

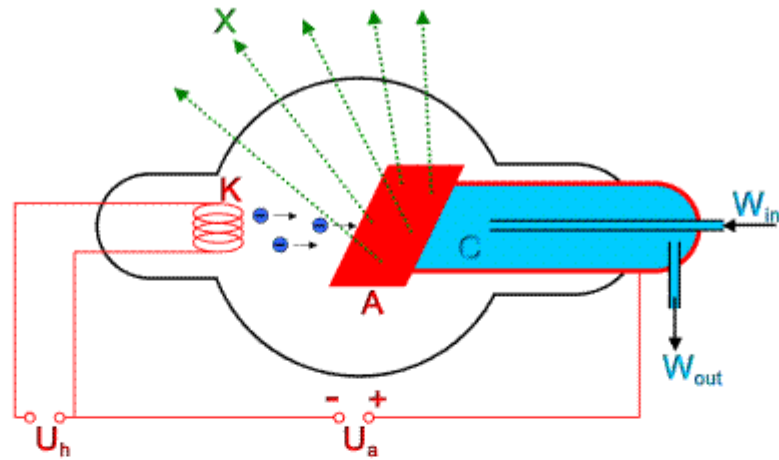
Physics 100 - March 7, 2007

█ X-ray and laser reading

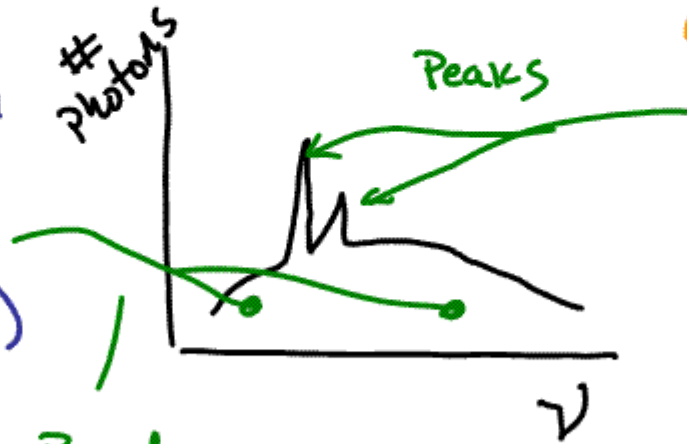
█ ~~Thermodynamics~~ for now

█ TA evaluations ... please do

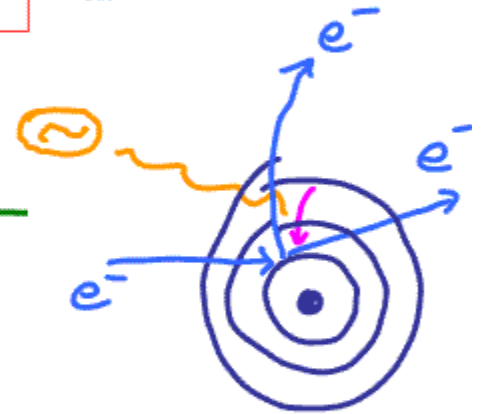
X-RAYS



electron beam
Slows down
→ radiates
(Bremsstrahlung)

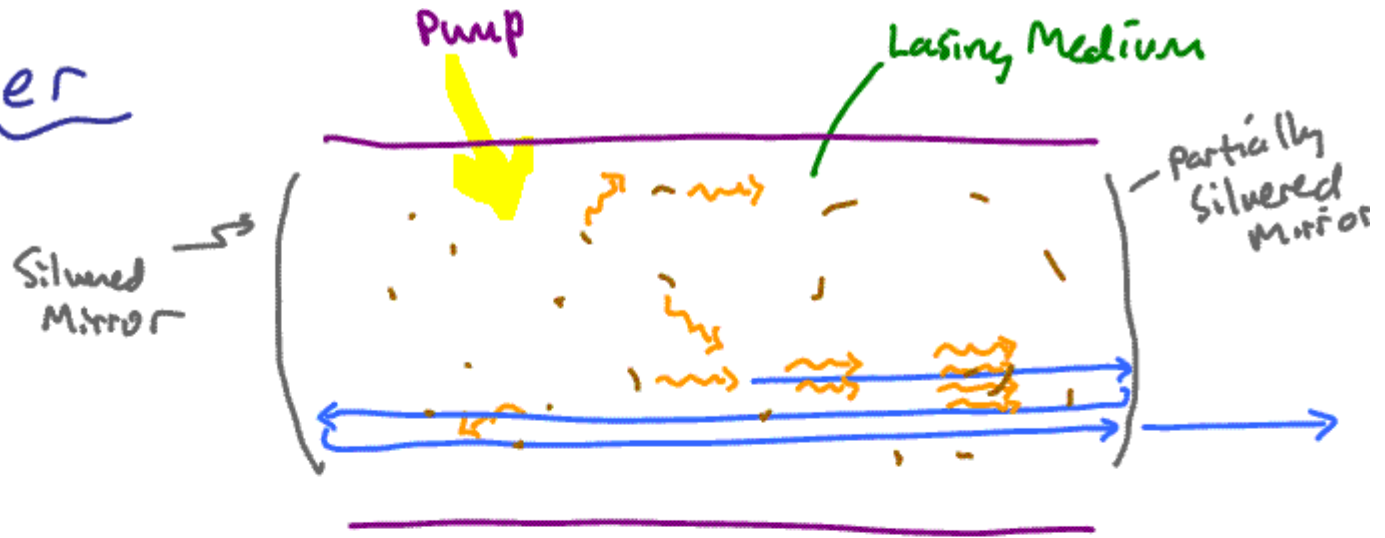



Broad
swath
of
frequencies



Knock out
inner shell e^-
in large- Z
nucleus
outer shell e^-
cascades
down

Laser



- ① Atoms in ground STATE
- ② Pump Atoms to higher energy STATE \rightarrow Population inversion
(for example - illuminate with Bright light)
- ③ Spontaneous emission of light happens
- ④ Stimulated emission causes "gain" in light
radiation is coherent 
- ⑤ Bounces back + forth Many times
Stimulated emission increases gain
- ⑥ Very intense coherent light penetrates partially
Silvered mirror

OUR Journey to date



layer 1
- Atom



Atomic Structure

10^{-9} m

Layer 2



10^{-15} m

Nuclear Physics
layer 3



Z protons
A-Z neutrons

$N\bar{u}k - l\bar{e} - \partial r$

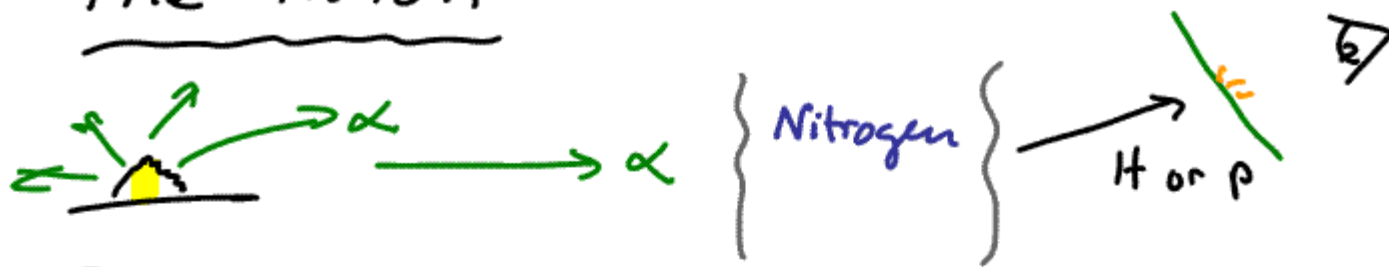
$N\bar{u} - k\bar{u} - l\partial r$

Please!!
Puhlease!!
Pronounce the
Word
Correctly

- Protons
- neutrons

with all these protons (positive charges) in such a small region... why does the nucleus NOT Blow apart??

The Proton



Rutherford 1918

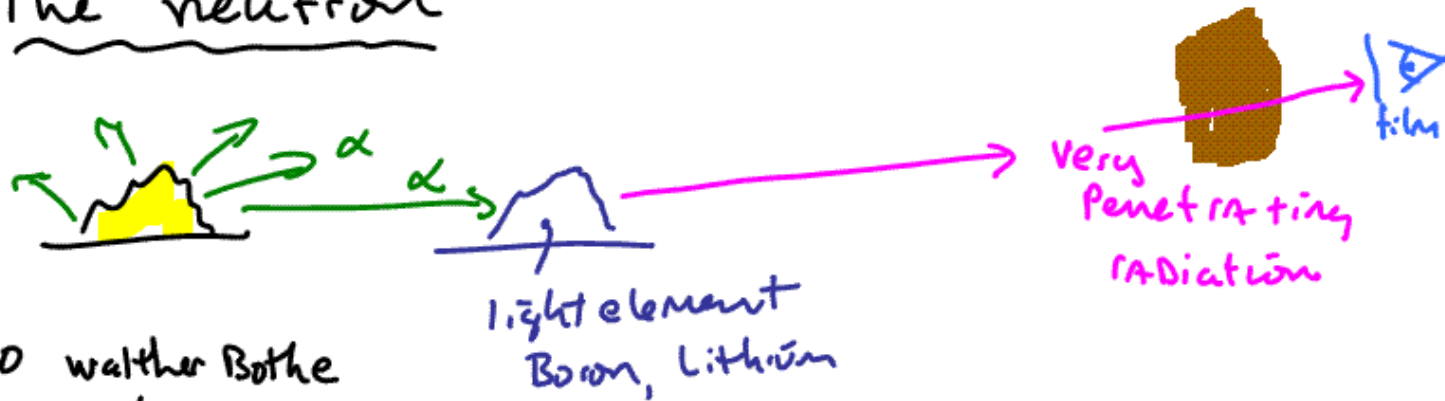


Mass $\sim 938 \text{ MeV}/c^2 = 1 \text{ AMU} \sim \text{Atomic Mass Unit}$
Charge = +1

for comparison Mass of electron $\sim 0.511 \text{ MeV}/c^2$

Spin = $\frac{1}{2} \rightarrow \text{Fermion}$

The neutron



1930 Walther Bothe

↙ H. Becker
(Germany)



1954 Nobel Prize in Physics

"For the coincidence method and his discoveries made therewith"

1932 James Chadwick
(England)



Showed this penetrating radiation to NOT be γ -rays

Showed had mass similar to proton and was uncharged

→ 1935 Nobel Prize "For discovery of the neutron"

Mass of neutron $\sim 940 \text{ MeV}/c^2$

electric chg = 0
Spin = $1/2$ fermion

New Attraction force

Strong Nuclear Force

Two Protons Approach each other

$P \rightarrow$

$\leftarrow P$

Coulomb repulsion until 10^{-15} m distance
then New Attraction force kicks in
and they stick together

nucleus



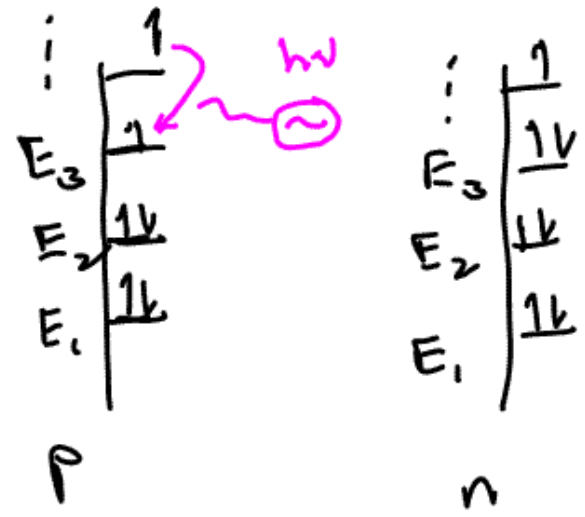
Solve Schrödinger's eqn

→ discrete Available energy level
(just like the atom!)

$P, n \equiv$ collectively called "nucleons"

Populate "Nuclear" orbitals w/ n, p

just like
e⁻ in
Multi-e⁻
Atom
:
Fill
shells
w/ P, n



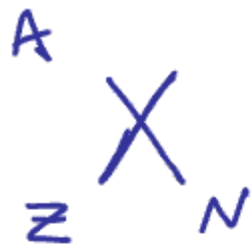
Chemistry
Energy
Scale

remove e⁻
ionization
O(10)eV to remove
"order"

TAKES
O(MeV)
Energy
to
Remove
from
nucleus

O(MeV) means "order of MeV" means
Approximately an MeV
~ 1 - 10 MeV

Note the Huge
difference ...
Nuclear reactions
require/release
energy 1 million
times greater
than
atomic
rxns



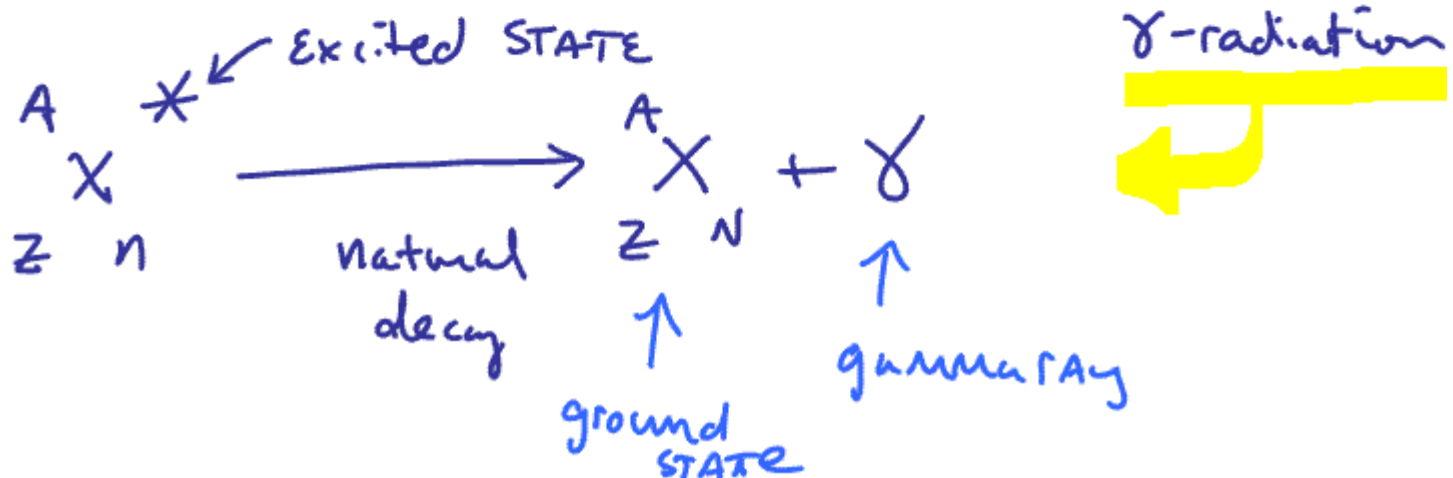
nuclei symbolized as

$X \equiv$ Atomic Element Symbol
determined by Z

$Z \equiv$ # protons

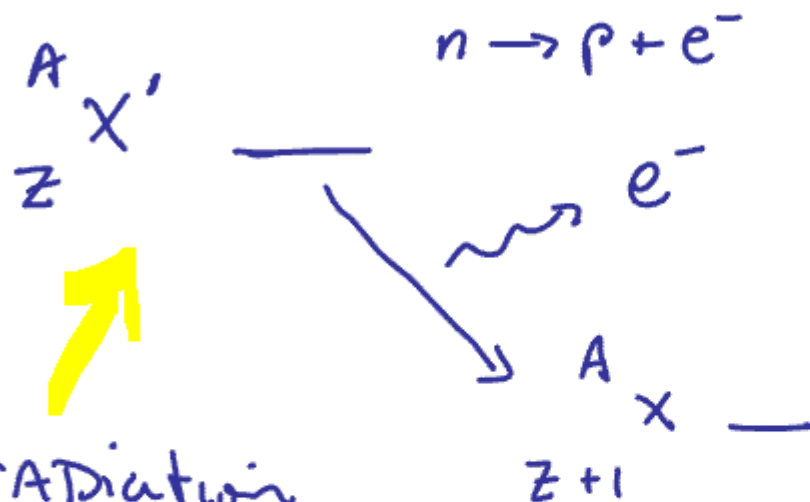
$A \equiv$ ATOMIC MASS = # p + # n

$N =$ # neutrons

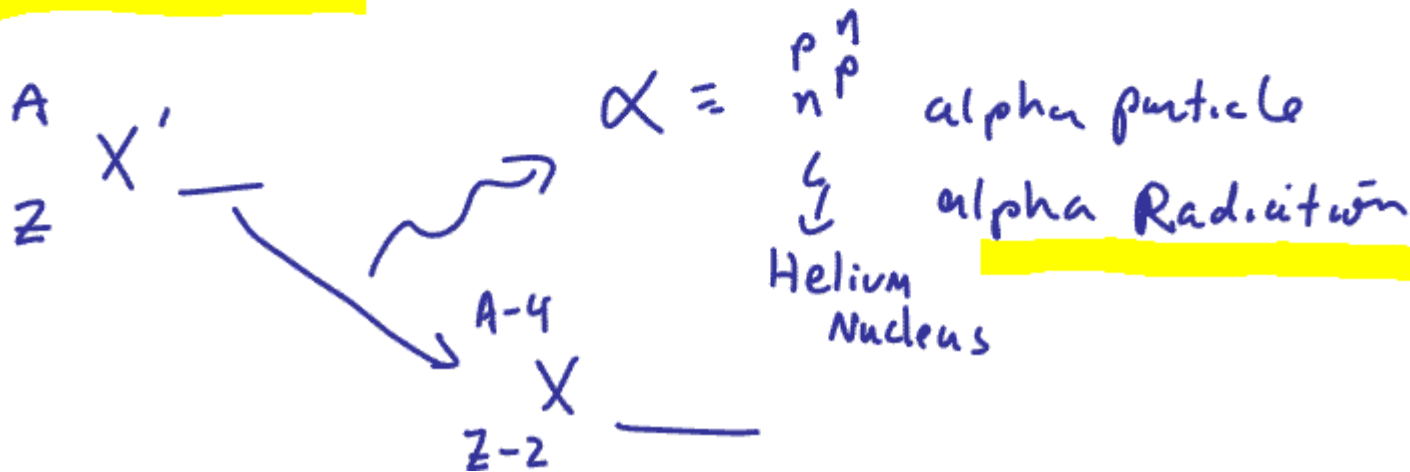
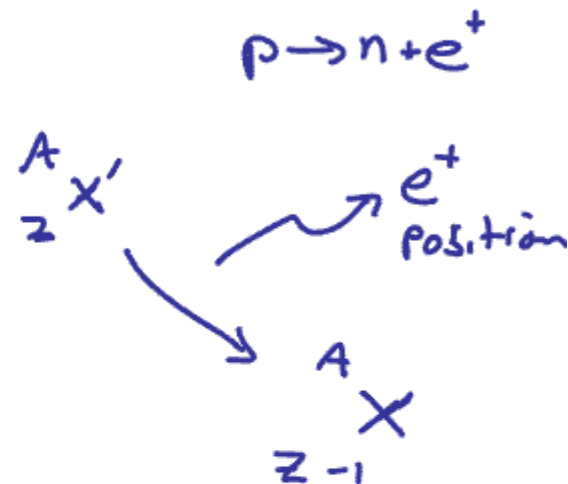


Naturally radioactive substance

$$E = mc^2$$



β -radiation

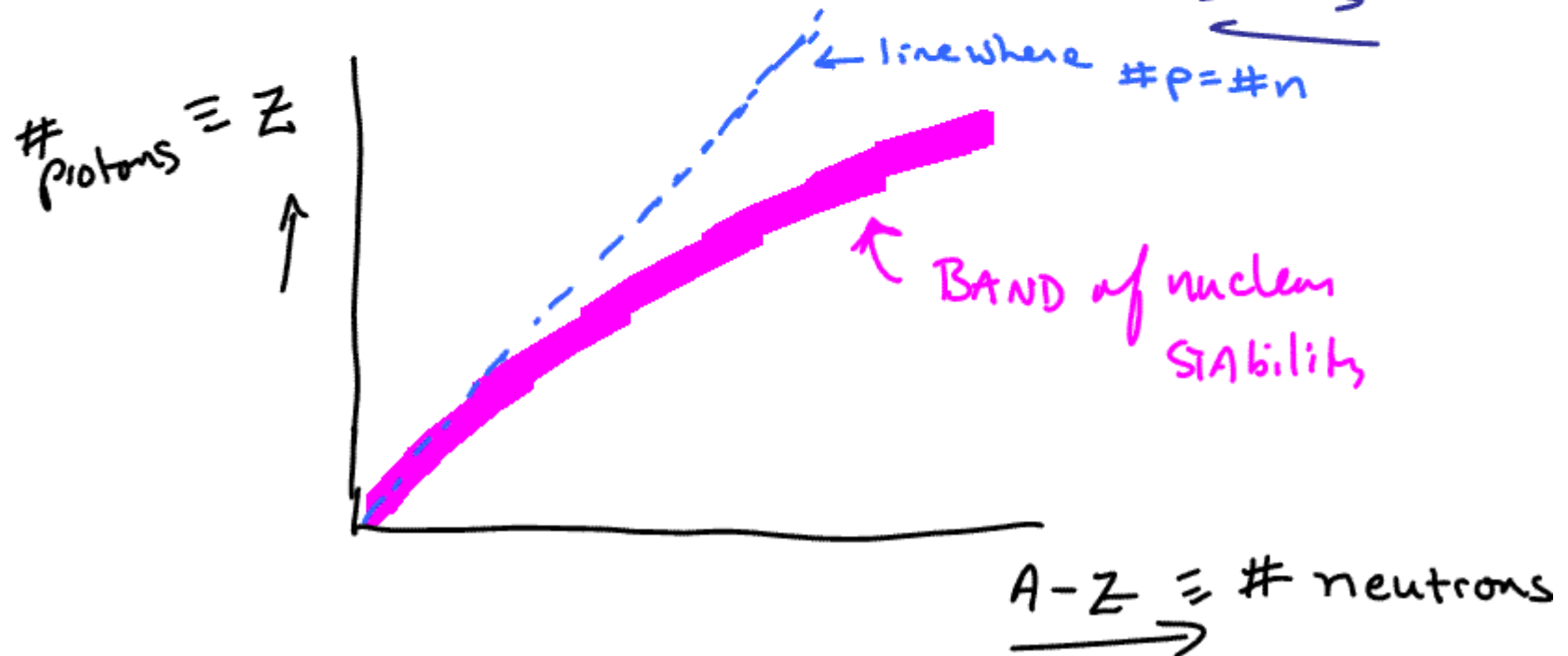


$Z \equiv \# \text{ protons} \rightarrow \text{defines chemistry}$



$\# \text{ neutrons can vary}$

same Z , diff $\# \text{ neutrons} \rightarrow \text{isotopes}$



Different nuclei \rightarrow different numbers of p and n
 \hookrightarrow different nuclear shell structures

Stable nuclei
naturally radioactive \rightarrow characteristic decay times

decay time characterized

$t_{1/2}$ = half-life \equiv time it takes for $\frac{1}{2}$ atoms in a given sample to decay

units of time

$$\text{activity} = \frac{\Delta N \equiv \# \text{decays}}{\Delta t} = \lambda N$$

$\lambda \equiv$ units of $\frac{1}{s}$ λ decay constant N # atoms in sample

$$t_{1/2} = \frac{.693}{\lambda}$$

Radioactive Dating Techniques

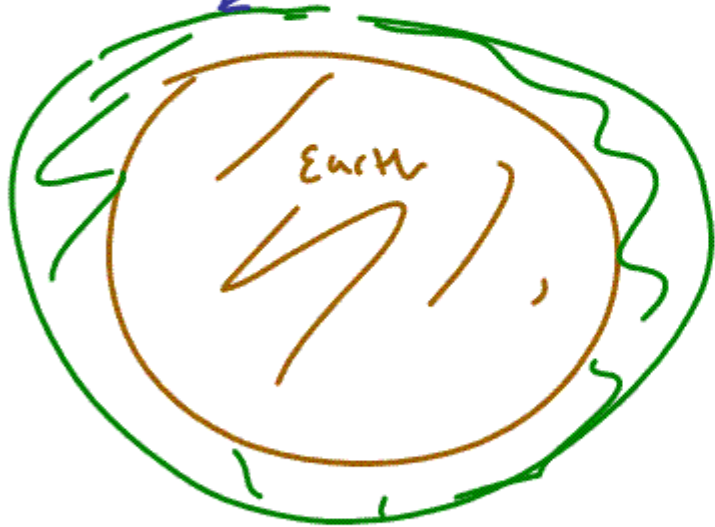
$$t_{1/2} \text{ of } ^{14}\text{C} = 5730 \text{ years}$$

^{14}C naturally decays
is incorporated into
living bodies. When
something dies the incorporation
of new ^{14}C stops.

^{14}C in dead material slowly
decays

Measure Amount of ^{14}C in sample provides a
measure of time since death.

Particles from
Sun and other
astronomical sources
 ^{14}C
 ^{14}C
 ^{14}C in Atmosphere
creates



Binding Energy
nucleon

Important

