

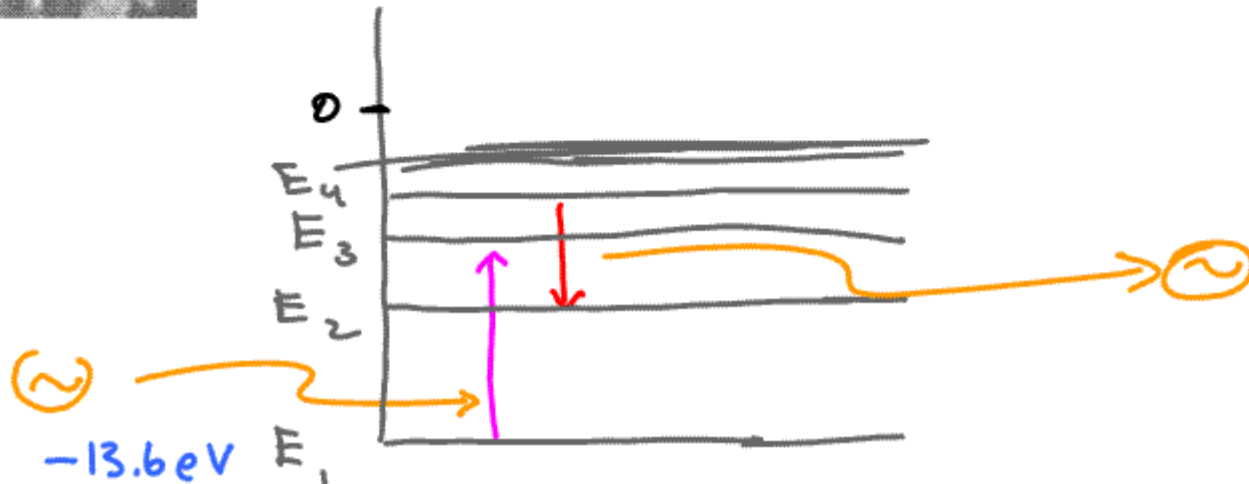
P114 - April 27, 2006

Last Time -

Bohr Atom - Semi-classical

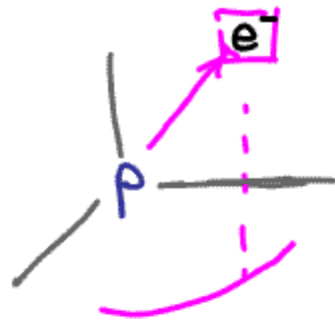


Coulomb
+
Circular Motion
+
quantization of
Angular Momentum



QM treatment of H atom

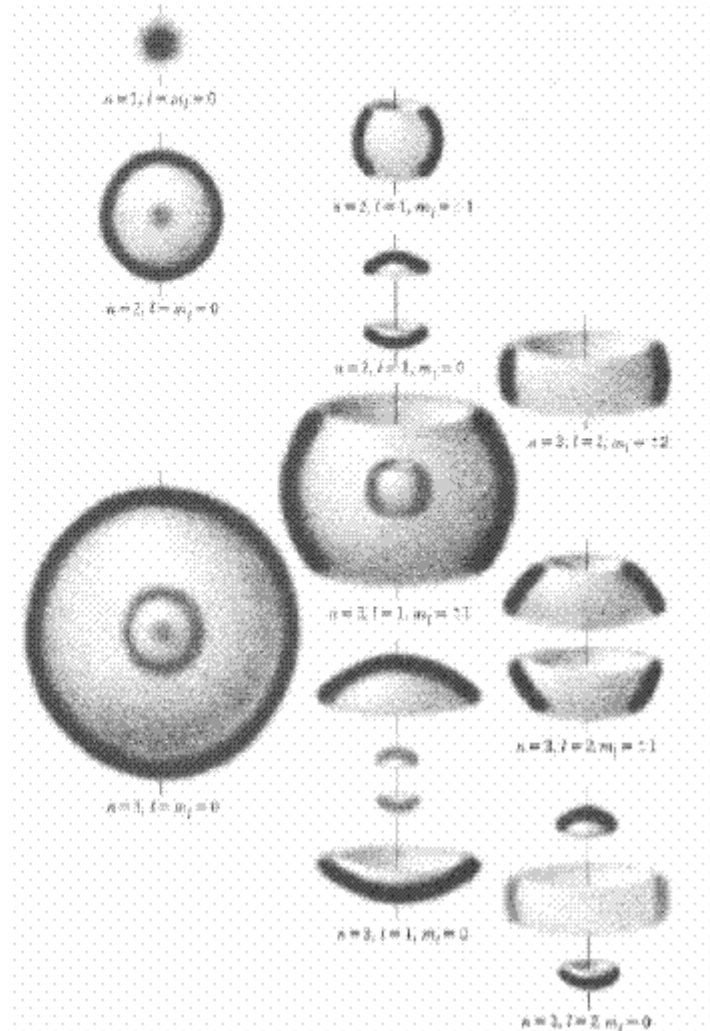
wave mechanics
Spherical Symmetry



Energy or principal quantum number
 $n = 1, 2, 3 \dots$

Orbital quantum number
 $l = 0, 1, \dots, n-1$

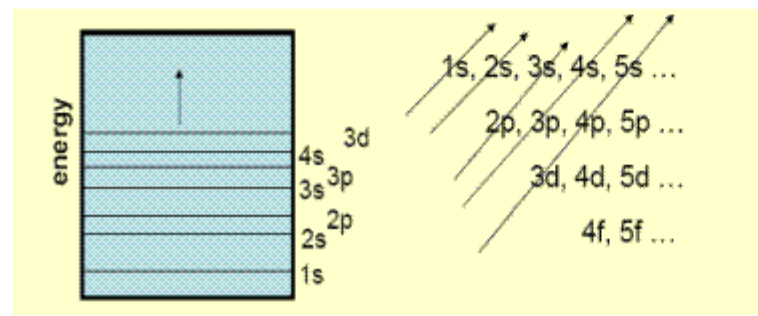
Magnetic quantum number
 $-l, -|l-1|, \dots, 0, 1, \dots, |l-1|, l$



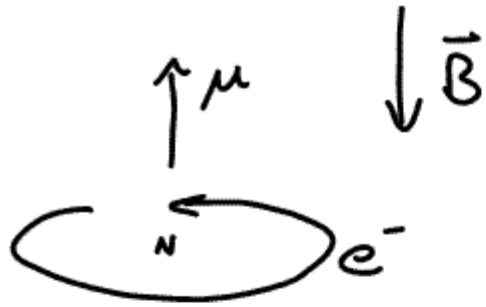
multi-electron Atoms

Energy level (n)	Sublevel (l)	K shell			L shell			M shell			...
		1	2	3	4	5	6	7	8	9	
z = #e ⁻		s	s	p	p	p	s	p	p	p	
				m _l =+1	m _l =0	m _l =-1					
1	H	1	—	—	—	—	—	—	—	—	1s ¹
2	He	1↓	—	—	—	—	—	—	—	—	1s ²
3	Li	1↓	1	—	—	—	—	—	—	—	1s ² 2s ¹
4	Be	1↓	1↓	—	—	—	—	—	—	—	1s ² 2s ²
5	B	1↓	1↓	1	—	—	—	—	—	—	1s ² 2s ² 2p ¹
6	C	1↓	1↓	1	1	—	—	—	—	—	
7	N	1↓	1↓	1	1	1	—	—	—	—	
8	O	1↓	1↓	1↓	1	1	—	—	—	—	1s ² 2s ² 2p ⁴

Aufbau ↑



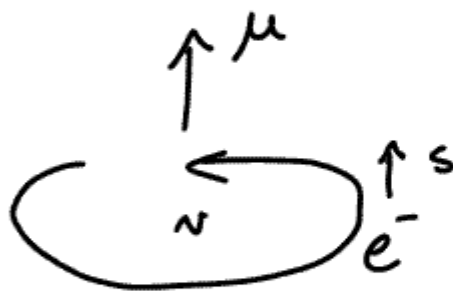
Hund's rule
 fermion symmetrization
 plus coulomb



Zee-man Effect

magnetic moment of e^-
in external field

\Rightarrow Split Spectral lines



Spin-orbit coupling

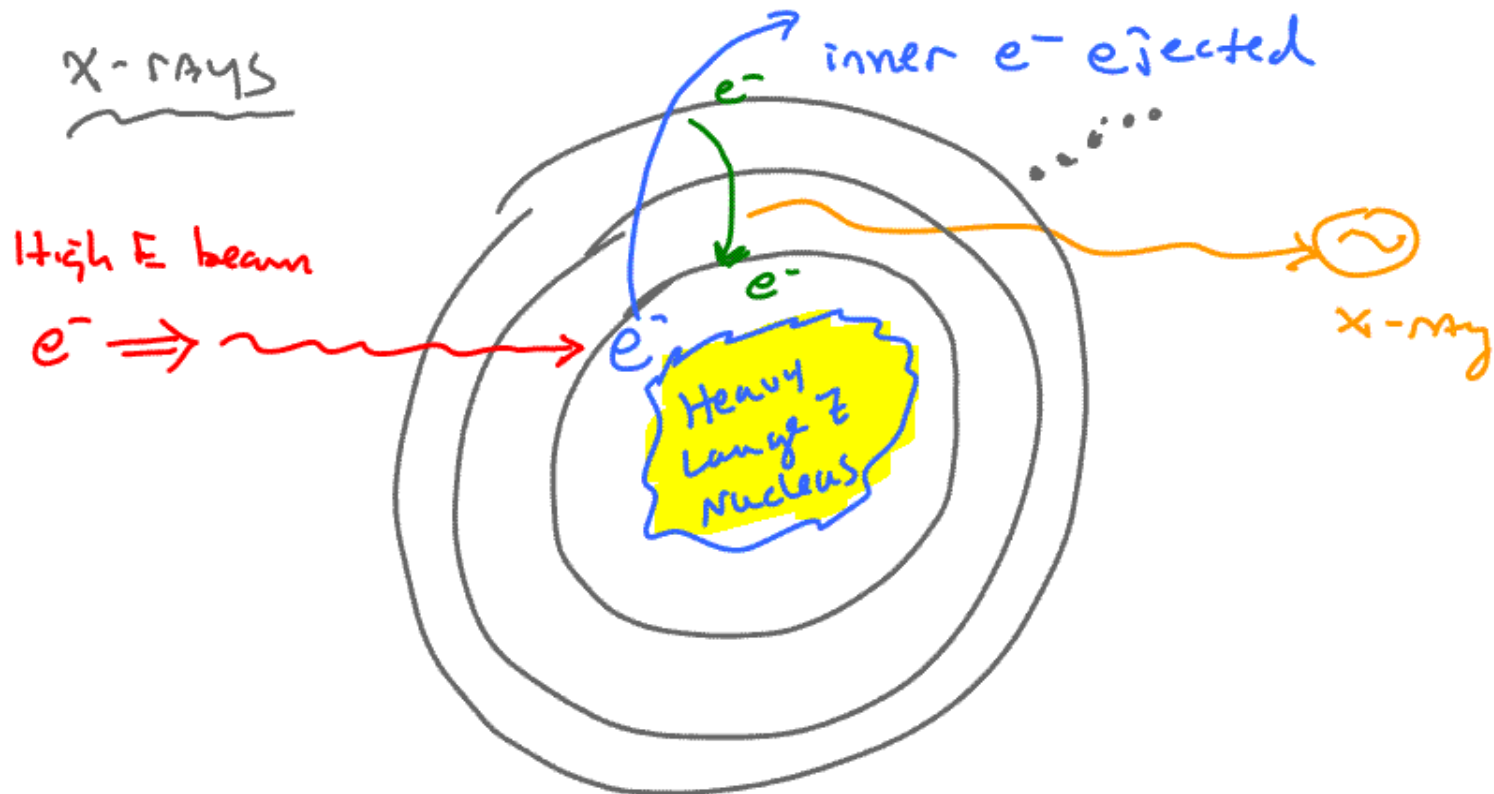
e^- energy depends on
orientation of spin
w/ respect to orbital
magnetic moment
(or Angular Momentum)

\Rightarrow Fine Structure

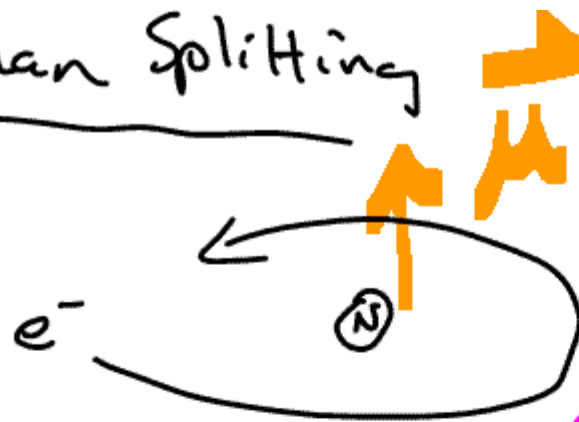


Hyperfine Splitting

Interaction between orbital magnetic moment of e^- and spin magnetic moment of nucleus



Zeeaman Splitting



Bohr
Magnetron

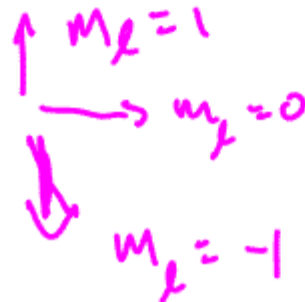
$$\vec{\mu} = \frac{\mu_B}{\hbar} \vec{L} = \left(\frac{e}{2m}\right) \vec{L}$$

2p state $\rightarrow l=1$

$m_l = +1, 0, -1$

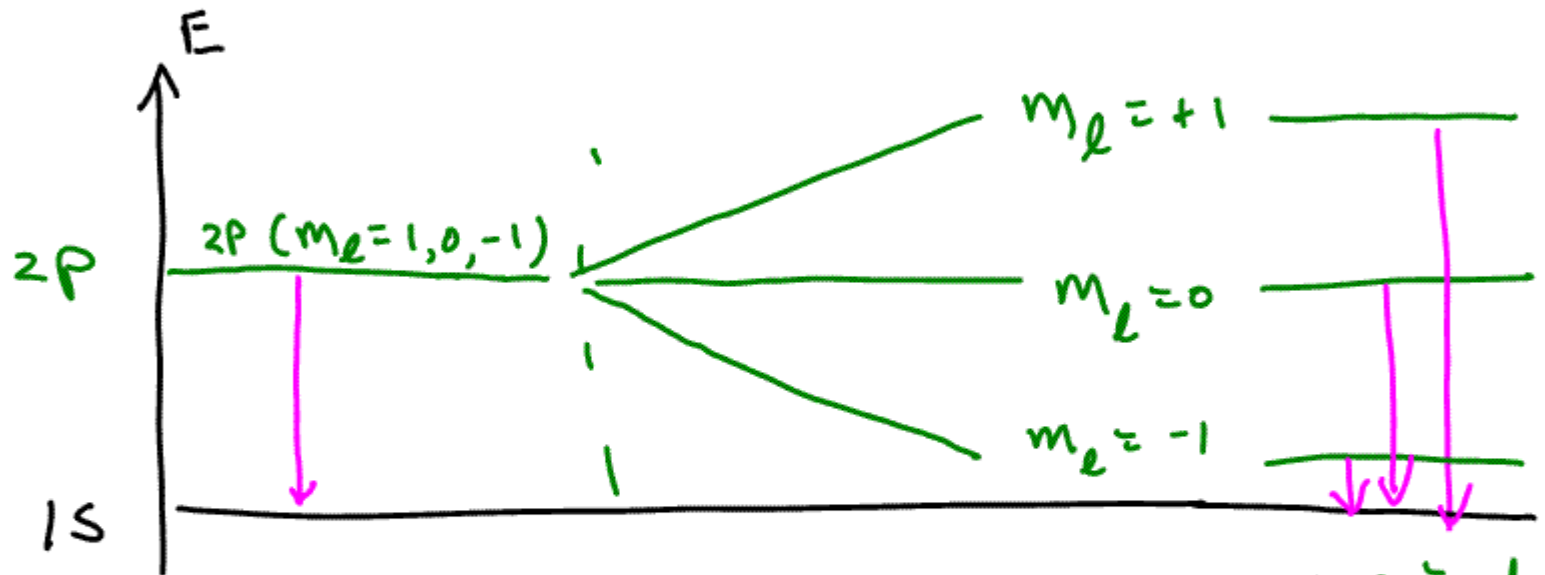
Let \vec{B} define Z axis

$\Delta x \Delta p \sim \hbar$



$$V = -\vec{\mu} \cdot \vec{B}$$

$$V \sim \frac{e}{2m} \vec{L} \cdot \vec{B} = \mu_B m_l B$$



$B=0$ B fixed
B → increase



$$\Delta\lambda = \frac{hc}{\mu_B B}$$

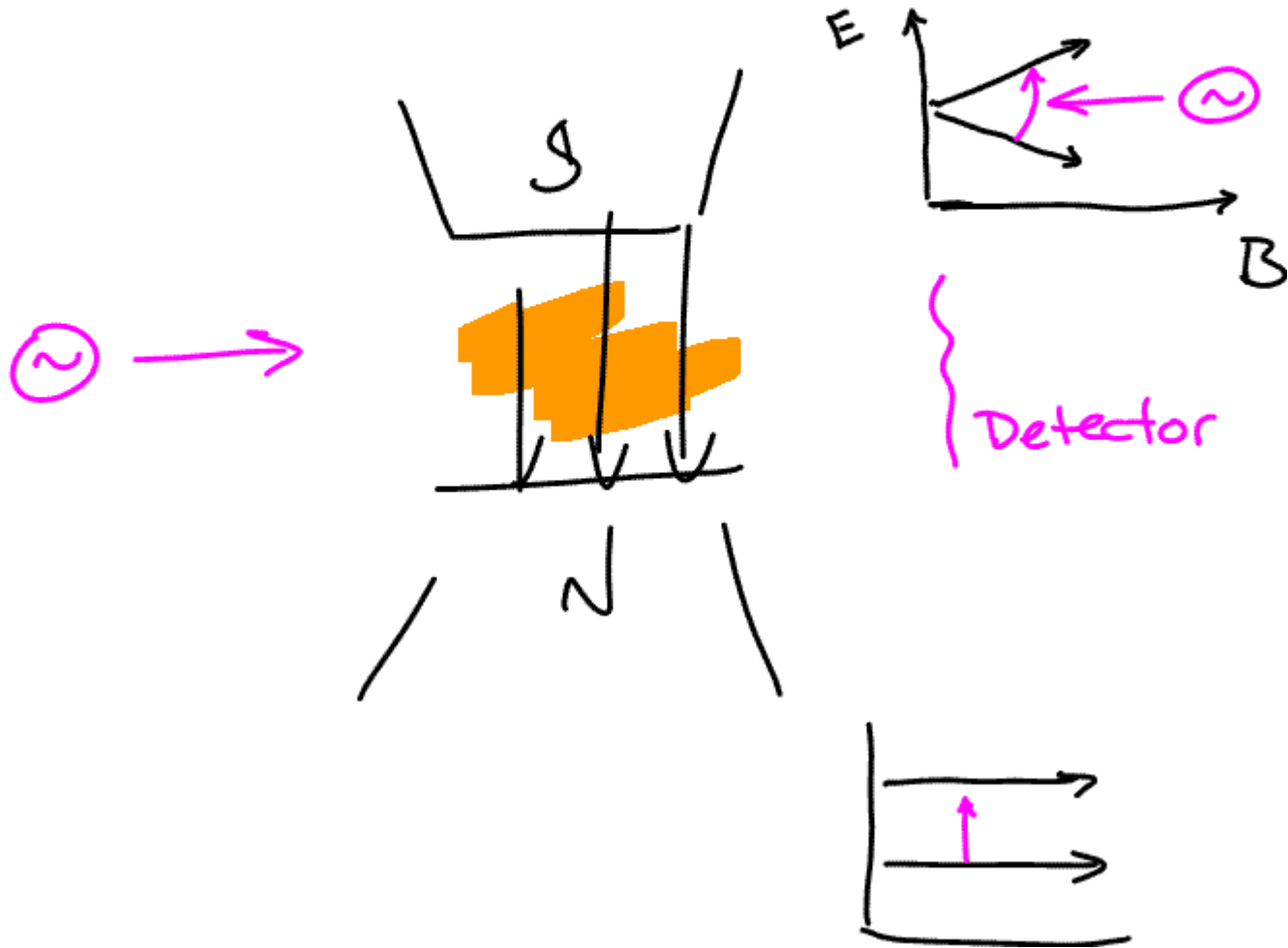


Electron Spin Resonance

ESR

Nuclear Magnetic Resonance

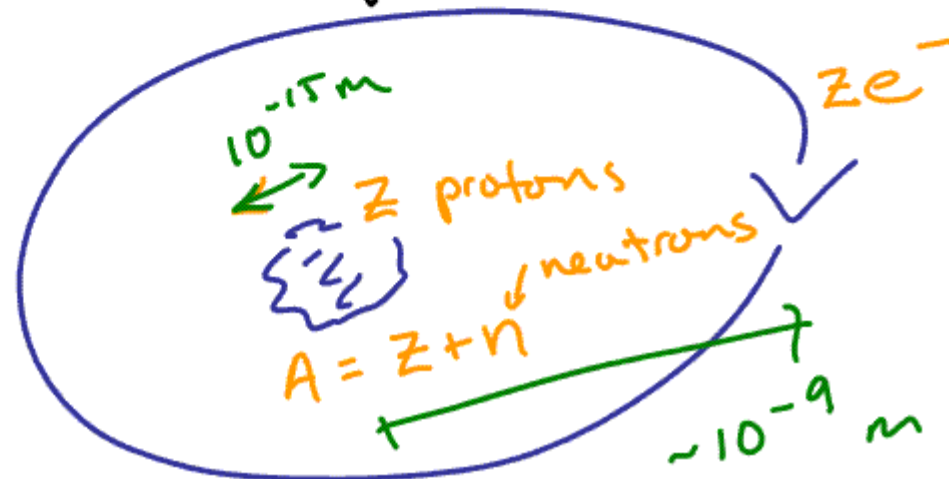
NMR



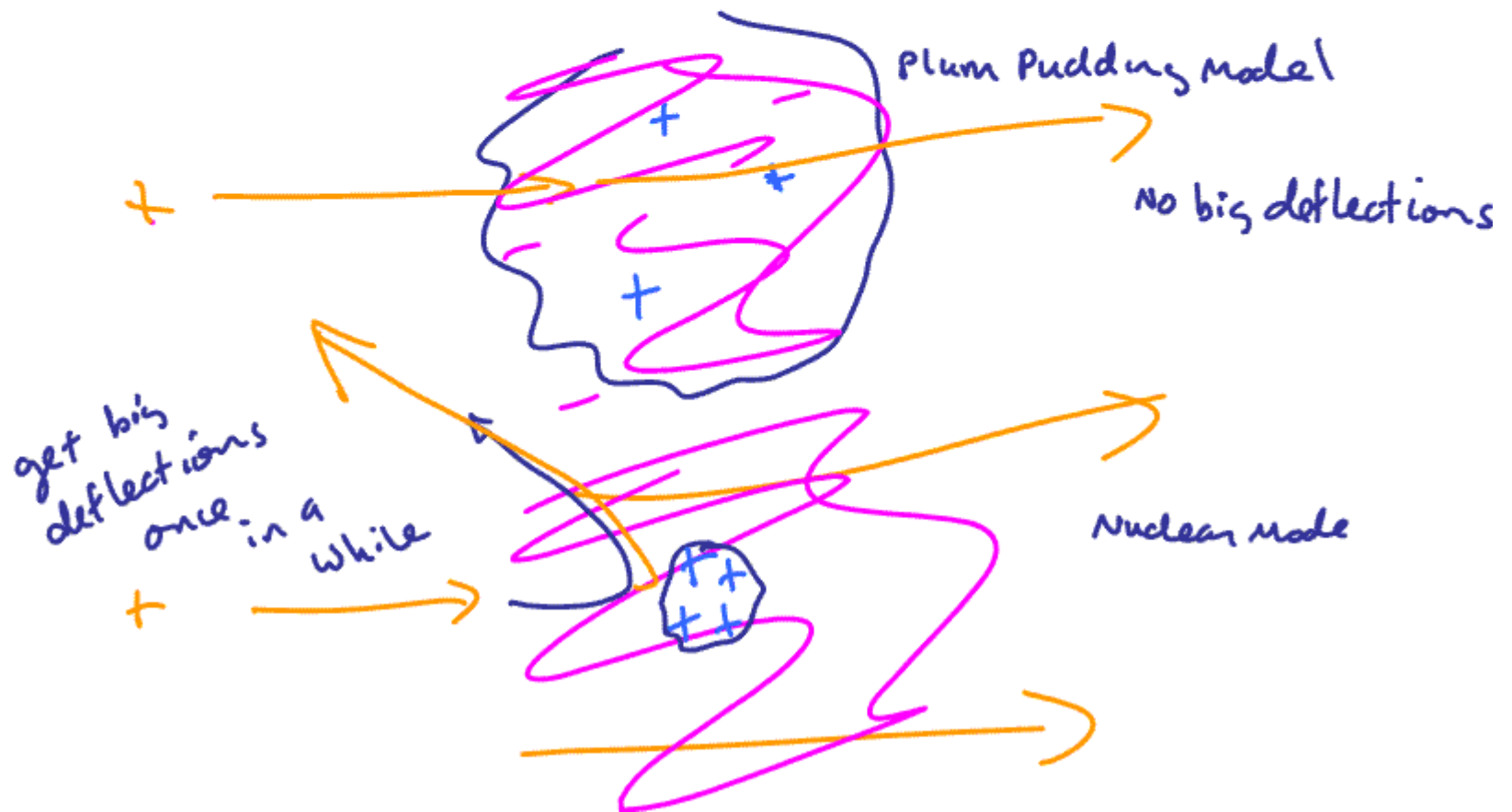


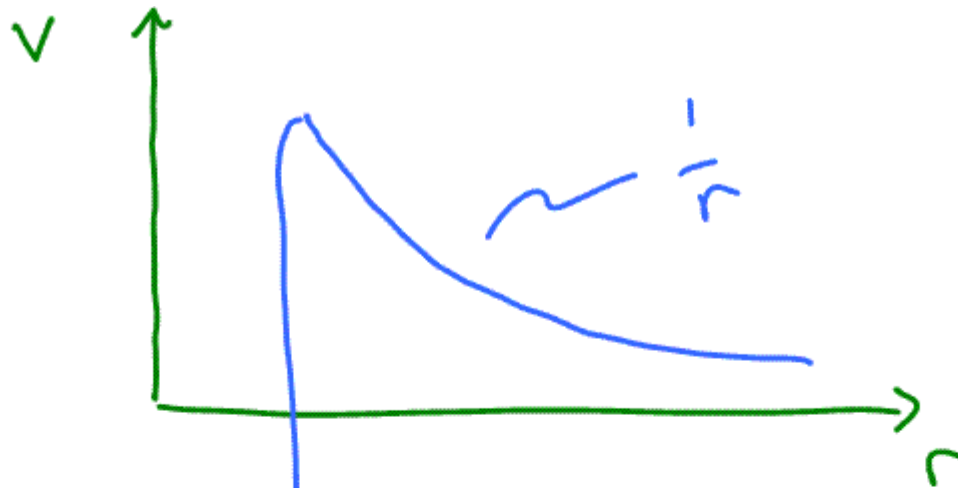
Nuclear Physics

Nuclear Model of Atom



Nuclear Model of Atom From Rutherford





New force

Strong Nuclear force

10^{-15} m

Nuclear radius $\sim 1-10$ Fermi (10^{-15} m)

Nuclear density is constant

$$\frac{A}{\frac{4}{3}\pi R^3} \sim \text{const.}$$

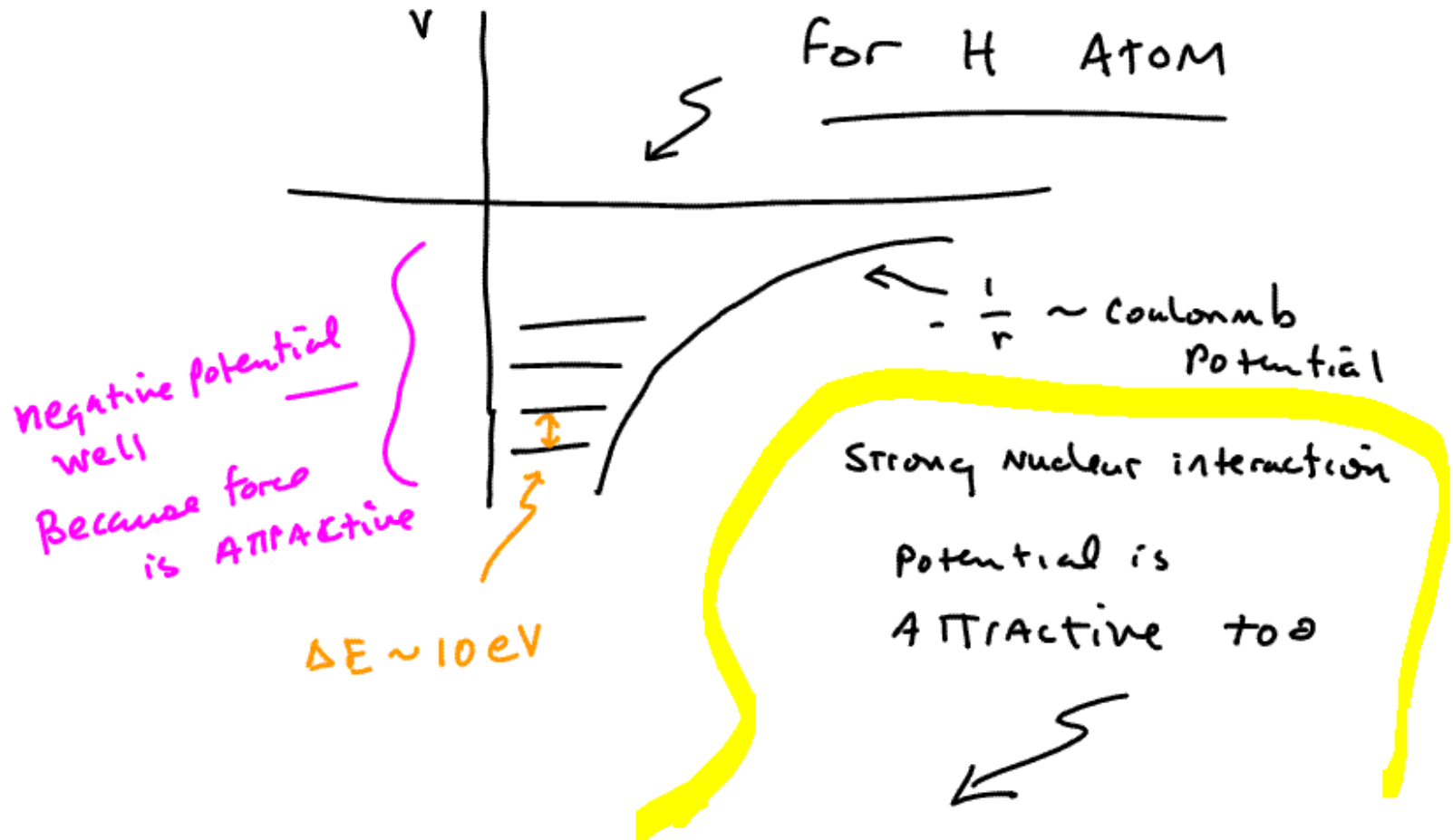
$$A \sim R^3$$

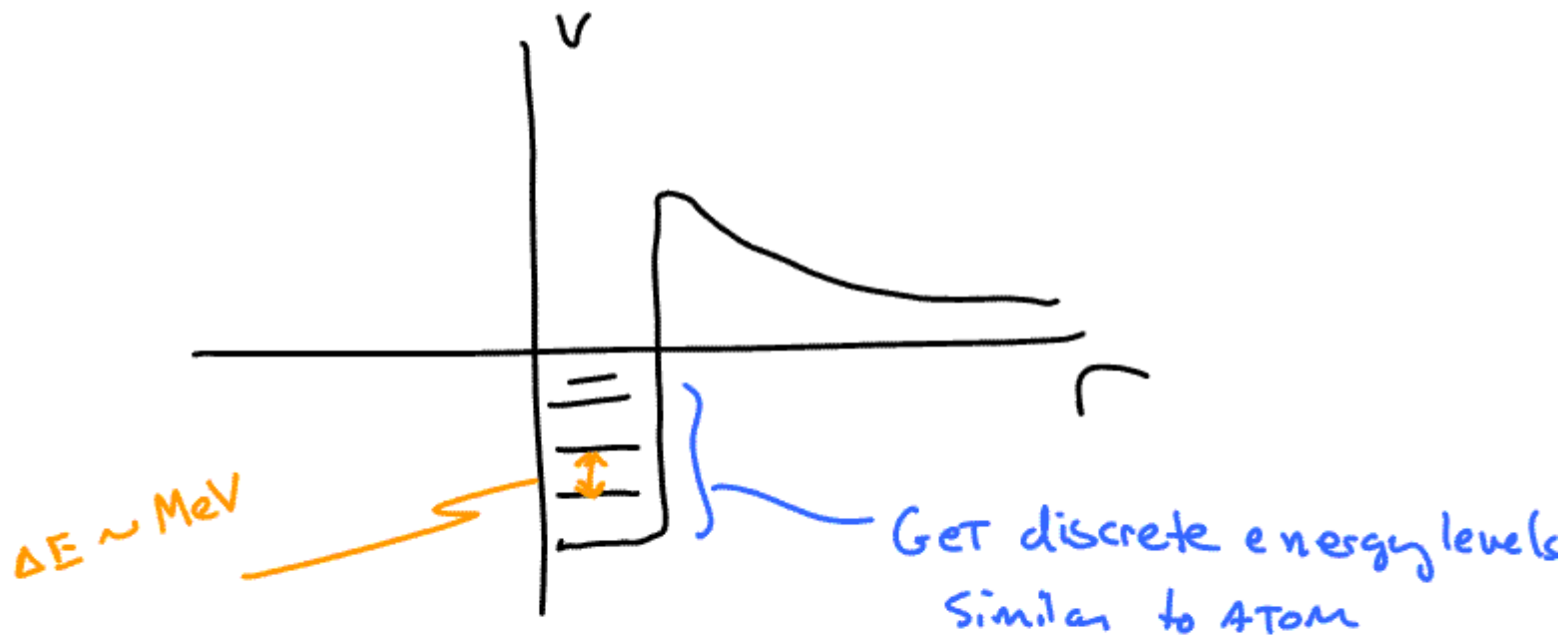
$$R \sim A^{1/3}$$

$$\rho \sim 10^{17} \text{ kg/m}^3$$

At this density

$M_{\text{Earth}} \rightarrow$ ball 140 m in radius





p, n Spin $1/2$

fill shells w/
Protons + neutrons
in much
the same way
we fill atomic
shells

