

Physics 100 - Nov. 5, 2007

Exam 2 - Nov. 14

One week from Wednesday

- Hoyt during regular class slot
- Calculator
- Note card
- Material from last time thru STARS ^{nuclear phys.}
(last lecture)
- Will email w/ material coverage soon
- Q + A session ... will do

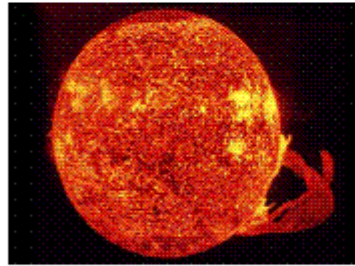
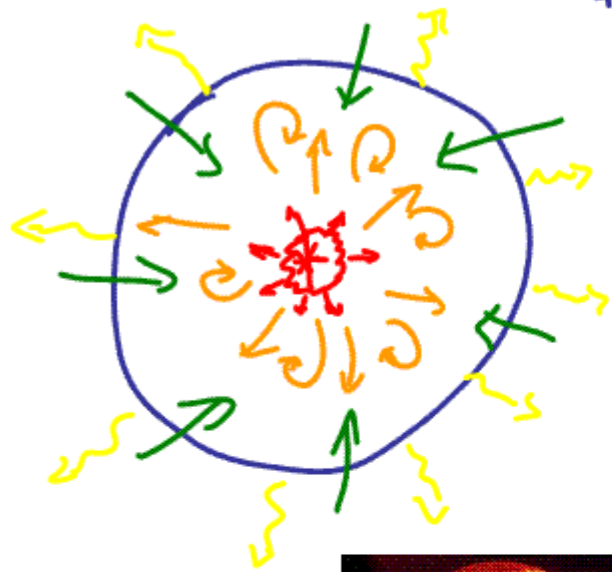
■ Would like to meet w/ presentation groups

... All before Thanksgiving

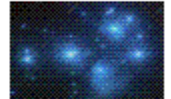
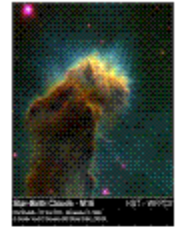
- Select spokesperson + get in touch w/ me
to set up time

last time

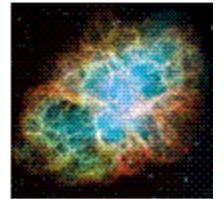
intimate relation between the large + small drive the stars and synthesize the elements



- gravity pulls in
- heat (radiation pressure) pushes out
- "Thermonuclear" fusion reactions in core



- STARTS w/ $H \rightarrow He$ but can work way up to ^{56}Fe
- Heavier elements formed in Supernovae

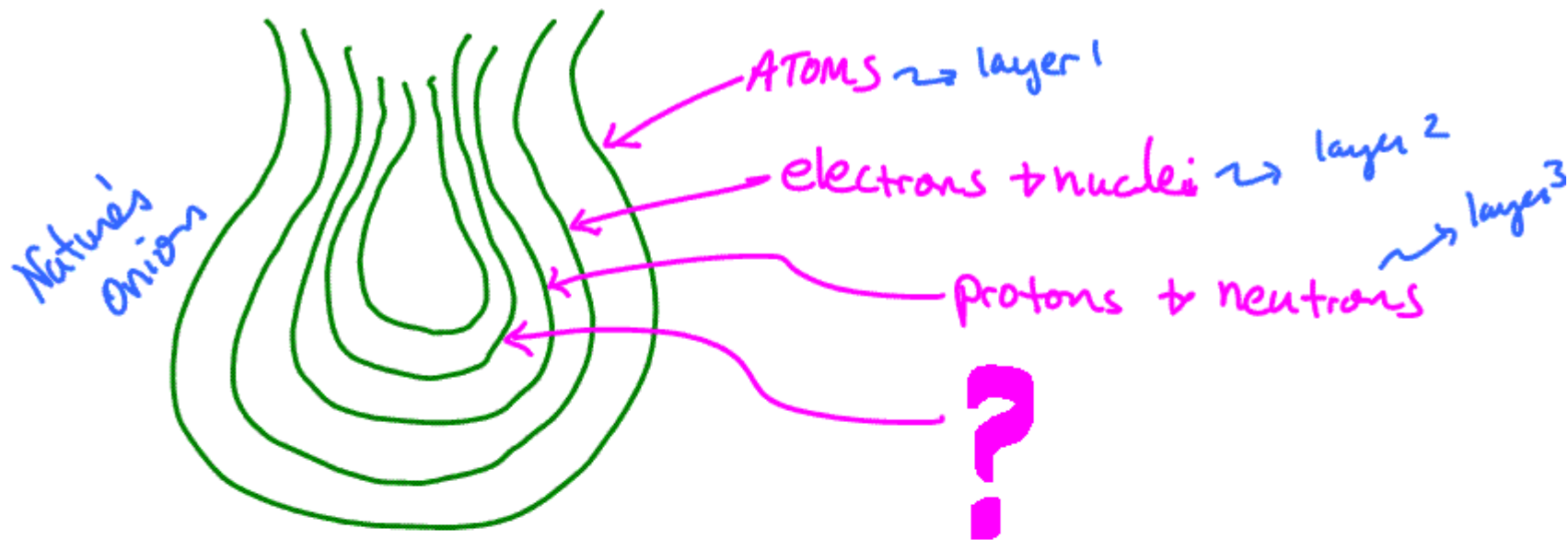


Rotation comes from Conservation of "Angular Momentum"

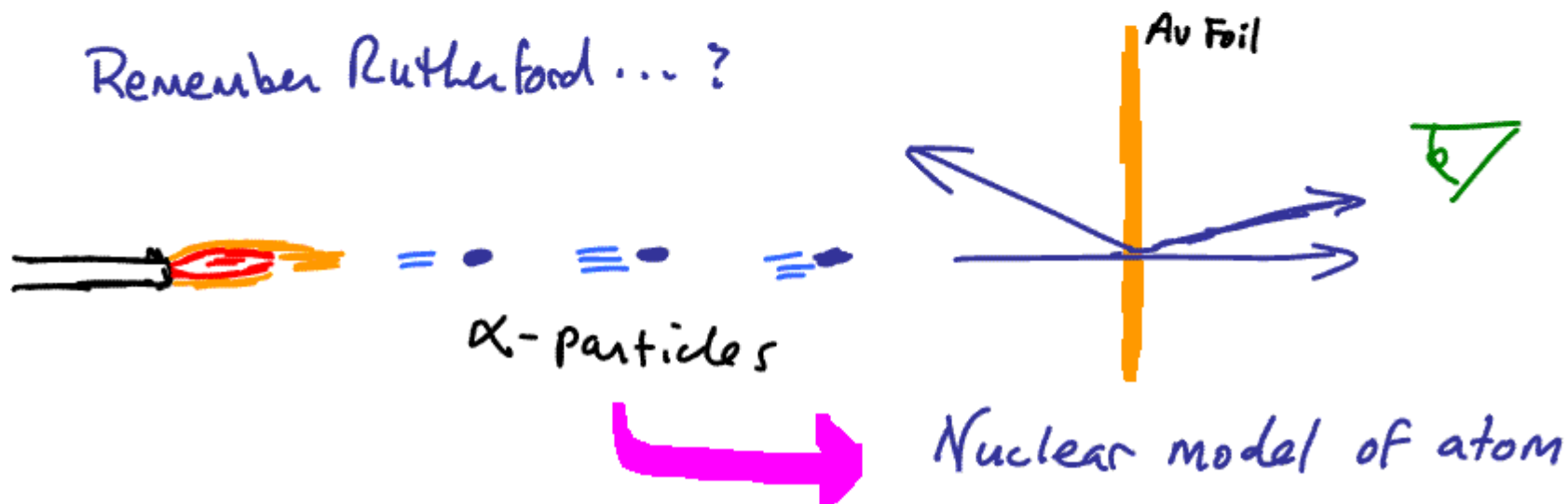


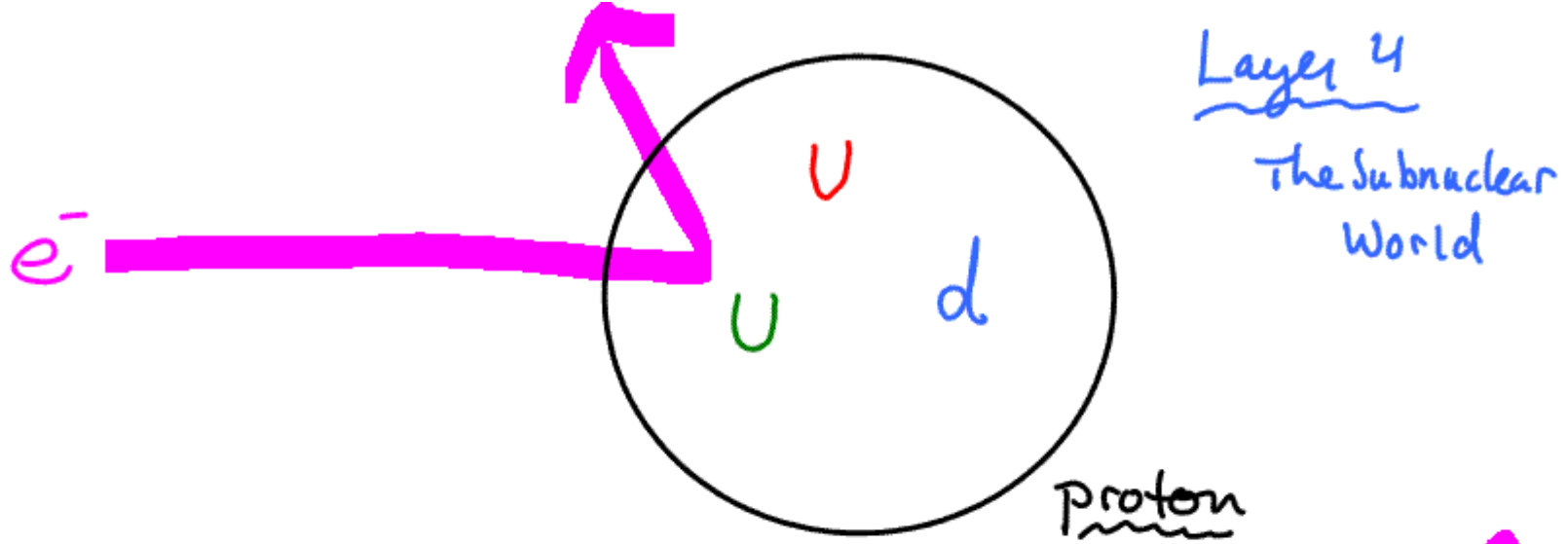
- white dwarfs, neutron stars, Black Holes

STAR forms
planets form



Remember Rutherford ... ?





$u \equiv$ up quark $q = +\frac{2}{3}|e|$

$d \equiv$ down quark $q = -\frac{1}{3}|e|$

Particle (high energy) Physics

New Taxonomy of Particles
New way of looking at forces

a little overwhelming at 1st ...

The Standard Model!

Places to learn more: Particle and nuclear physics links

<http://pdg.lbl.gov>

<http://particleadventure.org>

<http://www.slac.stanford.edu/gen/edu/aboutslac.html>

<http://www.bnl.gov/bnlweb/sciindex.html>

<http://www.bnl.gov/rhic/>

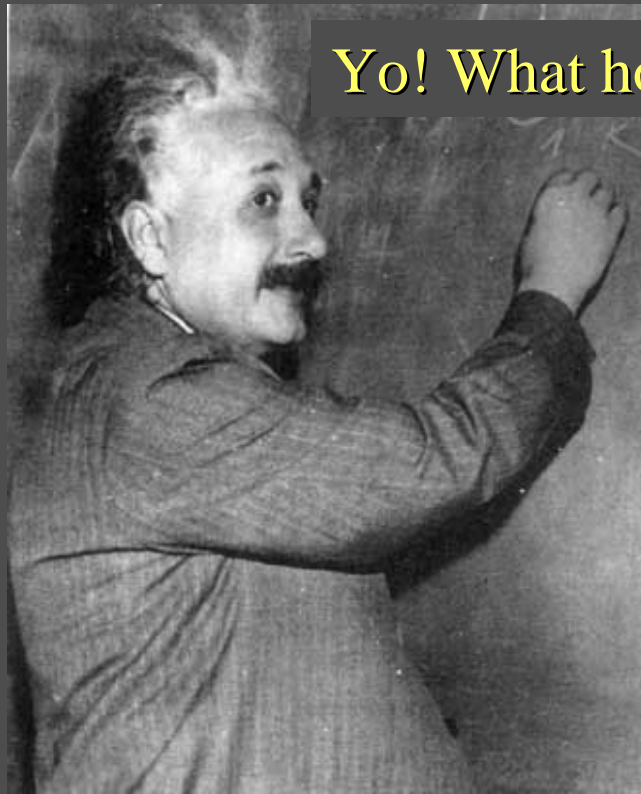
<http://public.web.cern.ch/public/>

<http://www.fnal.gov/>

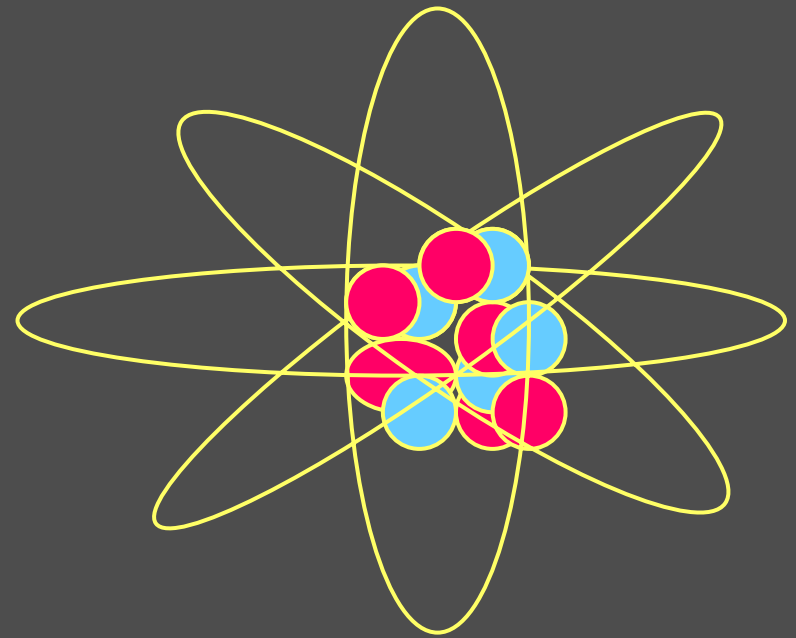
<http://www.er.doe.gov/production/henp/np/index.html>

<http://www.science.doe.gov/hep/index.shtm>

Inquiring minds want to know ...



Yo! What holds it together?

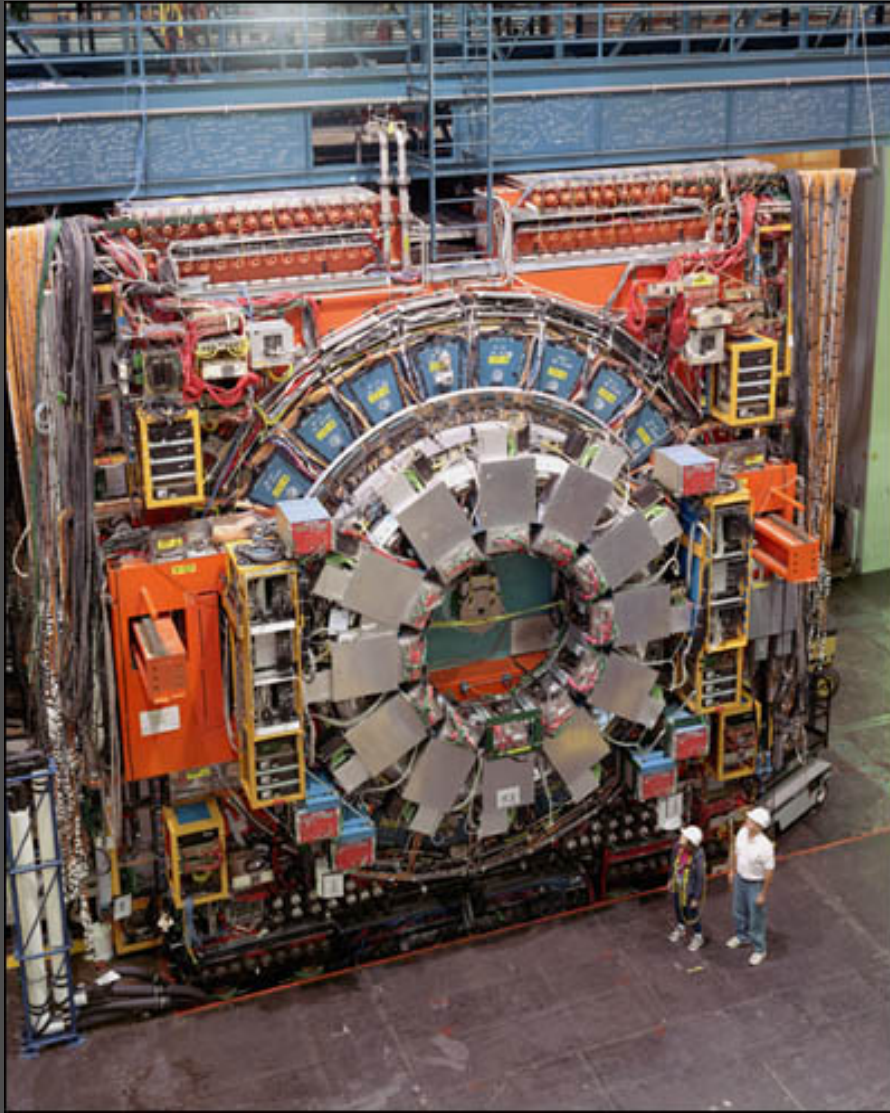




Fermi National Accelerator Laboratory (near Chicago)



CDF



Minos



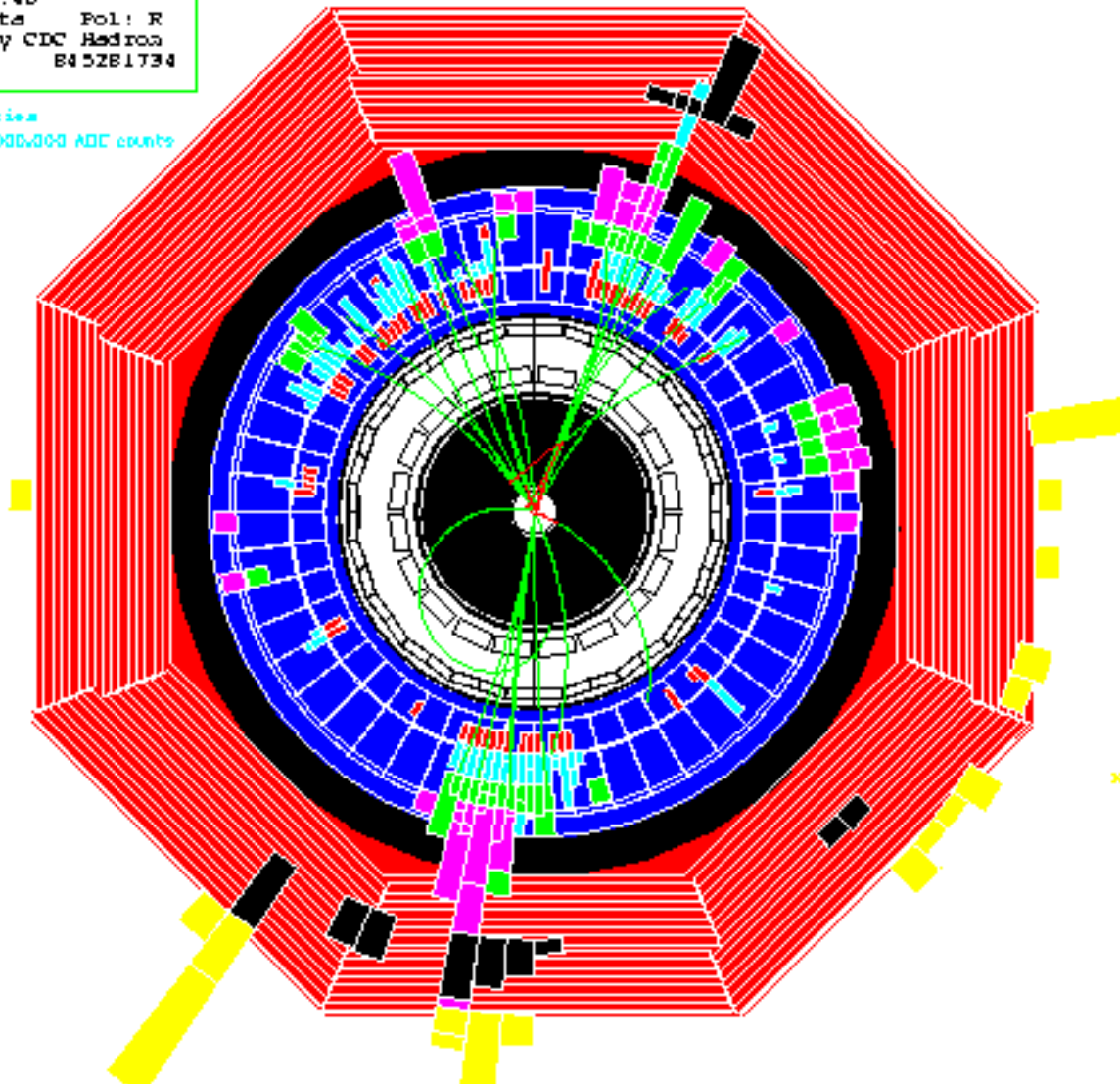
Stanford Linear Accelerator Center



Event display from the SLD experiment at SLAC

Run 2057E, EVENT 779
23-MAR-1993 12:40
Source: Run Data Fol: R
Trigger: Energy CDC Hadron
Beam Crossing 845281734

Xal hit properties
5.025 E_{had} $15000/000$ ADC counts



XAL Subsystems
IUM 0
IUM 1
IAC IM1
IAC IM2
IAC RAD1
IAC RAD2
MIC 1
MIC 2



$$I^G(J^{PC}) = 1^-(0^{-+})$$

Mass $m = 134.9766 \pm 0.0006$ MeV ($S = 1.1$)

$m_{\pi^+} - m_{\pi^0} = 4.5936 \pm 0.0005$ MeV

Mean life $\tau = (8.4 \pm 0.6) \times 10^{-17}$ s ($S = 3.0$)

$c\tau = 25.1$ nm

For decay limits to particles which are not established, see the appropriate Search sections (A^0 (axion), and Other Light Bosons (X^0) Searches, etc.).

week ending
4 JUNE 2004

$\rightarrow \omega\gamma(1S)$

E. Coan,² Y.S. Gao,² F. Liu,²
Dorjkhaidav,³ R. Mountain,³
Mahmood,⁴ S.E. Csorna,⁵
Das,⁷ A. Shaikro,⁷ W.M. Sun.⁷

S 30 MARCH 1998

ISS

mendolia,²⁷ D. Amidei,²⁰ J. Antos,³³
⁸ M. Atac,⁷ P. Azzi-Bacchetta,²⁵

1 MARCH 1999

Measurement
of π^0 radiative decay

Itow,¹ T. Kajita,¹ J. Kameda,¹
and S. Nishimura,¹ A. Oishi,¹

26 MAY 1975

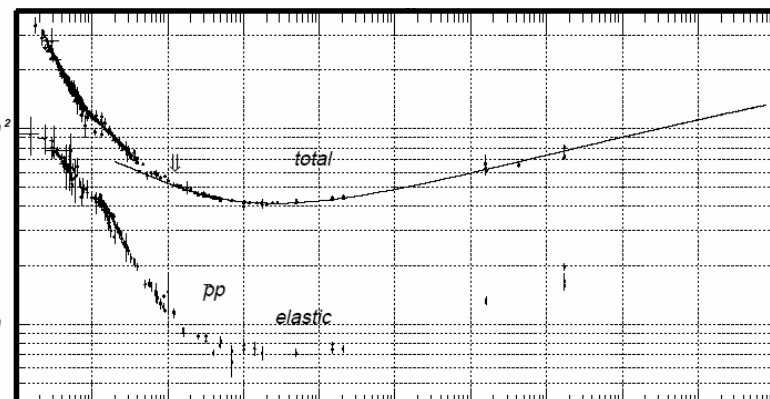
π^0 DECAY MODES	Fraction (Γ_i/Γ)	Scale factor/ Confidence level	p (MeV/c)
2γ	$(98.798 \pm 0.032) \%$	$S=1.1$	67
$e^+e^-\gamma$	$(1.198 \pm 0.032) \%$	$S=1.1$	67
γ positronium	$(1.82 \pm 0.29) \times 10^{-9}$		67
$e^+e^+e^-e^-$	$(3.14 \pm 0.30) \times 10^{-5}$		67
e^+e^-	$(6.2 \pm 0.5) \times 10^{-8}$		67
4γ	< 2	$\times 10^{-8}$ CL=90%	67
$\mu\bar{\mu}$	< 8.3	$\times 10^{-7}$ CL=90%	67
$\mu_0\bar{\mu}_0$	< 1.7	$\times 10^{-6}$ CL=90%	67
$\mu^+\mu^-\mu^+$	< 3.1	$\times 10^{-6}$ CL=90%	67
$\mu^+\mu^-\mu^-$	< 2.1	$\times 10^{-6}$ CL=90%	67
$\gamma\mu\bar{\mu}$	< 6	$\times 10^{-4}$ CL=90%	-

Charge conjugation (C) or Lepton Family number (LF) violating modes

3γ	C	< 3.1	$\times 10^{-8}$ CL=90%	67
μ^+e^-	LF	< 3.8	$\times 10^{-10}$ CL=90%	26
μ^-e^+	LF	< 3.4	$\times 10^{-9}$ CL=90%	-
$\mu^+e^- + \mu^-e^+$	LF	< 1.72	$\times 10^{-8}$ CL=90%	26

Cross section (mb)

Cross section (mb)



Laboratory beam momentum (GeV/c)

$\psi(3095)^{\dagger}$

ischer, D. Fryberger, G. Hanson,
, D. Lyon, C. C. Morehouse,
R. F. Schwitters,

ford, California 94305

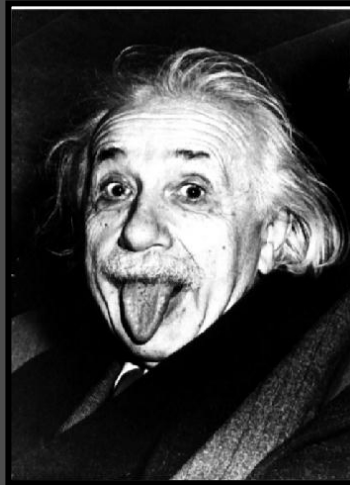
G. Golhaber, J. A. Kadyk,
Trilling, J. S. Whitaker,

ifornia, Berkeley, California 94720

near 3095 MeV. The

What forces exist in nature?

What is a force?



How do they interact?

How do forces change with energy or temperature?

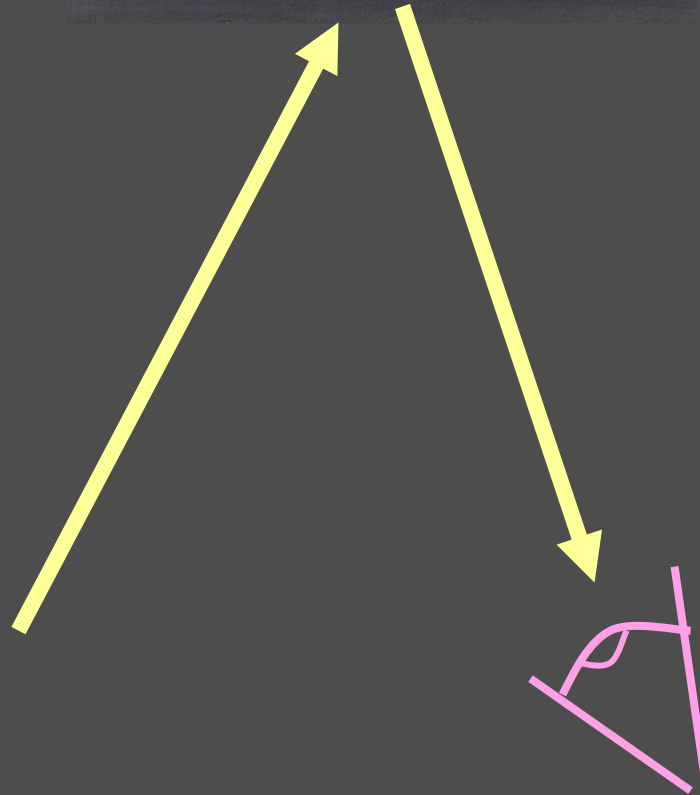
How has the universe evolved?

Mini-Ph.D. – Quantum Mechanics 101

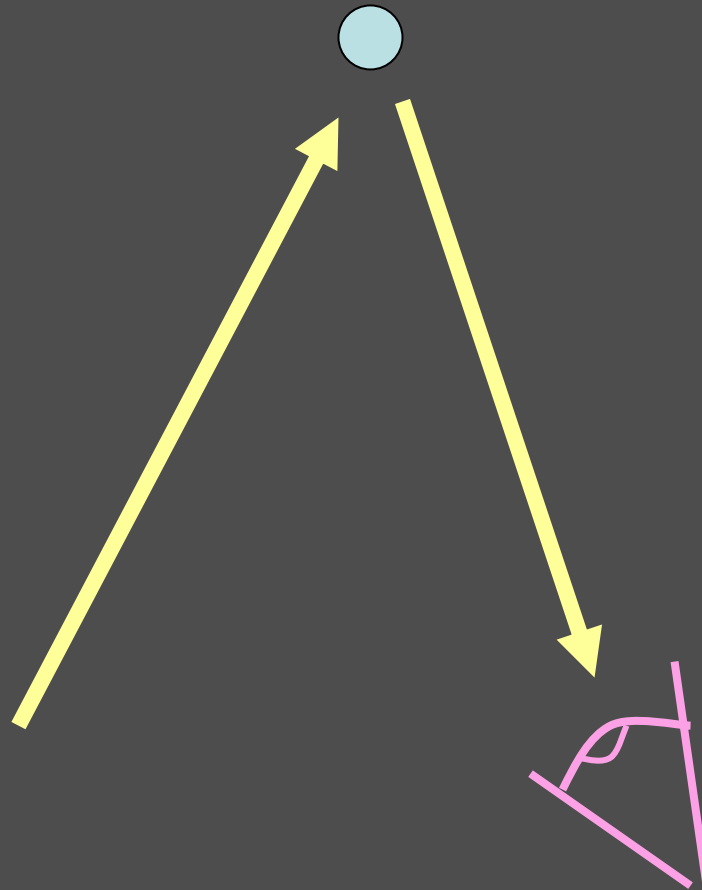
Lesson 1:

Size actually does matter.

Determine the position and velocity of a car ... no problem

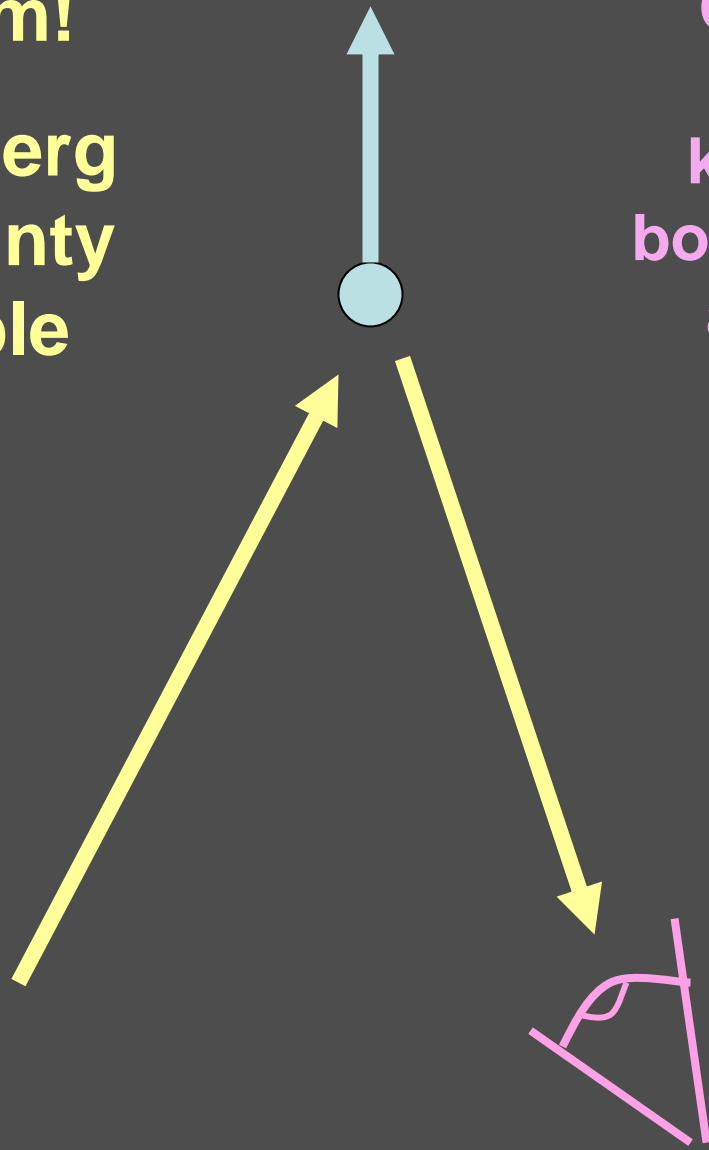


**Determine the position and velocity
of a small particle ... no problem**



Problem!
Heisenberg
uncertainty
principle

Cannot have
perfect
knowledge of
both the position
and velocity

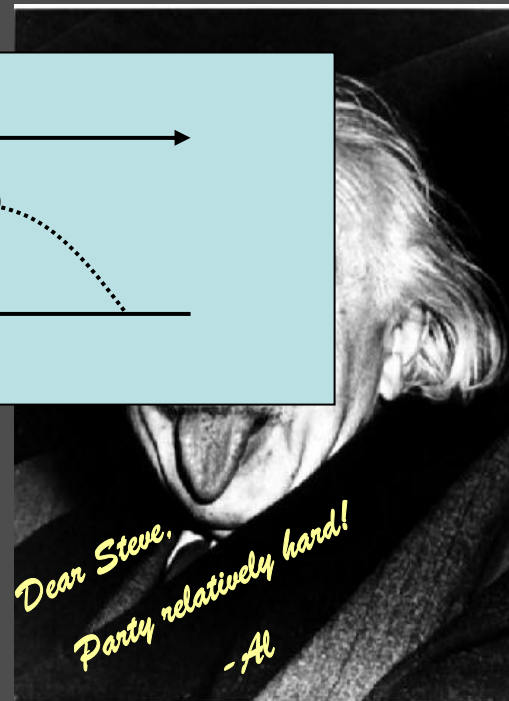
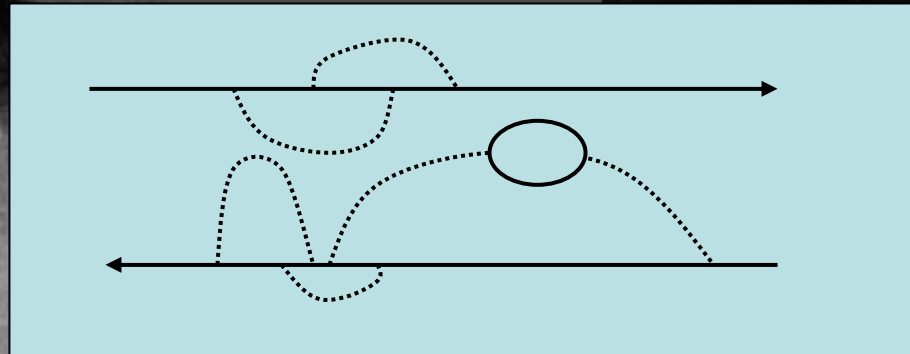
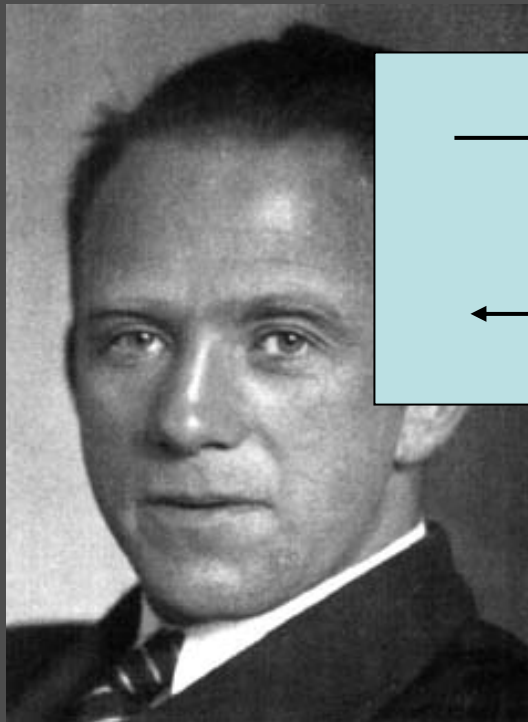
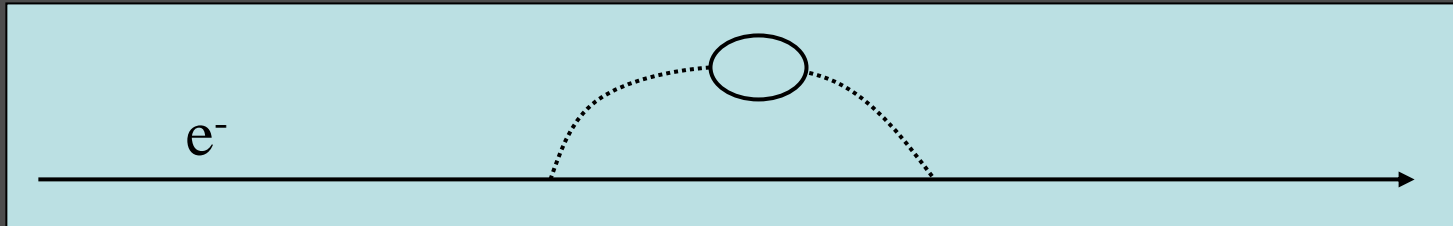


Heisenberg

The fundamental nature of forces: virtual particles

$\Delta E \Delta t \approx h$ Heisenberg

$E = mc^2$ Einstein



<i>Force</i>	<i>Source</i>	<i>Range</i>	<i>Strength</i>
<i>Gravitation</i>	mass	infinite	10^{-39}
<i>Electromagnetism</i>	Electric charge	infinite	10^{-2}
<i>Strong nuclear</i>	Color charge	10^{-15} m	1
<i>Weak nuclear</i>	Weak charge	10^{-18} m	10^{-5}

Frank and Ernest



The Standard Model of Particle Interactions

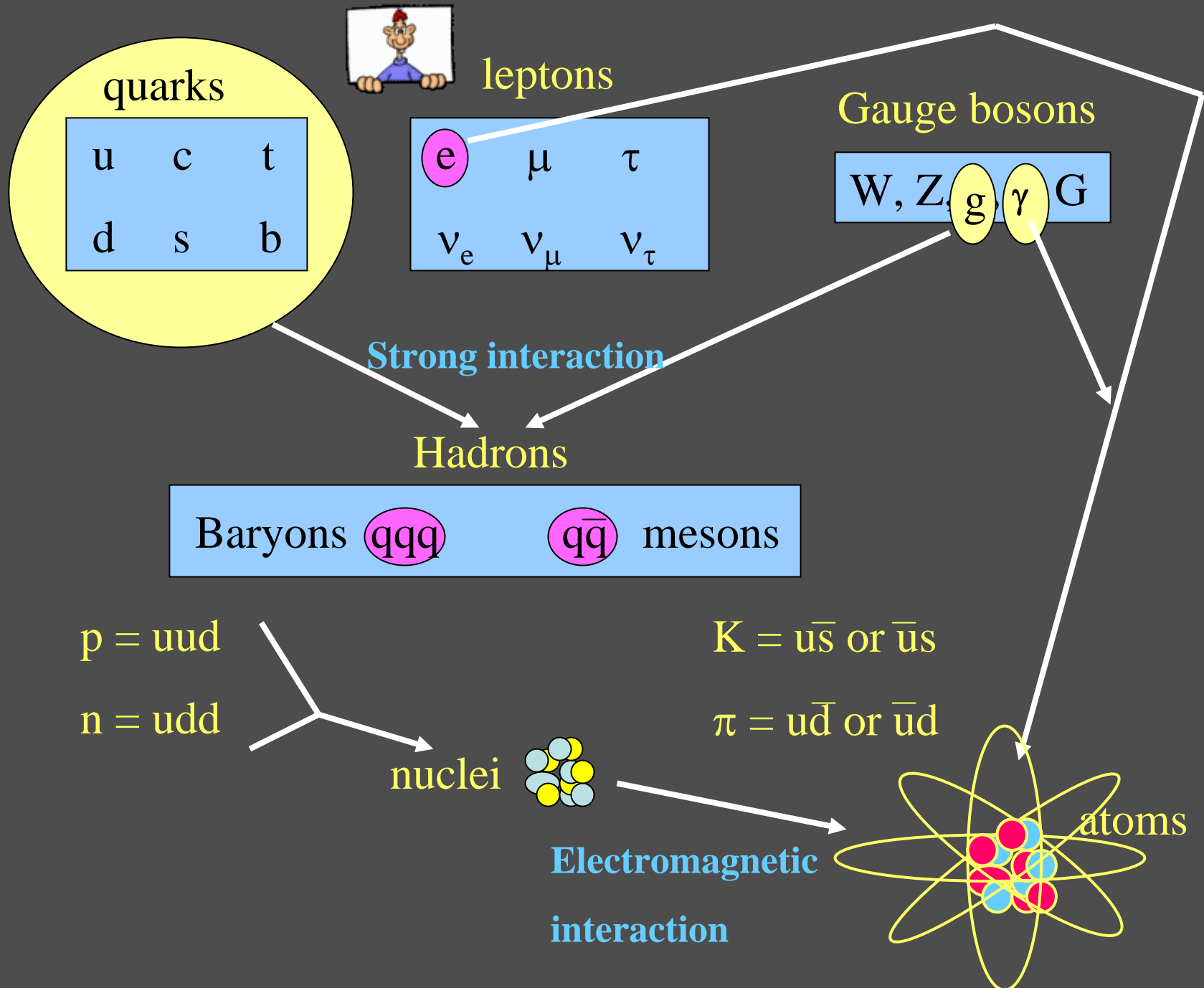
Three Generations of Matter

I II III

Leptons Quarks

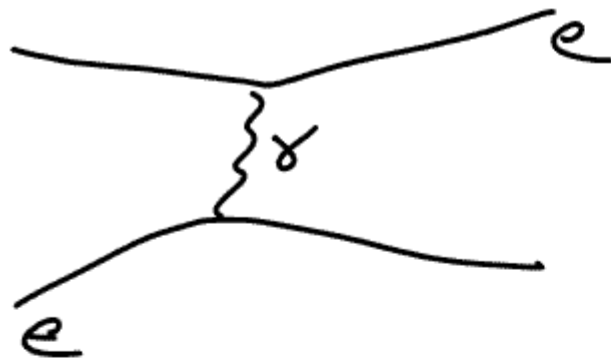
u <small>up</small>	c <small>charm</small>	t <small>top</small>	γ <small>photon</small>
d <small>down</small>	s <small>strange</small>	b <small>bottom</small>	g <small>gluon</small>
ν_e <small>electron neutrino</small>	ν_μ <small>muon neutrino</small>	ν_τ <small>tau neutrino</small>	Z <small>Z boson</small>
e <small>electron</small>	μ <small>muon</small>	τ <small>tau</small>	W <small>W boson</small>

Force Carriers



QUANTUM Field Theory \rightarrow Exchange force





$$\Delta E \Delta t \sim h$$

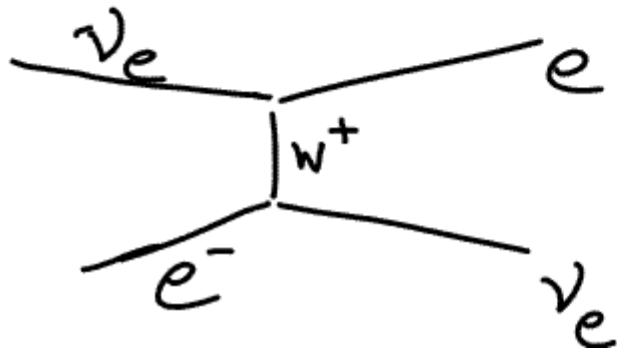
↑ ↖

Small big
Amount

massless $\gamma \rightarrow$ large (infinite range)

Electromagnetism Quantum electrodynamics

Weak nuclear interaction



$$\text{Mass } W = 86 \text{ GeV}$$

$$1 \text{ GeV} = 1000 \text{ MeV}$$

$$\text{Range} \sim 10^{-18} \text{ m}$$