

# Physics 100 - March 19, 2007

TA feedback/survey



Welcome  
Back



Presentations approacheth -

Time to set up meetings with prof. Manly

Plans, mode(s) of presentation,

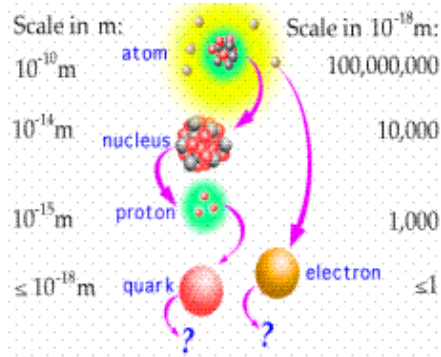
Timing, A/V, DEMOS, questions, sources ...

April 11, 16, 18, 23, 25

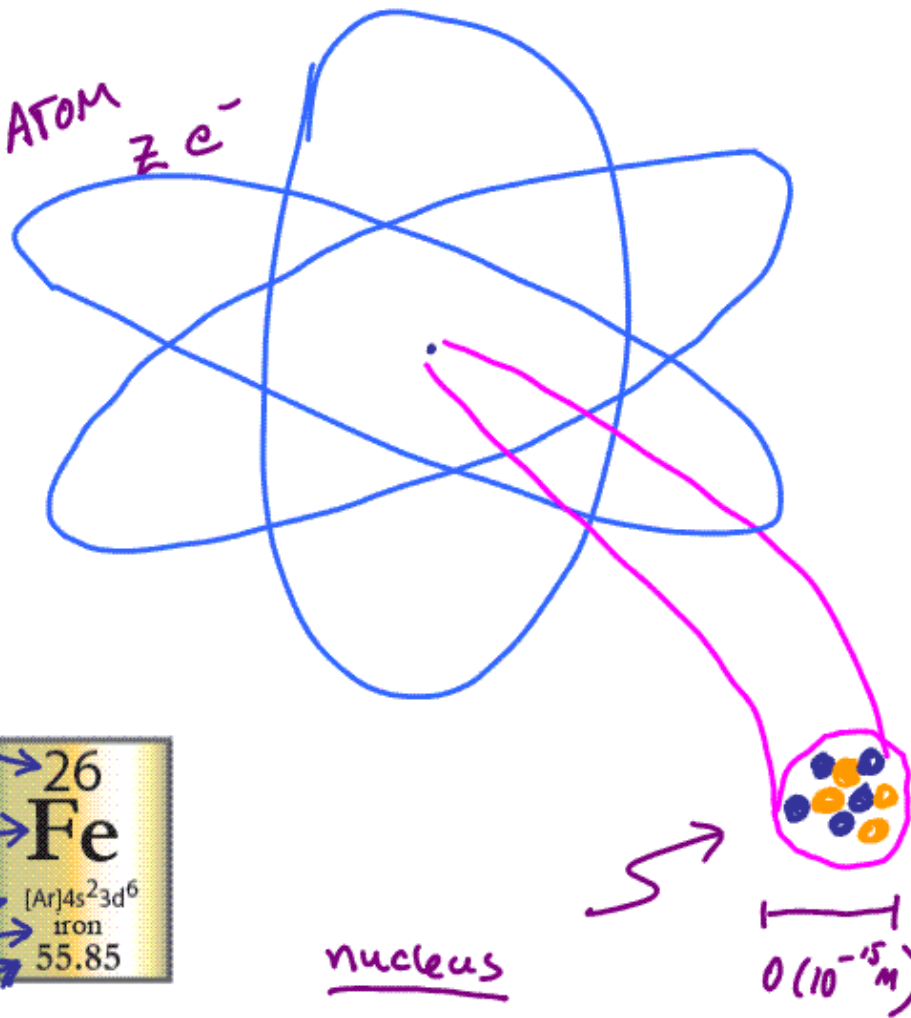
Two per day 1<sup>st</sup> come - 1<sup>st</sup> served Scheduling

Last Time

# nuclear physics



ATOM  $Z e^-$



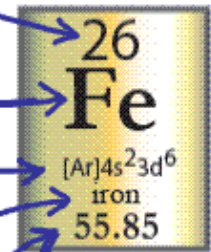
# protons  $\equiv Z$

Atomic Element Symbol

Short-hand for Electron orbital Configuration

Element name

Average Atomic Mass per Atom



nucleons [

nucleus

- Z protons
- A-Z neutrons
- A  $\equiv$  Atomic Mass

0 ( $10^{-15}$  m)

Why does nucleus NOT Blow Apart?

There exists a Strong nuclear force that is  
 Much Stronger than gravitation and electromagnetism  
 but has a range limited to  $\sim 10^{-15}$  m

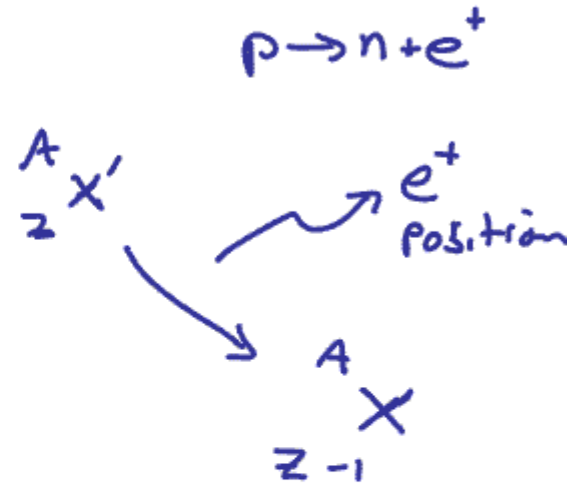
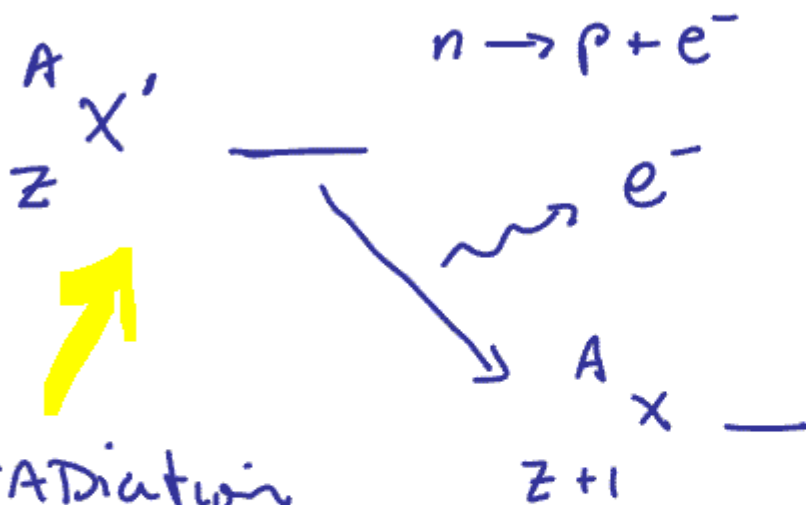
$\Rightarrow$  Pop this force into Schrödinger's equation  
 get discrete energy levels for nucleons

$A$   
 $Z$   $X$   $N$   
 $A \equiv \# \text{ protons} + \text{neutrons}$   
 $Z = \# \text{ protons}$   
 $N = \# \text{ neutrons}$   
 $X = \text{Atomic Element Symbol}$

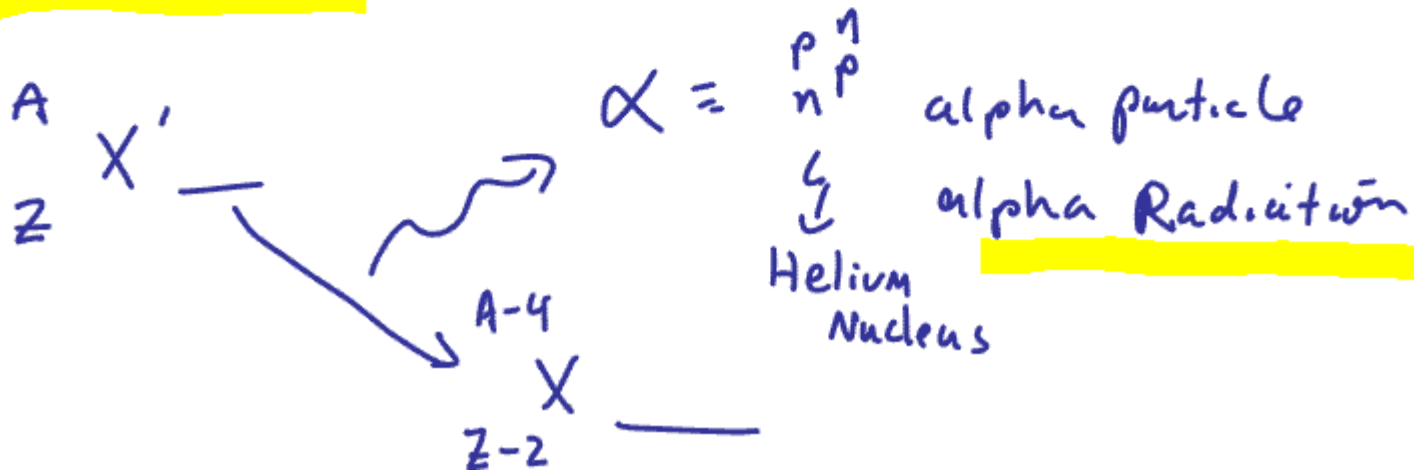
$A$   $X$   $*$   $\leftarrow$  Excited STATE  
 $Z$   $N$   
 $\xrightarrow{\text{natural decay}}$   $A$   $X$   $+ \gamma$   
 $Z$   $N$   $\uparrow$  gamma ray  
 $\uparrow$  ground STATE  
 $\gamma$ -radiation

# Naturally radioactive substance

$$E = mc^2$$



$\beta$ -RADIATION





$N$  atoms in sample  
at time = 0

$$\text{Activity} \equiv \frac{\# \text{decays}}{\text{second}} = \frac{\Delta N}{\Delta t} = \lambda N$$

↑  
decay CONSTANT

half life  $\equiv t_{1/2}$  = time for  $\frac{1}{2}$  sample to  
decay

$$t_{1/2} = \frac{0.693}{\lambda}$$

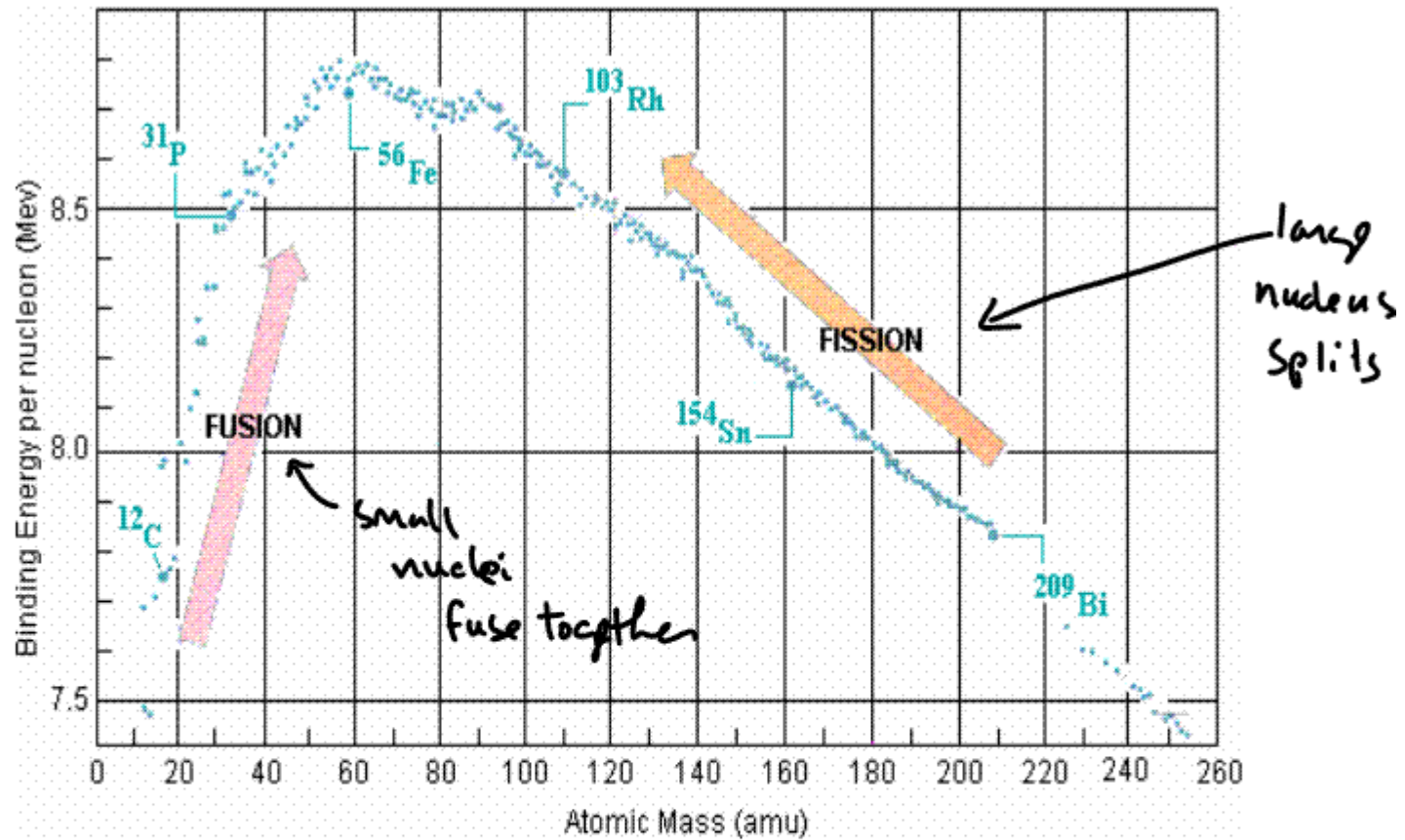
A sample of bone from an archeological dig has a  $^{14}\text{C}$  activity that is 25% of that in a living material.  
How old is the bone?

$$^{14}\text{C} \quad t_{1/2} = 5730 \text{ years}$$

1 half life : Activity falls to 50% of original

2 half lives : " " " 25% " "

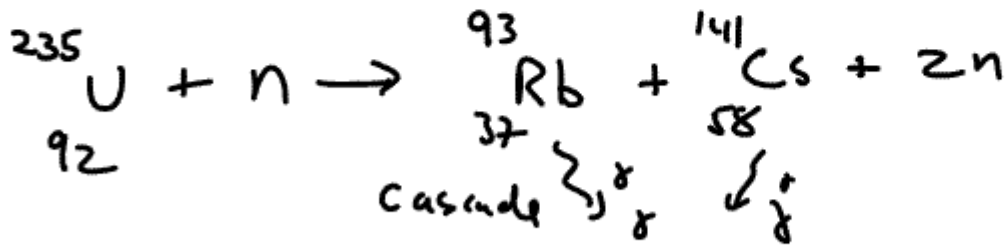
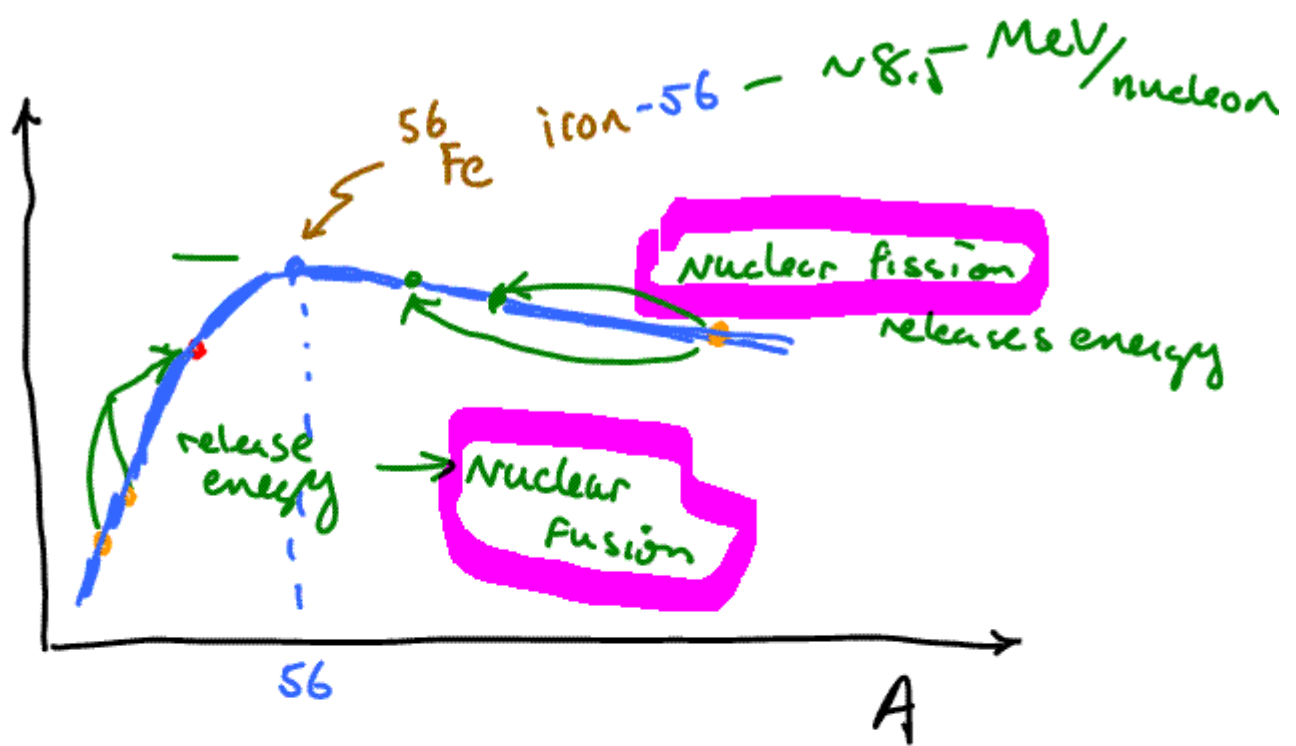
Age  $\sim (2) 5730 \text{ years} \sim 11,400 \text{ years}$



Binding energy/nucleon is a measure of how hard it is to remove a nucleon from the nucleus ... similar to how ionization energy is a measure of how hard it is to remove  $e^-$  from atom

Binding Energy  
nucleon

Important



M

M'

$$E = Mc^2$$

M > M'  
converting MASS  
to energy

Critical 1 split for each split

Subcritical < 1 split for each split

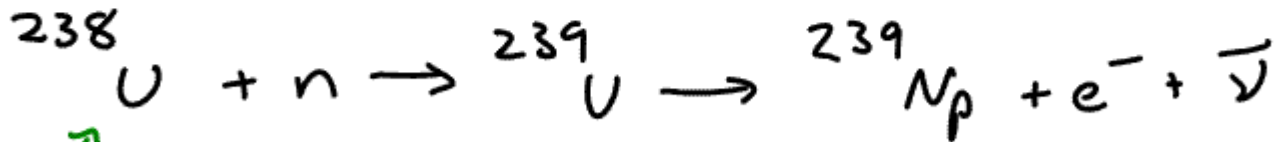
Super critical > 1 split " " "

$^{235}\text{U} \sim 0.7\%$  naturally occurring U

$^{238}\text{U} \sim 99.3\%$  " " "

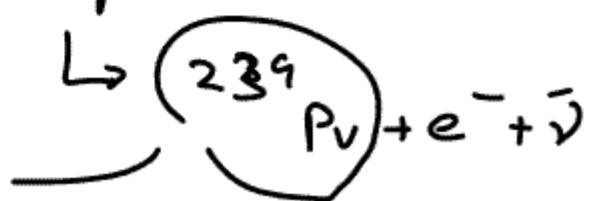
MUST "enrich"  
sample in  $^{235}\text{U}$   
to make useful  
for bombs or  
reactors  
- Hard to do

### Breeder reactor

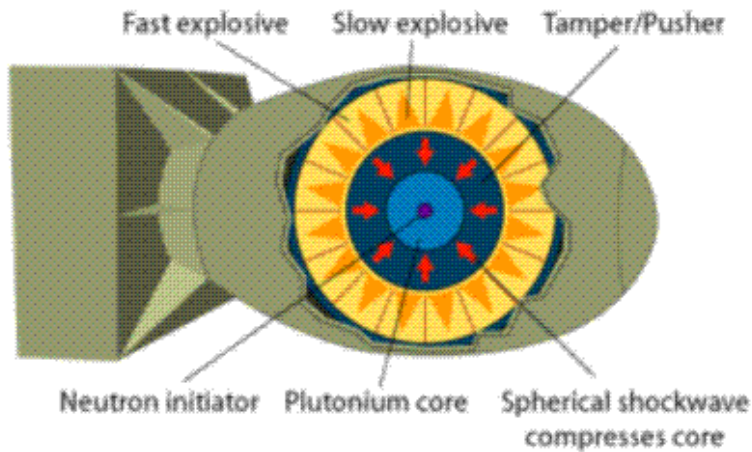


a not so useful  
isotope of Uranium

can use as  
nuclear  
fuel

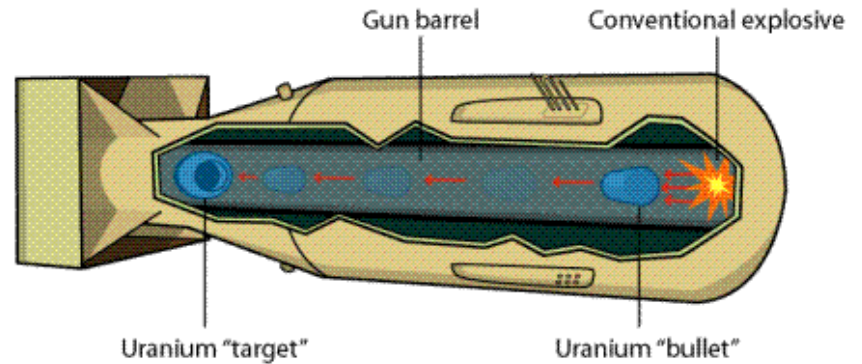






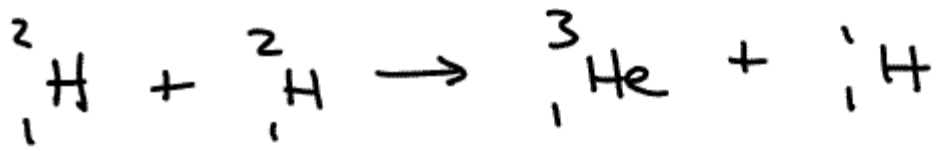
Similar to Fat Man  
used on Nagasaki  
Aug 9, 1945

Similar to "Little Boy"  
used on Hiroshima  
August 6, 1945



Diagrams  
from  
Wikipedia





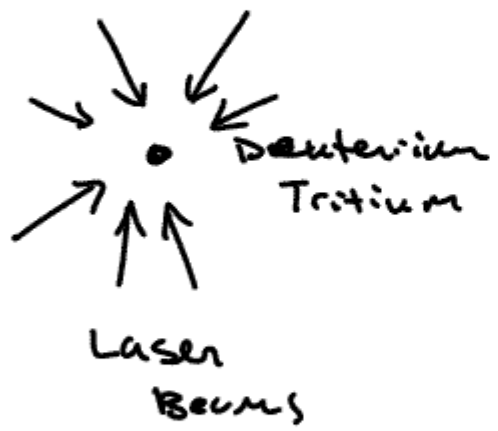
nuclear Fusion

small nuclei join to become larger nuclei

1. tritons  $\sim 1$  MeV energy

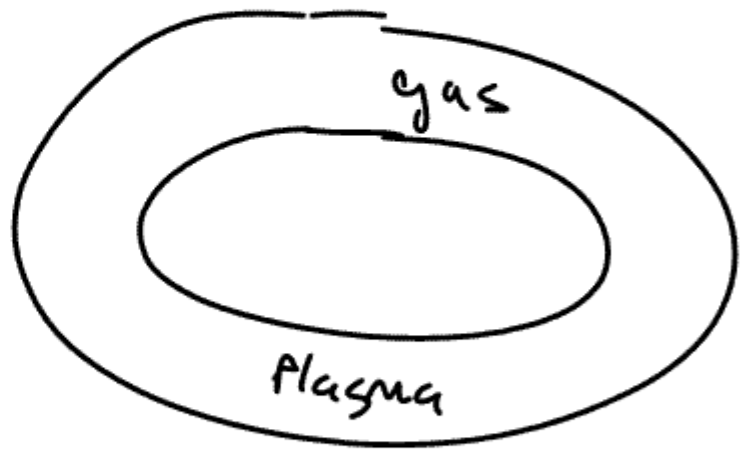
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NOVA



Inertial confinement fusion

Magnetic confinement



TOMAKAK

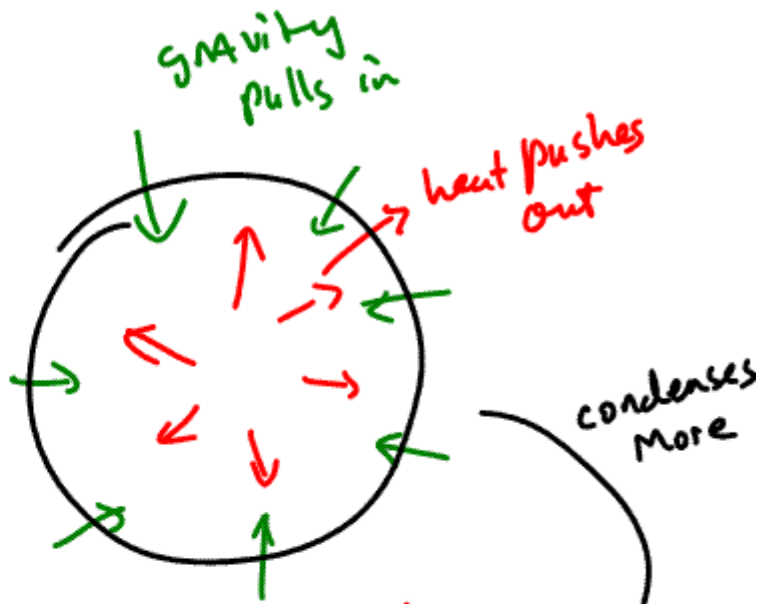
ITER

Takes a huge amount of energy or heat to force one (positive) nucleus close enough to another (positive) nucleus to fuse.

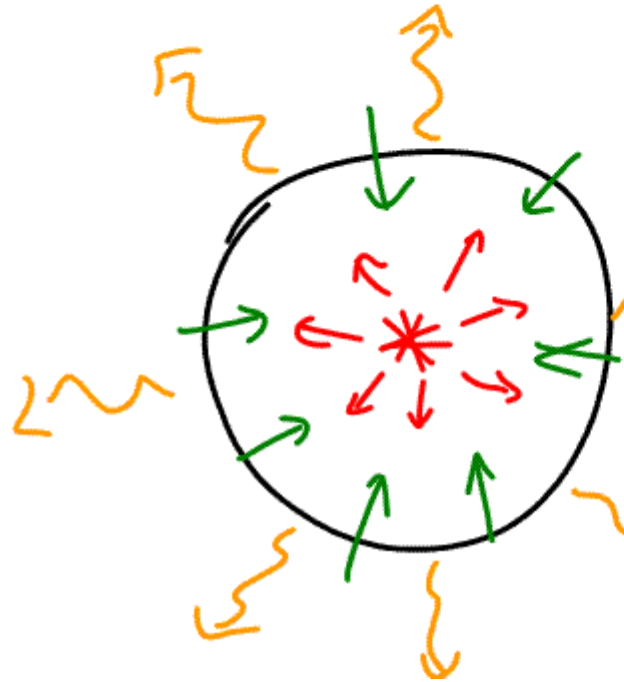
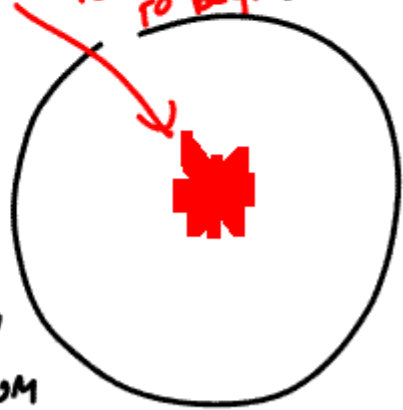
# STAR



condenses  
due to  
gravity



Becomes hot enough  
for fusion  
reactions  
to begin



Reach equilibrium  
between heat from  
fusion process in core  
pushing out and gravity  
pulling in.