

Physics 100 - October 24, 2007



Hello from Fermilab



<http://www.fnal.gov>

Last time

Heisenberg's Uncertainty Principle

$$\Delta x \Delta p \sim h$$

$$\Delta E \Delta t \sim h$$

Say Goodbye to the deterministic Universe

Say Hello to the Harry Potter universe

$\Delta E \Delta t \sim h$ \rightarrow gives us a loophole in energy conservation that holds the key to understanding the fundamental essence of forces in Quantum Mechanics

We'll come back to this...

OUR Journey to date



crud

layer 1
- Atom

Atomic
Structure

Layer 2



10^{-16} m

10^{-9} m

Nuclear
Physics
layer 3

Journey into Inner Space

- Protons
- neutrons

$N\bar{u}k - l\bar{e} - ar$

$N\bar{u} - k\bar{u} - lar$

Please!!
Pronounce the
Word
Correctly



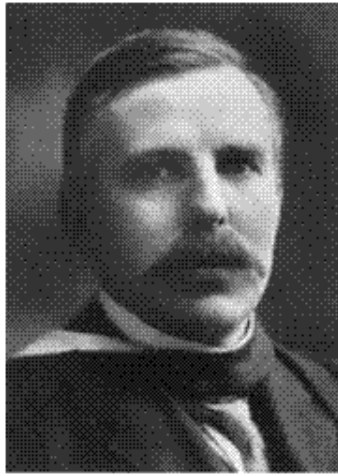
Ze^-



Z protons

A-Z neutrons

How did Bohr know to use "nuclear" model of Atom??



Ernest Rutherford

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(1871-1937)

(New Zealand Farm boy → England)

1908 Nobel Prize in Chemistry

Established Nuclear Model of Atom

Rutherford
(Manchester)



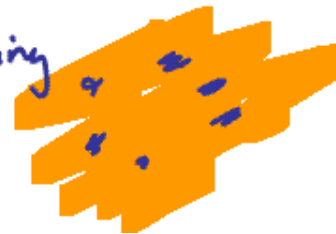
Nuclear

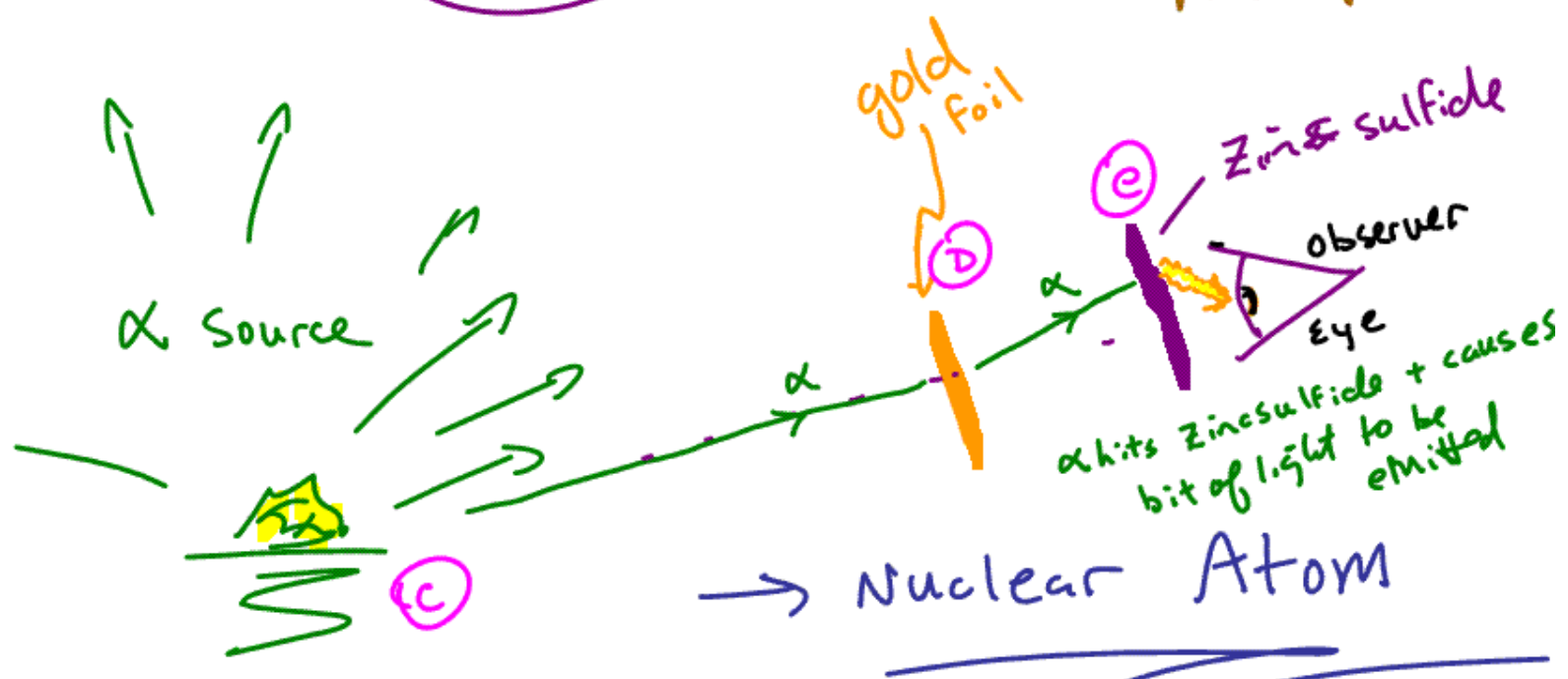
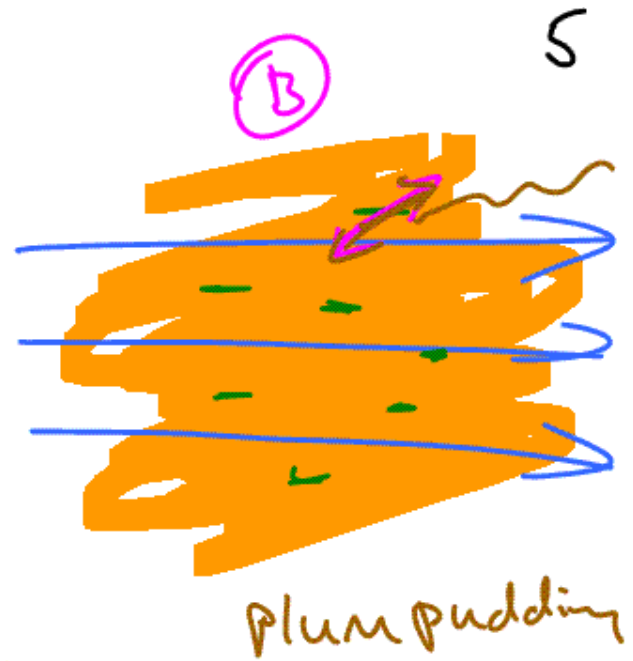
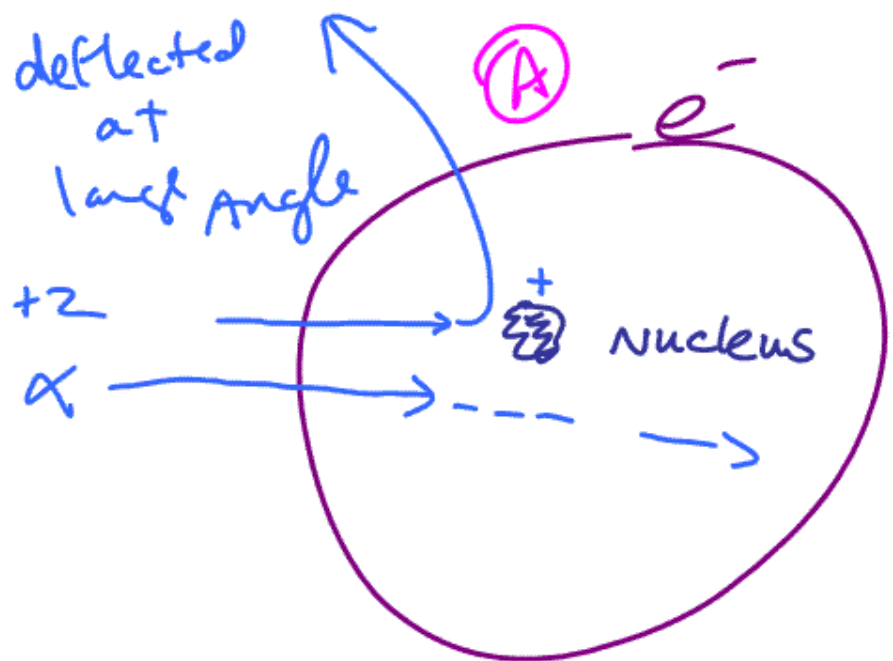
(A)

J.J. Thompson
(Cambridge)

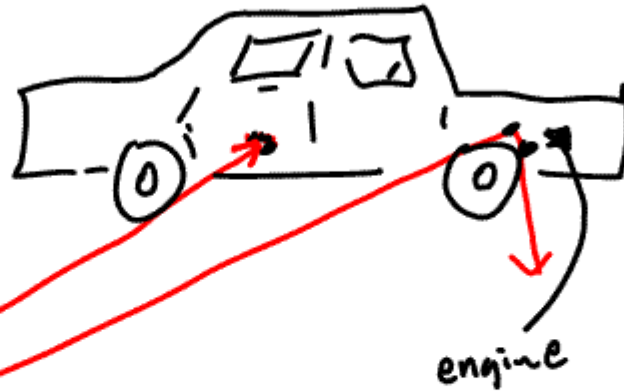
Plum
Pudding

(B)





Can think of Rutherford expt. as mapping out mass (charge)



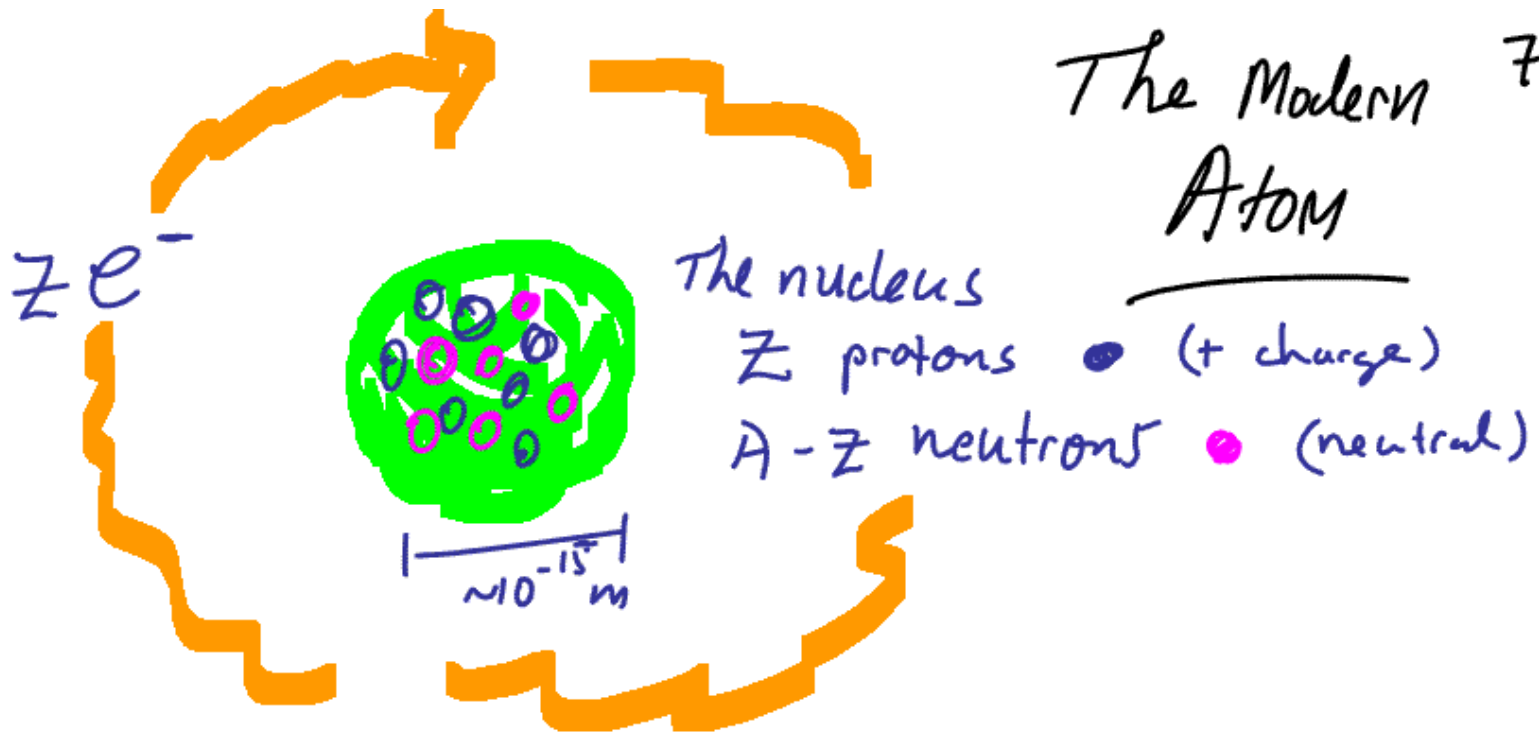
distribution
of atom
by probing
w/ "bullets"
and looking
at ricochets

could do same
with car and rifle

bullets would not go "thru" car when
hitting engine.



The Modern ^Z Atom

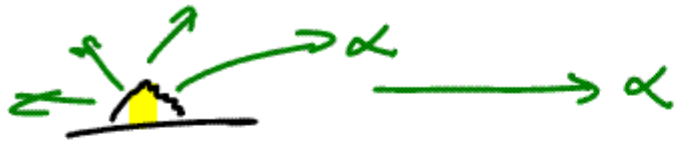


This is wacked! Why?

nucleus should blow apart with
all that + charge crammed
into such a small volume

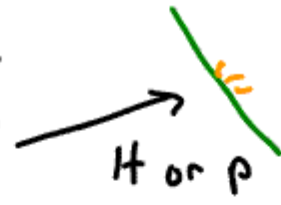
What's in the nucleus? What holds it together?

The Proton



Rutherford 1918

{ Nitrogen }



observer

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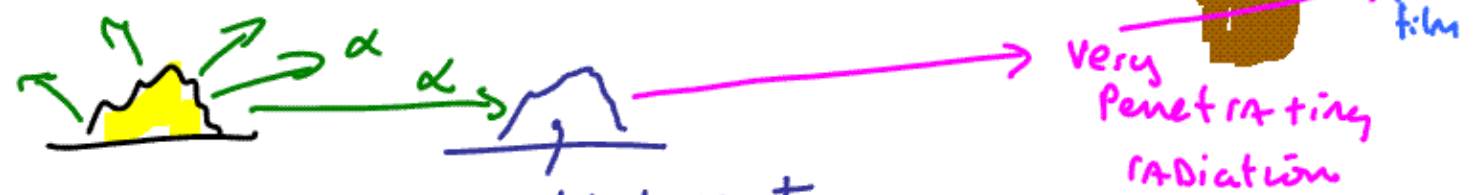


mass $\sim 938 \text{ MeV}/c^2 = 1 \text{ AMU} \sim \text{Atomic Mass Unit}$
charge = +1

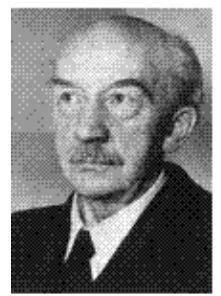
for comparison Mass of electron $\sim 0.511 \text{ MeV}/c^2$

Spin = $\frac{1}{2} \rightarrow \text{Fermion}$

The neutron



1930 Walther Bothe
H. Becker
(Germany)



1954 Nobel Prize in Physics

"For the coincidence method and his discoveries made therewith"

1932 James Chadwick → 1935 Nobel Prize "For discovery of the neutron"
(England)



Showed this penetrating radiation to NOT be γ -rays
Showed had mass similar to proton and was uncharged

Mass of neutron $\sim 940 \text{ MeV}/c^2$ electric chg = 0
Spin = $1/2$ fermion

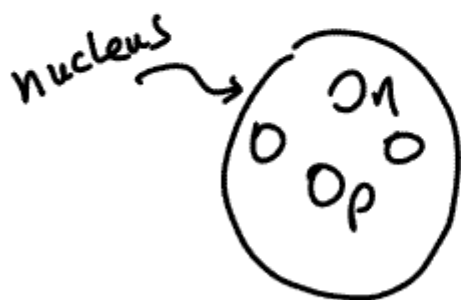
New Attraction force

Strong Nuclear Force

Two Protons Approach each other

$P \rightarrow$ $\leftarrow P$

Coulomb repulsion until 10^{-15} m distance
then New Attraction force kicks in
and they stick together



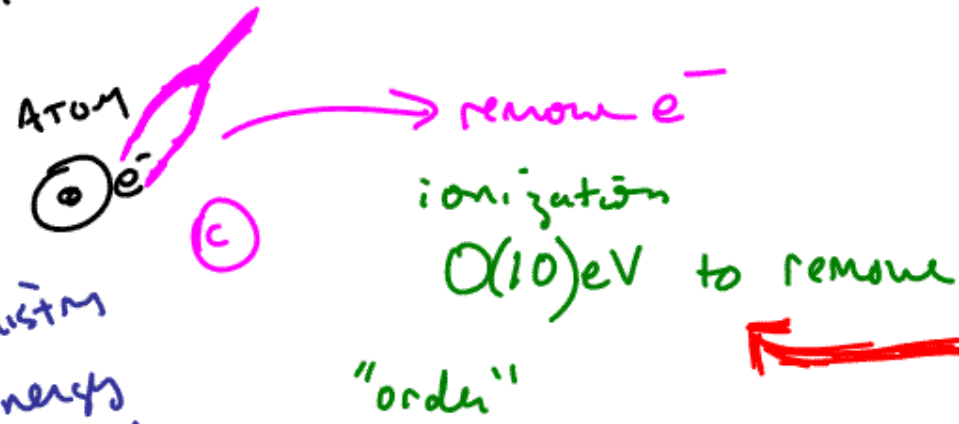
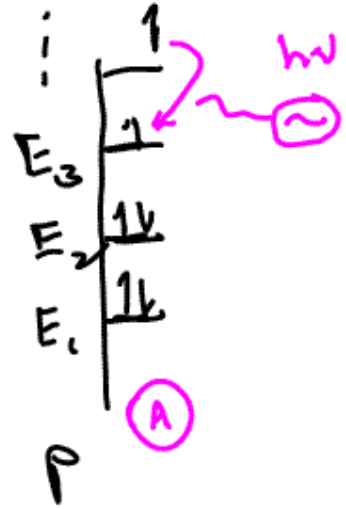
Solve Schrödinger's eqn

→ discrete Available energy levels
(just like the atom!)

$P, n \equiv$ collectively called "nucleons"

Populate "Nuclear" orbitals w/ n, p

just like e^- in Multi- e^- Atom
...
F. ll shells w/ P, n



Chemistry Energy Scale

Note the Huge difference ...
Nuclear reactions require/release energy 1 million times greater than atomic rxns

$O(\text{MeV})$ means "order of MeV" means
Approximately an MeV
 $\sim 1 - 10 \text{ MeV}$

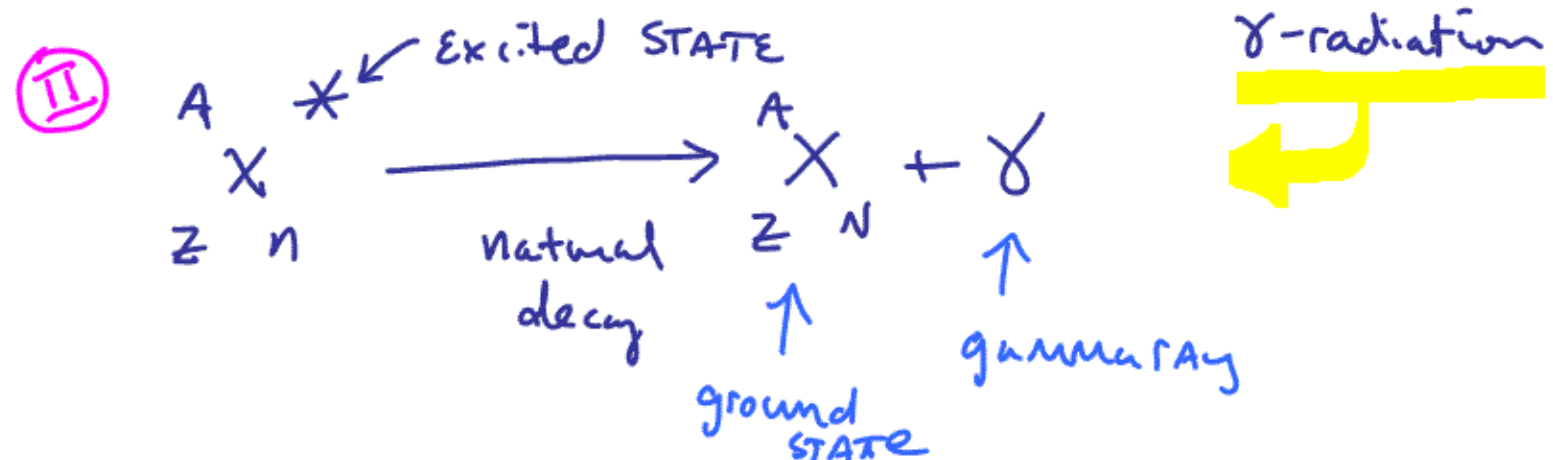


X ≡ Atomic Element Symbol
determined by Z

Z ≡ # protons

A ≡ ATOMIC MASS = # p + # n

N = # neutrons



Naturally radioactive substance - nucleus unstable

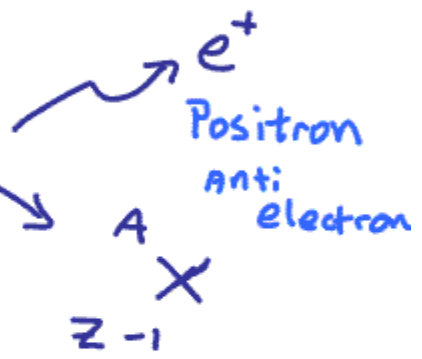
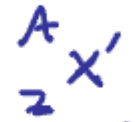
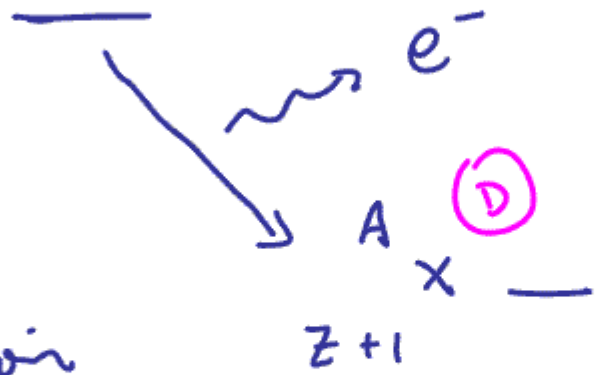
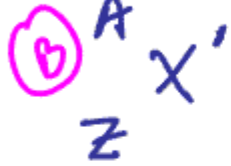
Will decay naturally

(A) $E = mc^2$

(C) $n \rightarrow p + e^-$

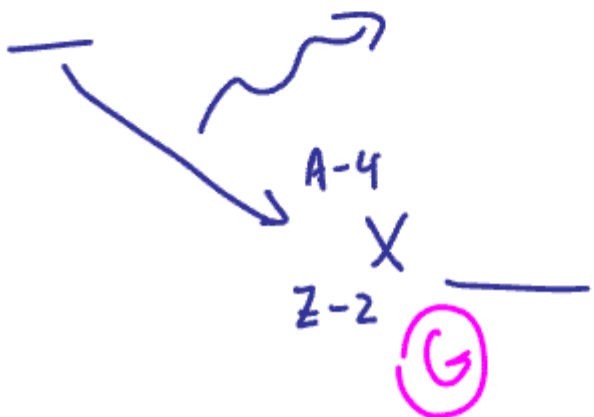
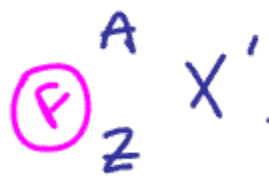
(E) $p \rightarrow n + e^+$

Beta decay



β -radiation

(I) $\alpha \equiv \begin{matrix} p & n \\ n & p \end{matrix}$ alpha particle
 \downarrow
 Helium Nucleus
alpha Radiation



Alpha decay

Different nuclei → different numbers of p and n

↳ different nuclear shell structures

Stable nuclei

naturally radioactive → characteristic decay times

decay time characterized

(A)

units of time

$t_{1/2}$ = half-life ≡ time it takes for 1/2 atoms in a given sample to decay



(B)

activity = $\frac{\Delta N}{\Delta t}$ ≡ # decays in time Δt = λN

λ ≡ units of $\frac{1}{s}$ decay constant # atoms in sample