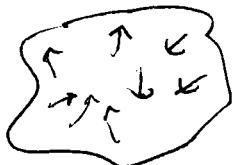


Magnetic fields and Matter

NOT really covered here but ---

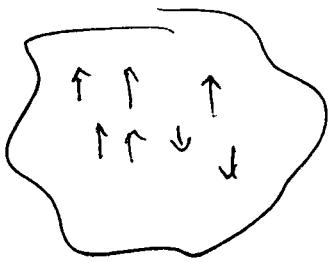
Paramagnetism

$$B=0$$



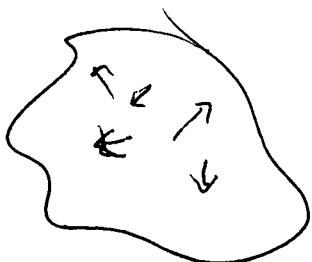
Inherent (randomly-oriented)
Magnetic Dipole

$$B_{ext}$$

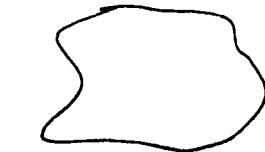


Moments line up
w/ field
(slight enhancement)

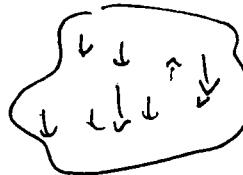
$$B=0$$



Diamagnetism



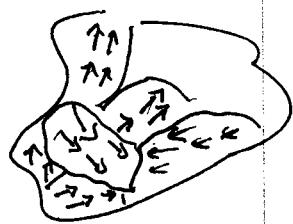
No inherent
Magnetic Dipole



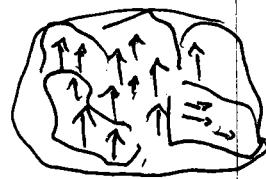
induced moments
opposite to
field

weakens field
slightly

Ferromagnetism

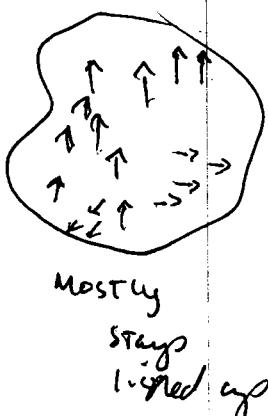


Strong Natural
Magnetic Dipoles
Domains



Domains mostly
line up

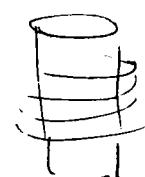
Strong
enhancement



MOSTLY
Stays
lined up

Electromagnets

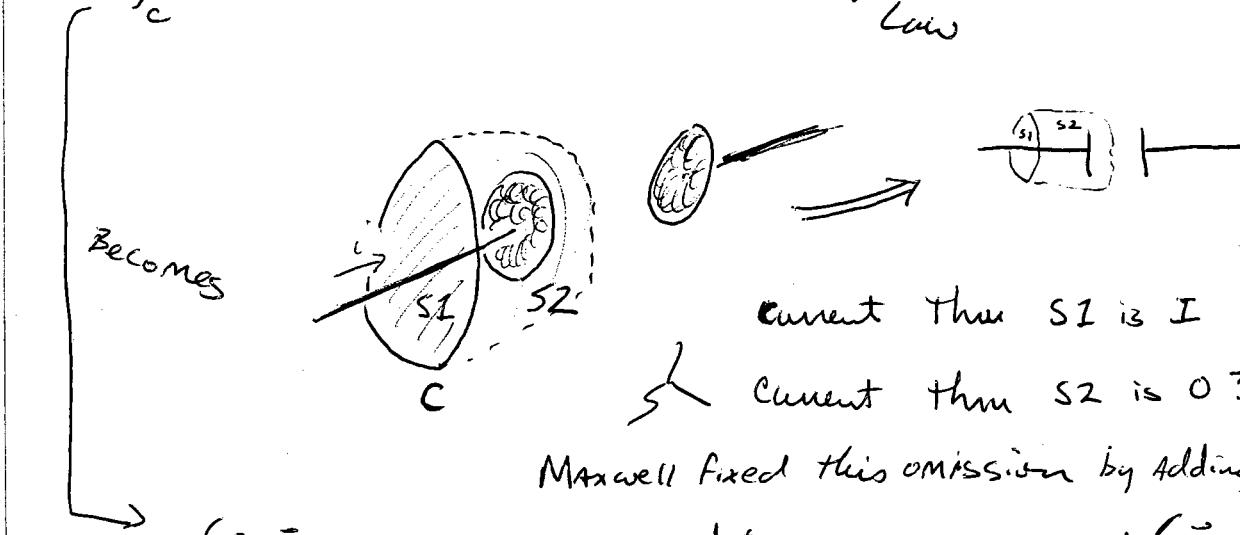
usually have an iron core



SATURATES
at
2 Tesla

~~Please~~ Review of what we have learned in E+M

- ① $\int_S \vec{E} \cdot d\vec{A} = \frac{Q_{\text{enc}}}{\epsilon_0}$ Gauss's Law
 (Gauss's law for B)
- ② $\int_S \vec{B} \cdot d\vec{A} = 0$ No Magnetic Monopoles
- ③ $\int_C \vec{E} \cdot d\vec{l} = -\frac{d\Phi_B}{dt} = -\frac{d}{dt} \int_S \vec{B} \cdot d\vec{A}$ Faraday's Law
 (induced EMF's from changing Magnetic Flux)
- $\int_C \vec{B} \cdot d\vec{l} = \mu_0 I_{\text{enc}}$ Ampere's Law



$$④ \int_C \vec{B} \cdot d\vec{l} = \mu_0 I + \mu_0 \epsilon_0 \frac{d\Phi_E}{dt} = \mu_0 I + \mu_0 \epsilon_0 \frac{d}{dt} \int_S \vec{E} \cdot d\vec{A}$$

Equations ①-④ are the integral form of Maxwell's Equations

These equations describe classical electromagnetic phenomena

This is most of what you have in your life experience

Note again the slight asymmetry between \vec{E} and \vec{B}

Implications of Maxwell's Equations - radiation, light

Conceptual:

$$\int_C \vec{E} \cdot d\vec{l} \sim \frac{d}{dt} \int_S \vec{B} \cdot d\vec{A} \rightarrow \frac{d\vec{B}}{dt} \rightarrow E \text{ varying w/t}$$

$$\int_C \vec{B} \cdot d\vec{l} \sim \frac{d}{dt} \int_S \vec{E} \cdot d\vec{A} \rightarrow \frac{d\vec{E}}{dt} \rightarrow B \text{ varying w/t}$$

Changing E induces B
Changing B induces E

Circular
Radiation

Mathematical:

Maxwell's Equations

Equivalent
TO
using Vector calculus
Tricks

Six
coupled
Partial
differential
Equations

$$\frac{\partial^2 E_x}{\partial x^2} + \frac{\partial^2 E_x}{\partial y^2} + \frac{\partial^2 E_x}{\partial z^2} = \frac{\partial^2 E_x}{\partial t^2} \mu_0 \epsilon_0$$

$$\frac{\partial^2 E_y}{\partial x^2} + \frac{\partial^2 E_y}{\partial y^2} + \frac{\partial^2 E_y}{\partial z^2} = \frac{\partial^2 E_y}{\partial t^2} \mu_0 \epsilon_0$$

$$\frac{\partial^2 E_z}{\partial x^2} + \dots$$

$$\frac{\partial^2 B_x}{\partial x^2} + \frac{\partial^2 B_x}{\partial y^2} + \frac{\partial^2 B_x}{\partial z^2} = \frac{\partial^2 B_x}{\partial t^2} \mu_0 \epsilon_0$$

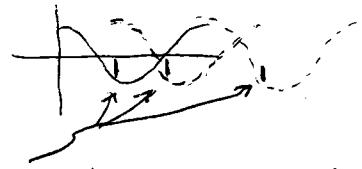
$$\frac{\partial^2 B_y}{\partial x^2} + \dots$$

$$\frac{\partial^2 B_z}{\partial x^2} + \dots$$

Looks quite complicated ... And can be complicated
lets try a simple situation and see what happens
Physics depends on "Boundary Conditions"

Consider the phase $\omega(t \pm \frac{z}{c}) = \text{constant}$

$$\frac{\partial(\text{phase})}{\partial t} = \omega$$



$$\frac{\partial(\text{phase})}{\partial z} = \pm \frac{\omega}{c}$$

Points w/ same phase as
wave propagates = $3 \times 10^8 \text{ m/s}$

$$\frac{\frac{\partial(\text{phase})}{\partial t}}{\frac{\partial(\text{phase})}{\partial z}} = \frac{\partial z}{\partial t} = \pm c$$

Velocity of propagation of condition of
constant phase

- or - phase velocity of light

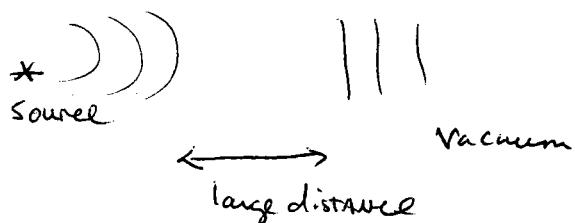
Maxwell's eqns predicts velocity of light in vacuum?

implications of Maxwell's eqns -

With somewhat more Mathematics than we'll want to use
one can derive from Maxwell's eqns in a vacuum

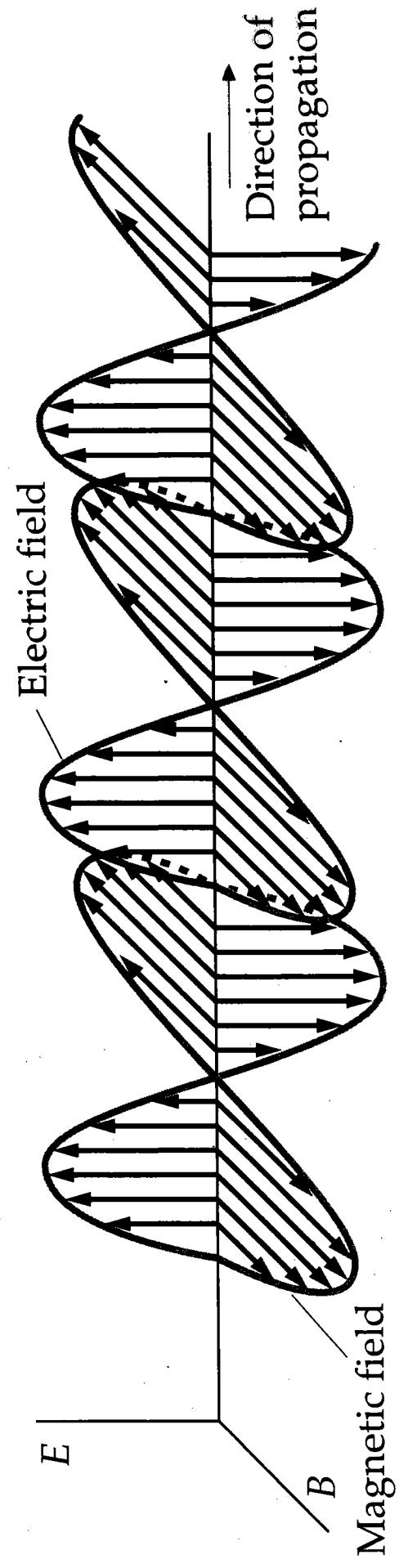
- Radiation (light, x rays, infrared, UV ... etc)
made up of electric and magnetic fields

For case



- Electromagnetic waves travel w/ $V = \frac{1}{\sqrt{\epsilon_0 \mu_0}} = c$

- $E + B$ fields \perp to each other and direction of travel
TRANSVERSE waves



~~-~~ - \vec{E} , \vec{B} in phase w.r.t one another

- $|\vec{E}| = C/B$

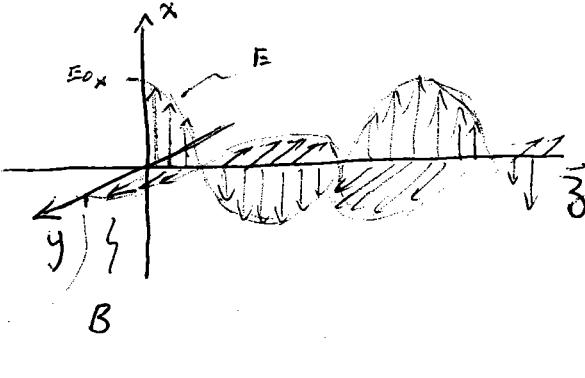
- Harmonic in both space + Time

(waves w.r.t space + time dependence)

e.g.

$$E_x = E_{0x} \cos [\omega_x (t \pm \frac{z}{c})]$$

$$B_y = \frac{E_{0x}}{c} \cos [\omega_x (t \pm \frac{z}{c})]$$



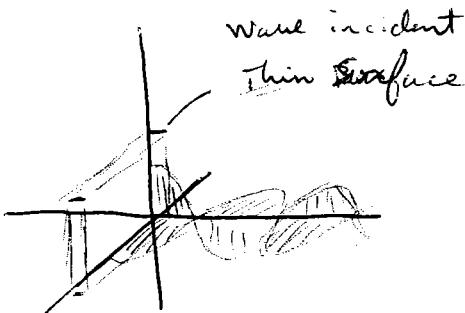
$$\frac{E_{0x}}{c} = B_{0x}$$

$$B$$

ω is frequency of the electromagnetic radiation

some eqns \rightarrow $\nu = \frac{1}{T}$ $\omega = \frac{2\pi}{T}$ $k = \frac{2\pi}{\lambda}$ $c = \omega/k$ $c = \lambda\nu$

Sometimes see phase written as $\omega t \pm k z$



- What is Power incident = intensity? Area

$$I = \text{Power} / \text{Area}$$

$$U = U_E + U_B = \frac{1}{2} \epsilon_0 E^2 + \frac{B^2}{2\mu_0}$$

$$C^2 = \frac{1}{\mu_0 \epsilon_0} \quad \text{and} \quad E = CB$$

$$= \epsilon_0 E^2$$

$$= \frac{EB}{\mu_0 C}$$

Replace $E_{\text{instantaneous}}$, $B_{\text{instantaneous}}$

w.r.t RMS values \sim Average values

$$E_{\text{RMS}} = E_0/\sqrt{2}$$

$$B_{\text{RMS}} = B_0/\sqrt{2}$$

$$\boxed{I = \frac{1}{2} \frac{\epsilon_0 B_0}{\mu_0 C}}$$

define $\vec{S} = \frac{\vec{E} \times \vec{B}}{\mu_0}$ Poynting vector

direction gives the direction of propagation of EM wave

$$|S| = \text{intensity}$$

EM waves carry energy

Power incident is the intensity = $\frac{\text{energy}/\text{Time}/\text{Area}}{\text{unit Area}}$

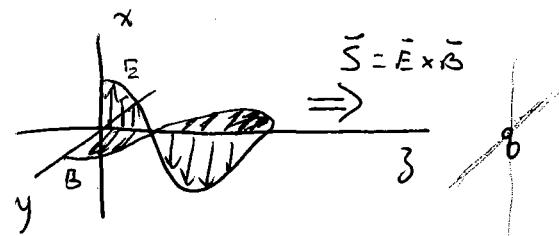
- Radio station energy to run radio (usually Amplified where picked up)

- Energy from sun for photosynthesis

- Energy to cause a photochemical reaction w/ film for x rays etc.

EM waves carry momentum

$$F = \frac{dP}{dt}$$



when

Force due to E is $gE\hat{i}$

g gets $\vec{v} = v\hat{i}$

F from $\vec{B} = (\vec{B})\hat{j}$

i's

$$F_B = g\vec{v} \times \vec{B} = gVB\hat{i} \times \hat{j} = gVB\hat{k}$$

Force is along direction of wave propagation

$$\text{relativity } F^2 = (mc^2)^2 + (pc)^2$$

$$F = pc$$

$$|P| = \text{Momentum carried by EM wave} = \frac{\text{Energy}}{c}$$

$$\text{Intensity} = \frac{\text{Energy}}{\text{Time}} / \text{Area}$$

$$\frac{\text{Momentum}}{\text{Time}} / \text{Area} = \text{Force}$$

$$\frac{I}{c} = \frac{\text{Energy}}{c} / \text{Time} / \text{Area} = \frac{\text{Force}}{\text{Area}} = \text{Pressure}$$

Radiation exerts pressure on charged particles!

- Sailing Space ships
- radiation pressure \leftrightarrow gravity

Equilibrium in Sun

Stellar evolution in a nutshell

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