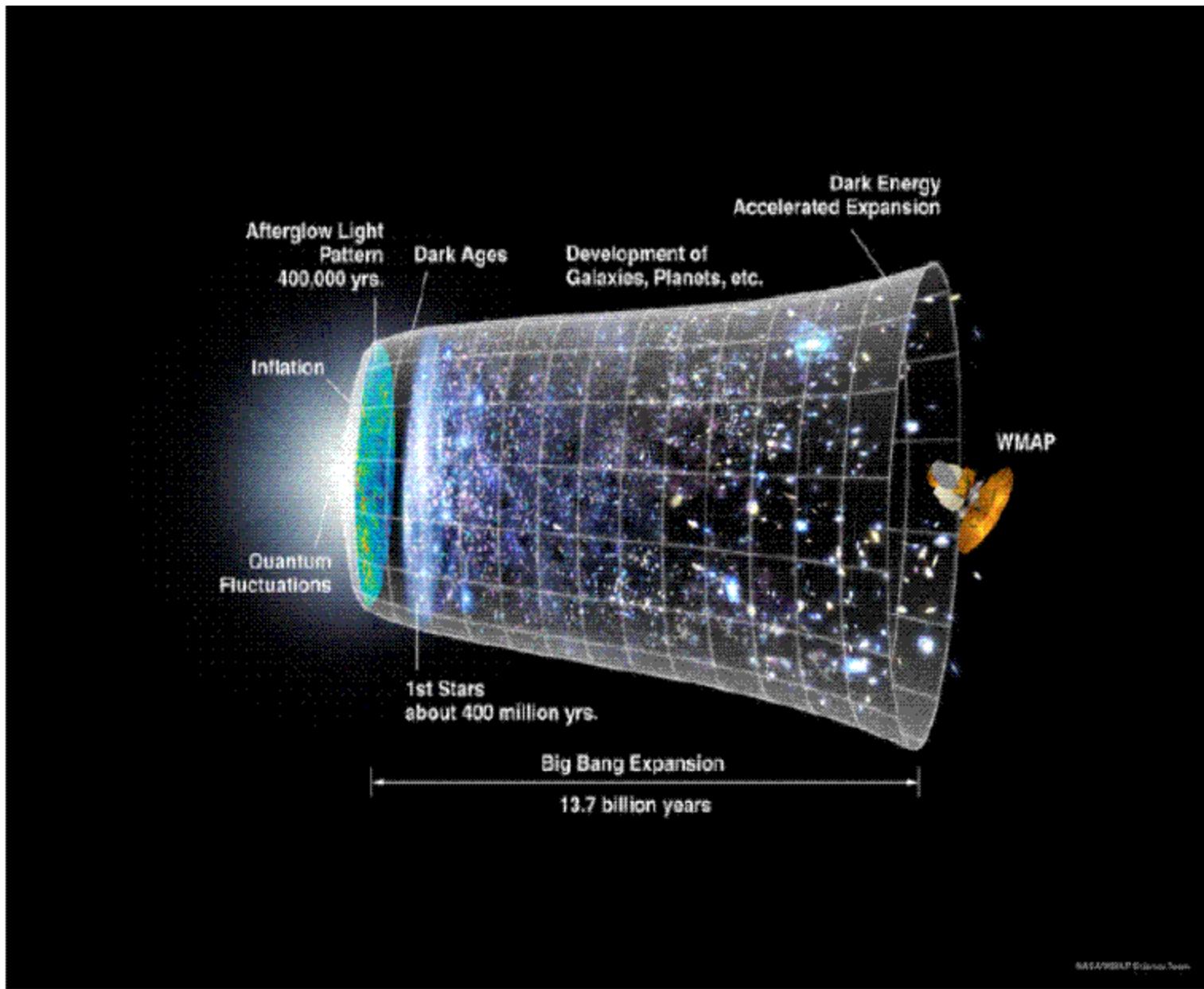
■ Energy Driving inflation Dumped into Matter and Radiation in early universe

■  $10^{-35} - 10^{-5}$  sec

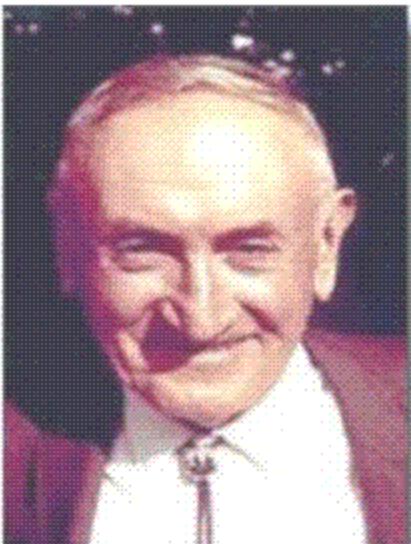
universe is primordial soup of quark-gluon plasma  
Subatomic particles quarks, Z, W,  $\gamma$ , gluons  
...

■  $10^{-5}$  sec quarks bound into  
Baryons  $\leadsto$  Protons, neutrons  
Mesons

- $t = 3$  minutes      light nuclei form  
Big Bang Nucleosynthesis
- $t = 400,000$  years      neutral atoms form  
Universe becomes Transparent  
light we see from this time called  
Cosmic Microwave Background
- $t = 400$  million years      First Stars
- $t = 13.7$  billion years      Buffalo wins  
Superbowl  
Now
- Expansion is accelerating



# DARK MATTER



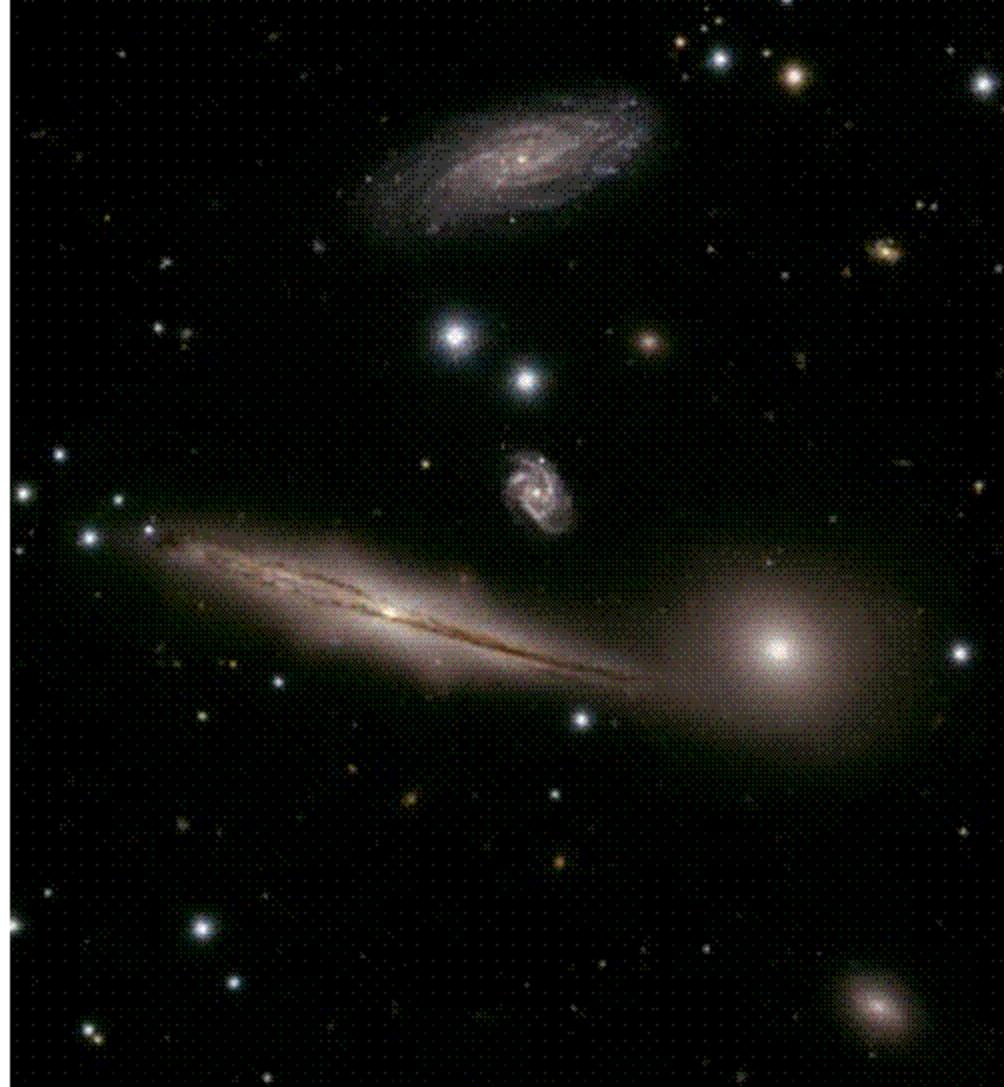
Fritz Zwicky (1898-1974)  
Cal Tech astrophysicist

- jet engines
- "Spherical bastards"
- Suggested galaxies could act as gravitational lenses
- Dark Matter

Zwicky compared mass of galactic cluster using two methods

① number + brightness of galaxies in cluster

② motion of galaxies at edge of cluster



Mass | >> method 2

Mass | method 1

galactic cluster

DARK Matter

# Dark Matter

Can relate velocity  
radius and force  
in orbits.

Have seen that  
orbits in stars  
and galactic clusters  
require stronger  
gravitational force  
than can be explained  
by conventional  
observable "visible"



ORBITS

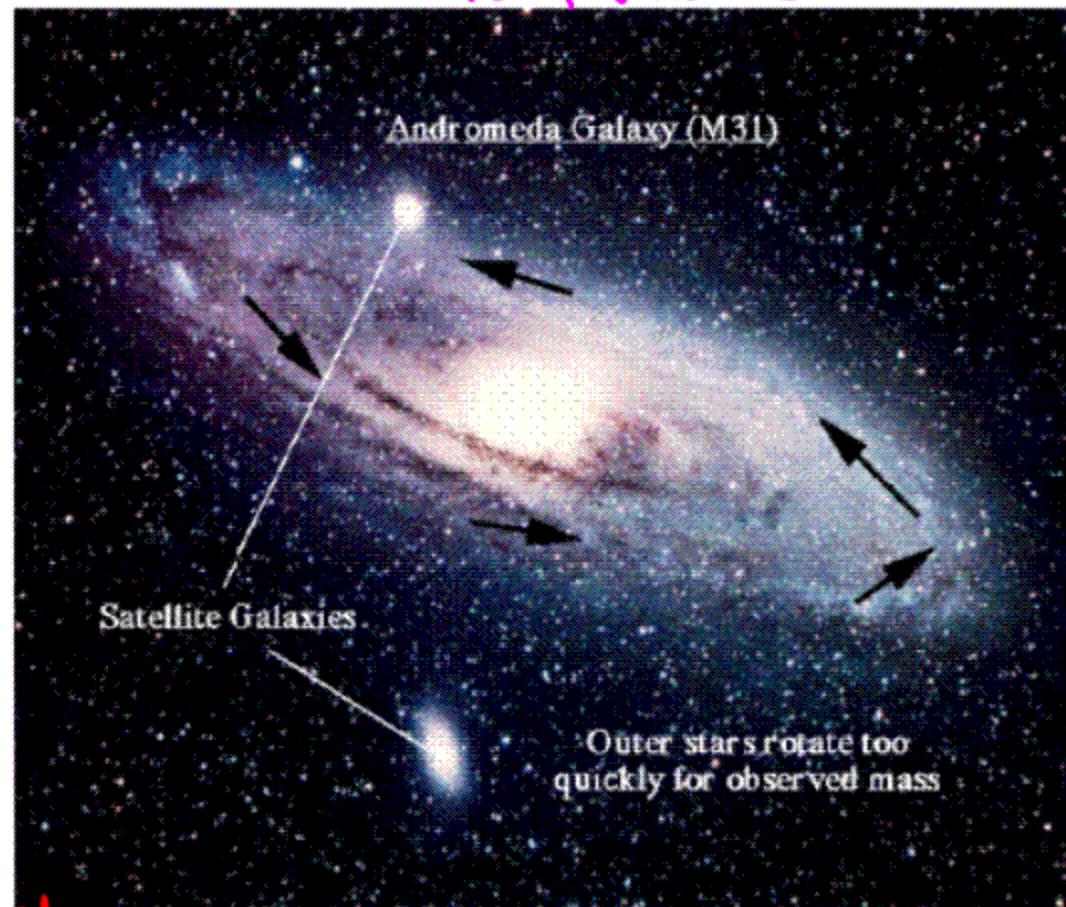
$$F = \frac{mv^2}{R}$$

Circular Motion

$$F = \frac{GMm}{R^2}$$

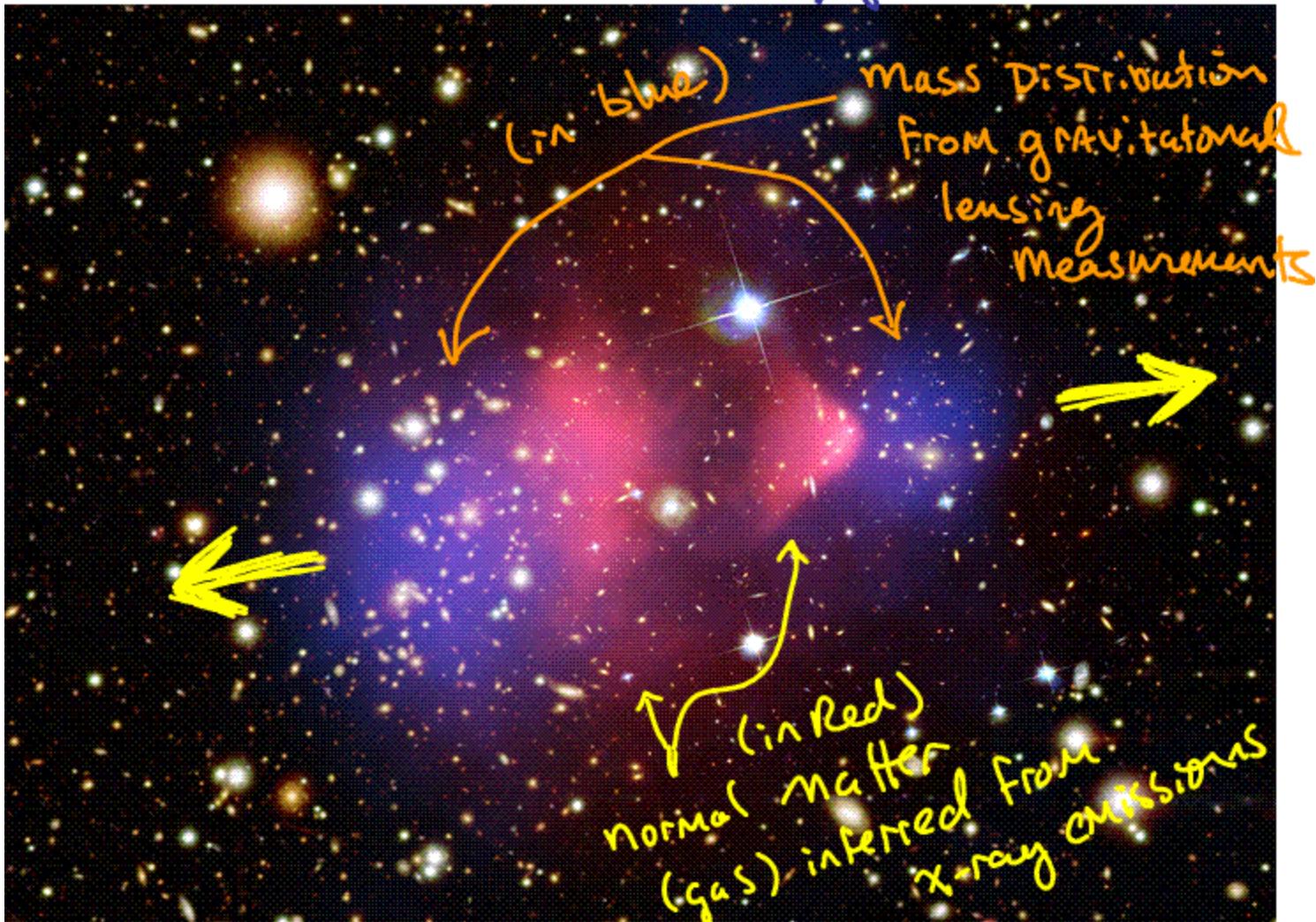
$$\frac{mv^2}{R} = \frac{GMm}{R^2}$$

1975 - Kent Ford, Vera Rubin



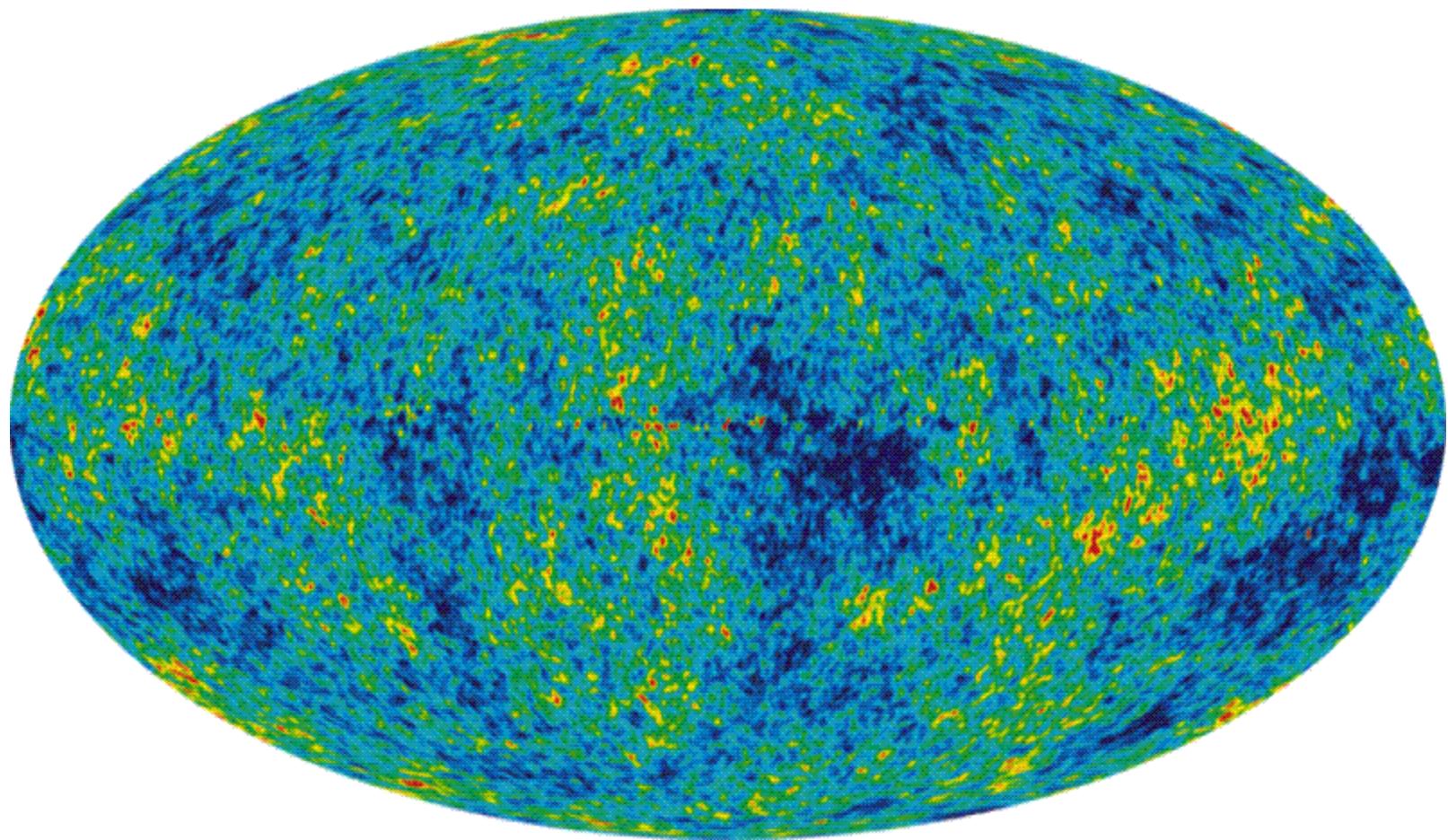
-P. Cushman

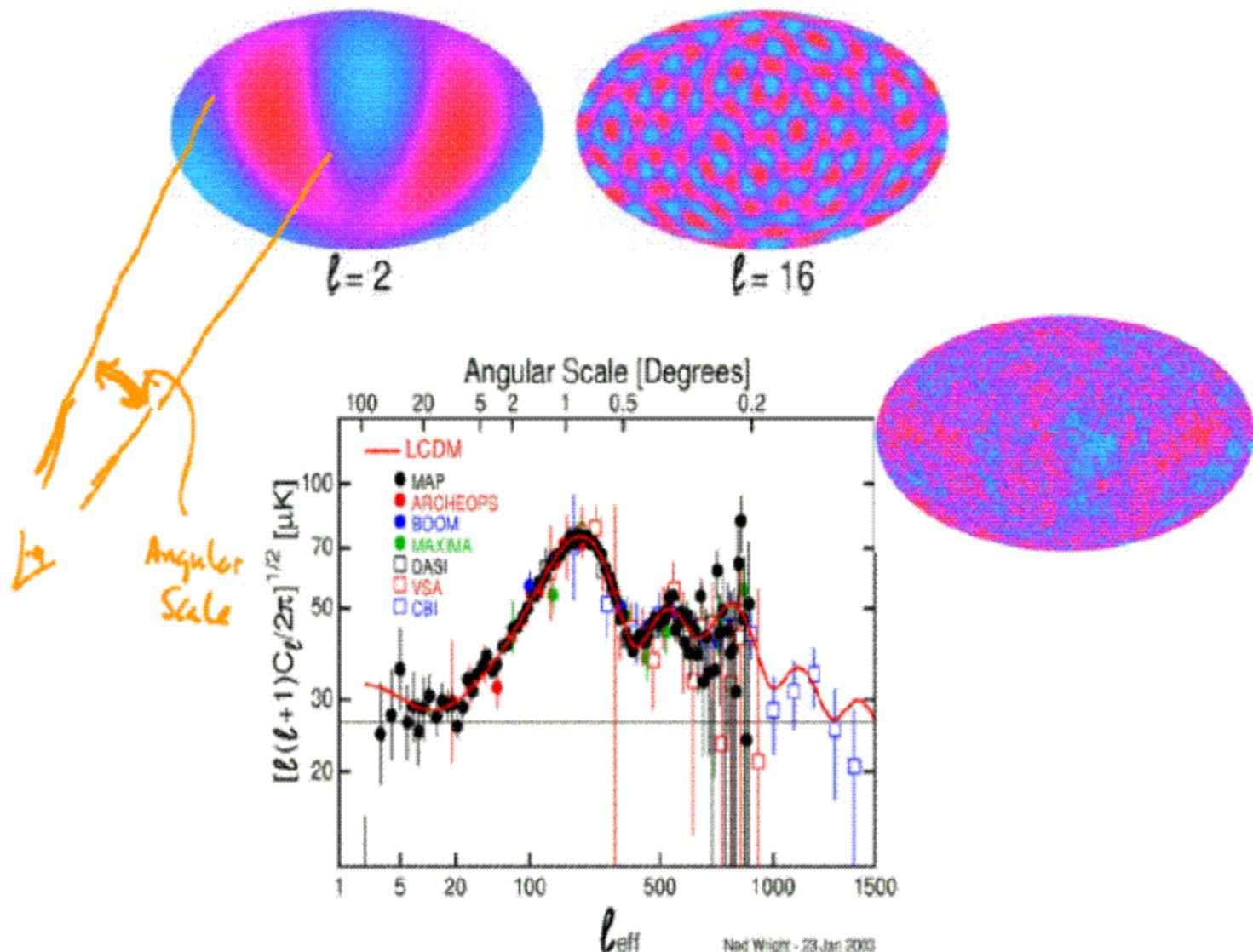
## Bullet Cluster - colliding galactic clusters



The universe at  $t = 400,000$  years

CMB  
from WMAP

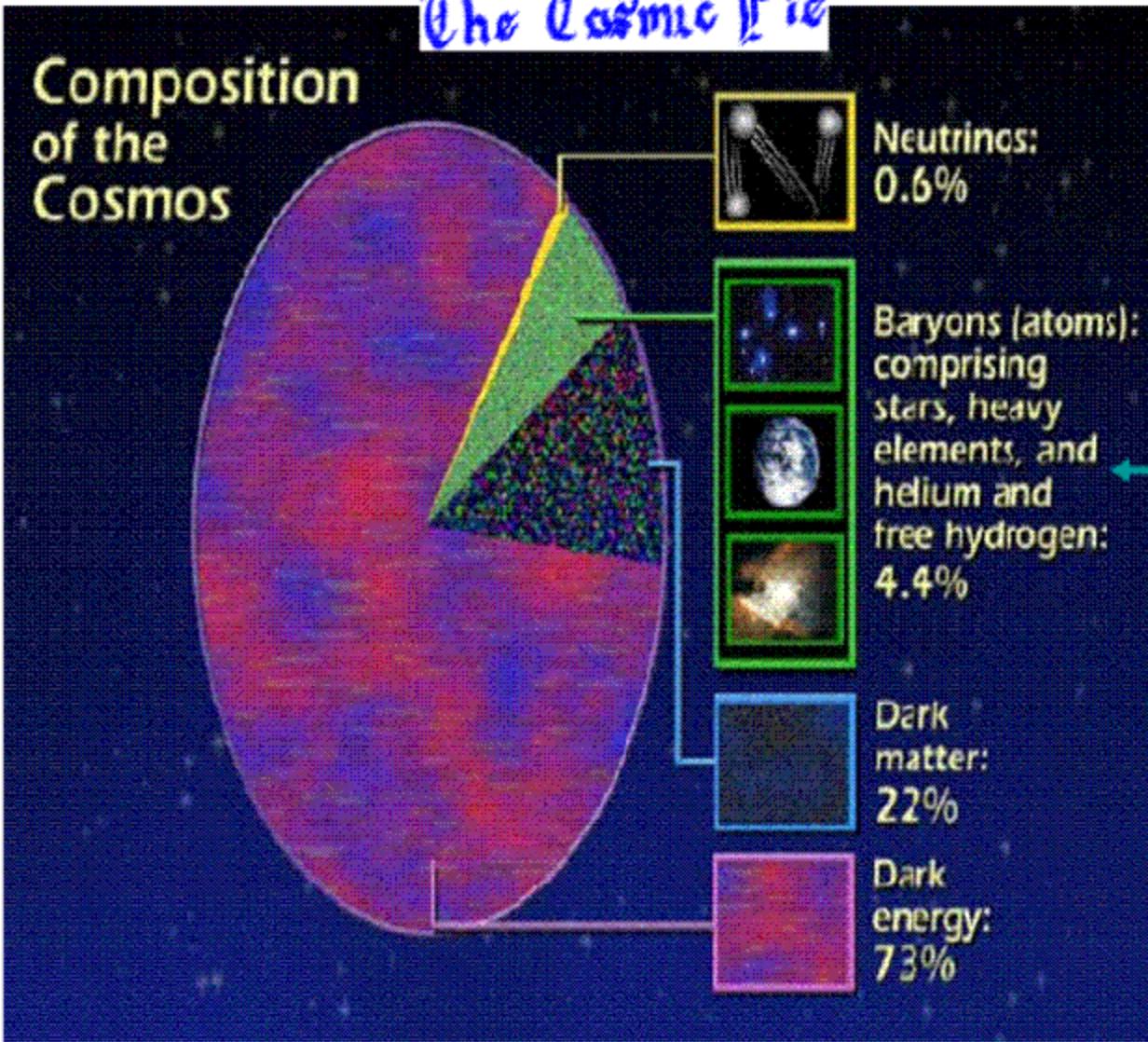




"Power spectrum" (size) of temperature fluctuations  
sensitive to different matter/energy components of the  
universe

# The Cosmic Pie

## Composition of the Cosmos



Us

STScI

95% of the universe is unknown!

Figure from E. Linder  
LBL

Inflationary  
Big Bang Model

STANDARD  
Model  
of Particle  
Physics

Dark  
Matter

- interacts gravitationally  
but NOT via electromagnetism  
or strong force

Accelerating  
Expansion of  
Universe  
↳ "Dark Energy"

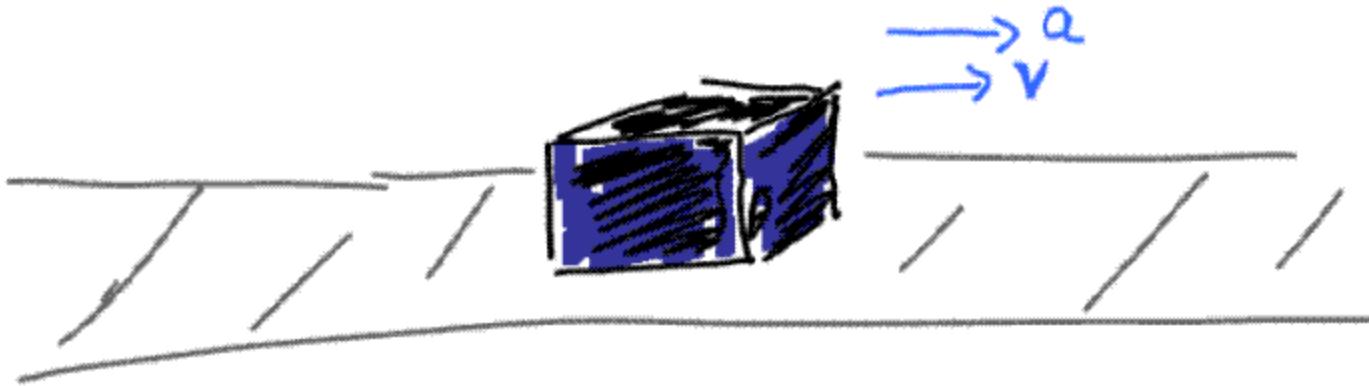
Andromeda Galaxy (M31)

Satellite Galaxies

Other stars recede too  
quickly for observed blues

dark energy → sort of an  
inflation-like pressure (slower)  
on space expansion - Exact  
Source unknown as yet.

# Dark Energy Analogy



See box accelerating down sidewalk

We Know there is a force + source of energy.

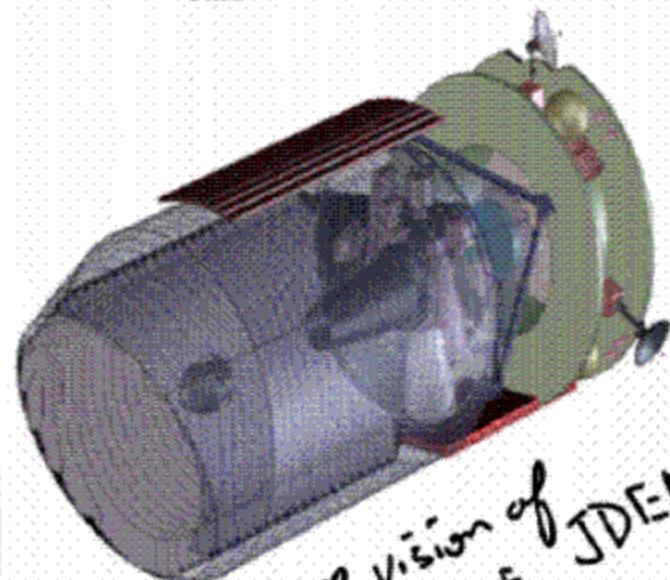
$$(F)(\text{dist}) = \text{work} = \text{Energy}$$

We have ideas ... but we don't know for sure the nature of the force / energy

1/20

Where do we go from here?

SNAP  
SuperNova  
Acceleration  
Probe



one vision of  
NASA DOE JDEM  
Joint Dark Energy Mission



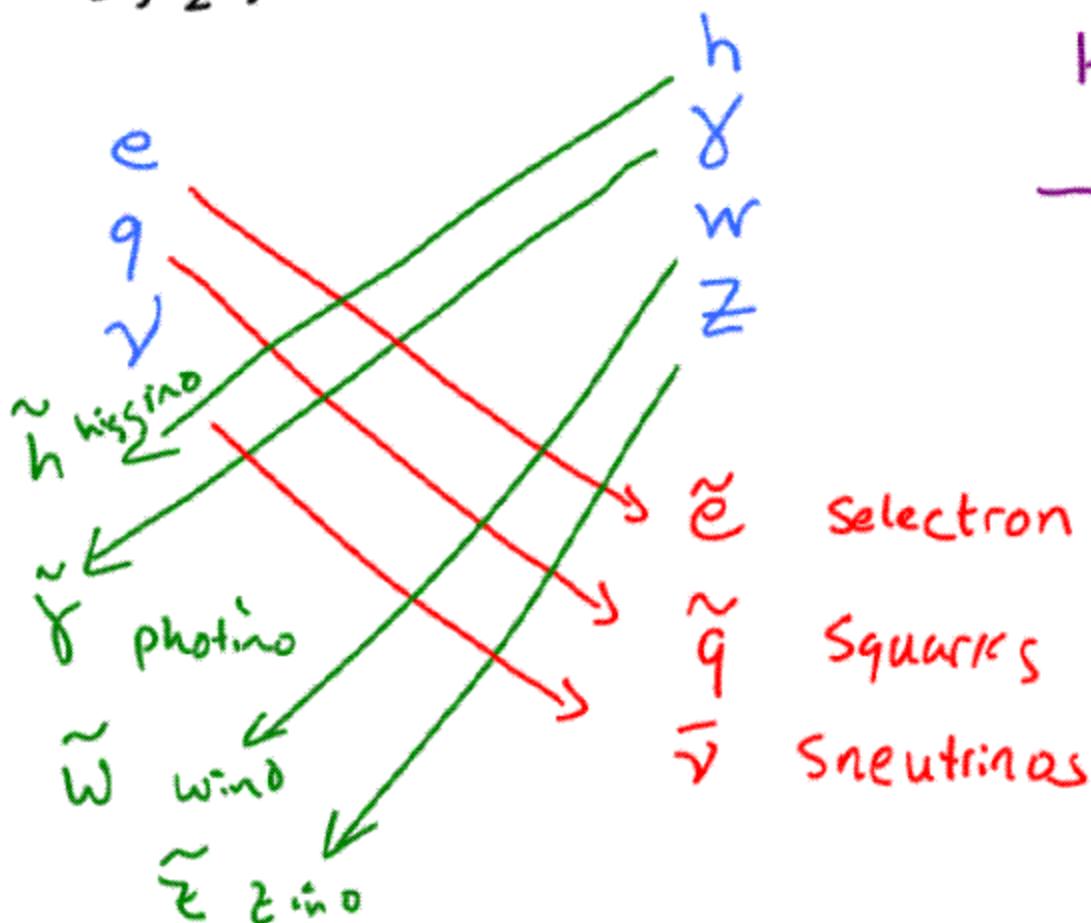
Large Hadron  
Collider (LHC)

# Supersymmetry

fermion  $\longleftrightarrow$  boson

$\frac{1}{2}, \frac{3}{2}, \dots$

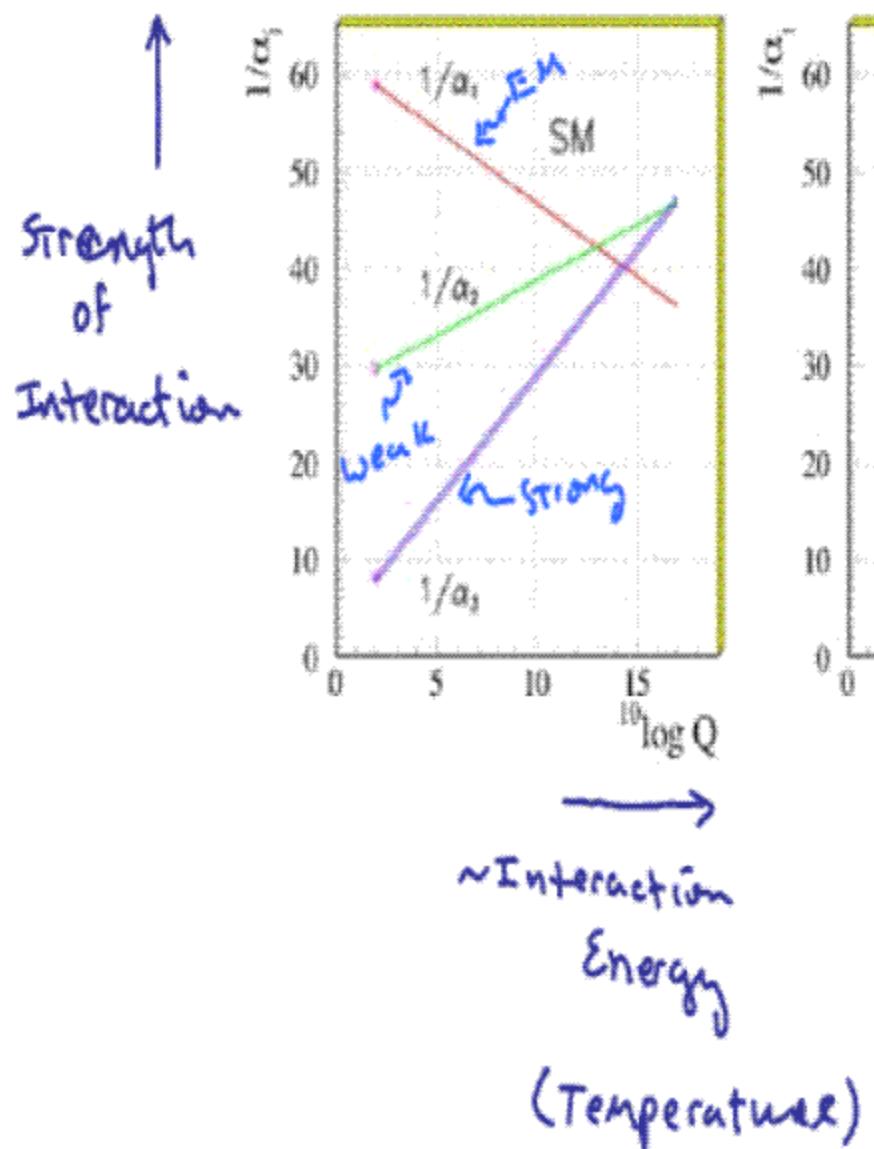
0, 1, ...



Intro to particle Spec  
R parity  
LSP  
Gauge hierarchy  
Coupling Unification  
Dark energy?



## STANDARD Model



## Minimal Supersymmetric Model

